# GUIDE TO LUNAR ORBITER PHOTOGRAPHS

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## **Preface**

This document provides information on the location and coverage of each photograph returned by the Lunar Orbiter series of spacecraft. Small-scale maps show the overall coverage of each mission and the areas of common coverage among sites of different missions. Large-scale maps show coverage of the individual photographs at each target area. The characteristics of the cameras and of the various orbital sequences utilized are given for background information pertinent to an understanding of Lunar Orbiter photography.

#### Introduction

The Lunar Orbiter program initiated in early 1964 consisted of the investigation of the Moon by five identical unmanned spacecraft. Its primary objective was to obtain detailed photographs of the Moon. This document presents information on the location and coverage of all Lunar Orbiter photographs and is one in a series of four NASA Special Publications documenting Lunar Orbiter photography. The others are references 1 to 3. Reference 1 contains 675 photographic plates and provides coverage of the complete Moon with more detail than any other publication. Reference 2 is a collection of approximately 180 selected photographs and portions thereof at enlarged scale, and includes captions for each photograph. Reference 3 shows each named feature on the near side on annotated highresolution frames from mission IV. It also includes (1) an alphabetical index of features, (2) cross-indexes between listings in the catalog of the University of Arizona and the catalog of the International Astronomical Union which was published in 1935, and (3) listings of named lunar features on the near side covered during missions I, II, III, and V, and their photograph numbers.

The objectives of the Lunar Orbiter program were-

- (1) Photography.—To obtain detailed lunar topographic and geologic information of various lunar-terrain types to assess their suitability for use as landing sites by Apollo and Surveyor spacecraft and to increase man's scientific understanding of the Moon.
- (2) Selenodesy.—To provide precision trajectory information which would improve the definition of the lunar gravitational field.
- (3) Moon environment.—To provide measurements of micrometeoroid and radiation flux in the lunar environment for spacecraft performance analysis.

These objectives were accomplished by the flights of five spacecraft during the 13-month period from August 1966 to September 1967. In addition to references 1 to 3 on Lunar Orbiter photography, the interested reader is directed to references 4 to 7 for results of the program.

The five Lunar Orbiter spacecraft returned over 1654 highquality photographs taken from lunar orbit. Each spacecraft was similarly equipped with two cameras which operated simultaneously and had the same line of sight but different fields of view and resolutions. The cameras utilized a common supply of 70-mm film and the dual images they recorded are referred to as medium-resolution frames and high-resolution frames.

Of the 1654 Lunar Orbiter photographs, 840 are of areas photographed on the basis of Apollo program requirements and were obtained primarily during missions I, II, and III. They were taken from low flight altitudes and provided detailed coverage of 22 areas located along the equatorial region of the near side of the Moon. The remaining 814 photographs were taken primarily during missions IV and V and include 703 of the near side of the Moon, 105 of the far side of the Moon, and 6 of the Earth. These photographs were taken from flight altitudes ranging from approximately 44 km over the near side to approximately 6000 km over the far side, and provide broad coverage of essentially the entire Moon and detailed coverage at 88 areas on the near side.

This document contains tables and maps which catalog the various types of Lunar Orbiter photography conducted and aid the user in procuring photographs of selected areas. The maps were prepared by the U.S. Air Force Aeronautical Chart and Information Center, in support of preliminary photo analyses performed immediately following each Lunar Orbiter mission.

The National Space Science Data Center (NSSDC) at Goddard Space Flight Center, Greenbelt, Md., is responsible for dissemination of Lunar Orbiter photographs and other scientific data. Scientists requiring high-quality Lunar Orbiter photographs for study can obtain them from that Center. Persons interested in Lunar Orbiter photographs for other reasons should direct their requests to NASA, Public Information Division, Code FP, Washington, D.C. 20546.

## Lunar Orbiter Photographic System

A sketch of the photographic system of the spacecraft is shown in figure 1. The system was housed in a pressurized, thermally controlled container, and included the cameras, film and film handling, film processor, and readout equipment and environmental controls. The system was designed to expose, develop, and read out images for transmission to Earth by the communications system.

The two cameras simultaneously placed two discrete frame exposures on a common supply of 70-mm aerial film. Each camera operated at a fixed aperture of f/5.6 with controllable shutter speeds of 0.01, 0.02, or 0.04 second. One of the lenses had a 610-mm focal length; the other, an 80-mm focal length.

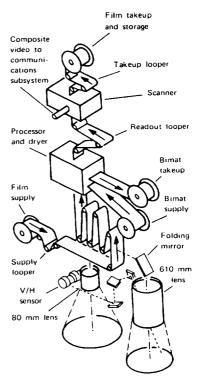


FIGURE 1.—Photographic subsystem.

Shutter, platen, and image-motion compensation were provided for each camera; the film, film advance, and shutter operation were common to both. The film was developed onboard by using a method which passed the film into contact with a web that contained a single-solution processing chemical. After the film was dried, it was stored ready to be read out and transmitted to Earth.

Figure 2 shows a schematic of the readout system which used a line-scan tube as the light source for scanning the negative image on the spacecraft film. The line-scan tube electronically scanned the beam of light a distance of 2.667 mm in the lengthwise direction of the film. The sweep of this line across the film was accomplished by a mechanical drive of the scanner lens which focused the line. One traverse of the

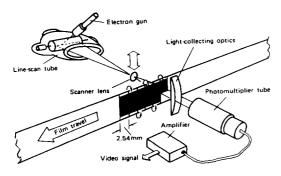


FIGURE 2.—Readout scanner.

scanner lens across the film required approximately 22 seconds; during this time the line scan was repeated over 17 000 times. The sections of film that were read out with this type of scan were referred to as framelets and were 2.54 mm wide and over 55 mm long. At this rate, 10 minutes were required to read out one medium-resolution frame, and 34 minutes for one high-resolution frame. The transmitted light was sensed by a photomultiplier tube and the resulting electrical signal was mixed with synchronization and blanking pulses and fed to the communication system modulator for transmission to Earth. The video signal received on Earth was fed into the ground reconstruction electronics (GRE) where it was converted into an intensity-modulated line on the face of a cathode-ray tube. This line was used to expose 35-mm film in a continuousmotion camera to reconstruct the framelets. The scale of the reconstructed framelets (GRE scale) was 7.18 times spacecraft scale; the framelets were approximately 18 mm wide and 40 cm long. The framelets were then reassembled. Medium-resolution frames were reassembled in their entirety; high-resolution frames were reassembled into three component sections.

The video signal was also recorded on magnetic tapes which were subsequently used to make additional 35-mm framelets. These framelets had generally improved tonal qualities over the framelets reconstructed during the missions and were used to make master negatives for use by the NSSDC in providing copies to the public.

The film supply of the spacecraft consisted of 79 meters of unperforated 70-mm Kodak aerial film, type SO-243. This film

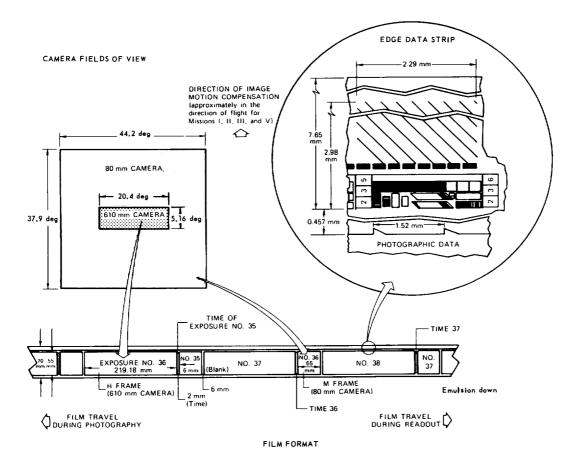


FIGURE 3 .- Spacccraft film format.

is a fine-grained, low-speed film with an aerial index of 3.0, which makes it relatively insensitive to space environment radiation. The film was provided with image-motion compensation (IMC) by a velocity/height (V/H) sensor which utilized the 610-mm lens. The V/H sensor also controlled the spacing of shutter operations during multiple-exposure sequences.

The full fields of view (shown in fig. 3) for the 80-mm camera and the 610-mm camera were 44.2° by 37.9° and 20.4° by 5.16°, respectively. The placement of the images of the two cameras on the spacecraft film is shown in figure 3. Images recorded by the 80-mm camera are referred to as medium-resolution frames (M frames); those recorded by the 610-mm camera are referred to as high-resolution frames (H frames).\* A folding mirror was employed in the optical path of the 610-mm camera and therefore the H-frame images are reversed left to right with respect to the M frames. Exposure times were recorded on the film by a binary-coded arrangement of lamps. These timing lights were located on the 80-mm camera platen and recorded exposure times to tenths of a second.

The angular resolution of the 610-mm cameras was 4.4 seconds of arc; for the 80-mm cameras, 34 seconds of arc. The resolution of the images recorded by both cameras was 76 lines/mm (spacecraft scale) which translates to an image resolution of approximately 11 lines/mm for the reconstructed 35-mm framelets (GRE scale). The ground resolution of vertical photographs taken from an altitude of 46 km is approximately 1 meter for high-resolution frames and 8 meters for medium-resolution frames.

Prior to launch, the spacecraft film was preexposed along one edge as shown in figure 3. The preexposures included calibration data which were used to monitor inflight system operation and to evaluate final data quality. The preexposed data array included a 0.3 background density to provide a reference level for setting readout gain, diagonal focus lines to indicate optimum readout scanning-spot focus, resolution charts to evaluate readout quality independent of camera image quality, a gray scale for sensitometric calibration, and an identification number. In addition, the film used on missions II, III, IV, and V was preexposed with a geometrical pattern extending across the entire format for geometric calibration purposes.

## **Photographic Mission Parameters**

Table 1 (p. 7) gives the flight log of the five Lunar Orbiter photographic missions and table 2 summarizes the photographic accomplishments. The orbits for each mission were ellipses with orbital parameters selected according to the various tasks of each mission. Mission-site location and the type and extent of required coverage were major considerations in determining the most suitable orbital parameters. Apollo landing sites were to be located within the equatorial region on the near side of the Moon and therefore the first three missions utilized close-in orbits inclined slightly to the lunar equator to provide optimum coverage of these areas. Perilune altitudes for these missions were as low as 44 km, limited primarily by uncertainties in execution errors in maneuvers, uncertainties in the lunar gravitational field and elevations, and the operating range of the V/H sensor. Missions IV and V were devoted to increasing scientific understanding of the Moon and utilized highly elliptical, near-polar orbits for access to areas at high latitudes with proper illumination. Each spacecraft orbited the Moon in the same sense as the rotation of the Moon.

During each mission, photography of the near side was conducted near perilune with morning illumination and photography of the far side near apolune with evening illumination. Photographs were sequenced by using one of two exposure interval rates. Spacecraft photographic maneuvers were based on the requirements that the camera axes must point directly at the target position at the midpoint of a sequence and that image-motion compensation is provided (when required) by proper orientation of the spacecraft with respect to the flightpath. These maneuvers usually consisted of a three-axis rotation from the normal Sun-Canopus celestial reference several minutes prior to the picture taking. Each photograph in a sequence would be taken with the camera axes in the same reference after which the spacecraft would be returned to the celestial reference. Since photographs could be taken much faster than they could be processed, a looper having a capacity of 20 dual frames acted as a buffer between the cameras and the readout section. Site photography proceeded from east to west as the Moon rotated under the stationary, inertially fixed orbit of the spacecraft.

## Photographic Coverage

#### AREAS OF PARTICULAR INTEREST TO APOLLO PROGRAM

Photography during Lunar Orbiter missions I, II, and III was conducted primarily to locate and confirm suitable manned landing sites for the Apollo program. The requirements for these sites were as follows:

Zone of interest.—The sites had to be located within the zone specified by the Apollo program;  $\pm 45^{\circ}$  longitude and  $\pm 5^{\circ}$  latitude.

Site locations.—Multiple sites which permitted at least three launch opportunities within any Apollo launch window had to be located. Launch opportunities were anticipated to occur on alternate days; thus, suitably lighted sites separated in longitude by 23°±3° were required. In addition, the capability to launch during each month of the year required sites to be located along both the northern and southern portions of the zone mentioned.

Site characteristics.—Apollo landing sites had to cover an elliptical area with major and minor axes approximately 8 km and 5 km, respectively, and had to be relatively free of protuberances, depressions, or slopes that would constitute a hazard to the Apollo landing vehicles. The landing-approach terrain was to be reasonably unmodulated to accommodate the guidance system of the vehicle.

Areas were originally selected from Earth-based observations that appeared to offer candidate Apollo landing sites. Areas photographed as candidate landing sites were designated primary (P) sites. Prescheduling of P sites left certain periods when photographs had to be taken in order to satisfy film-handling constraints. Sites photographed in compliance with this constraint were designated secondary (S) sites; it should be noted, however, that this designation had relevance only with respect to the mission objectives and the mission plan and not to the value of the photography.

Mission I photography was conducted from three different orbits characterized by the parameters listed in table 1. Nine primary sites concentrated in the southern part of the Apollo zone were photographed. The 610-mm camera failed to operate satisfactorily at close range and consequently most of the high-resolution frames were smeared. The medium-resolution frames were of good quality and provided coverage of extensive areas with an increase in resolution of two orders of magnitude over astronomical photographs. Secondary-site photography provided coverage of numerous areas along the equatorial region on both the near and far sides. The 610-mm

<sup>\*</sup>Other terms used in literature in referring to Lunar Orbiter photographs are for medium-resolution frame: wide-angle frame, low-resolution frame, moderate-resolution frame; for high-resolution frame: telephotoframe. This paper uses the words "photograph" and "frame" interchangeably.

camera generally operated satisfactorily during photography of the far side; consequently, both the medium-resolution and high-resolution photographs of these areas were of excellent quality. Ground resolution of the high-resolution frames was approximately 30 meters. Two oblique exposures (nos. 102 and 117) were taken during mission I. Both are very similar in nature, cover approximately the same area, and show views of the crescent Earth and part of the far side of the Moon just beyond the eastern limb (as seen from Earth). In each case both the medium-resolution frame and the high-resolution frame are of good quality.

Mission II photography was conducted from a single orbit having the parameters listed in table 1. Photographic targets were concentrated in the northern part of the Apollo zone of interest and included 13 primary sites and 17 secondary sites. Most of the primary sites were photographed by taking multiple-exposure sequences during consecutive passes. The secondary sites provided equatorial coverage of areas near the equator on both the near and far sides. With the exception of a few photographs which were incompletely read out, no problems were encountered with the operation of the photographic system

During missions I and II, all the primary sites were photographed by using standard techniques; that is, coverage was obtained by taking sequences of 4, 8, or 16 vertical photographs during one or more passes of the spacecraft over the site. This procedure provided stereoscopic medium-resolution coverage and high-resolution coverage useful primarily for interpretation and photometric analysis. An experiment conducted during mission II determined that even better stereoscopic coverage could be obtained with the 610-mm camera by photographing the same area during two consecutive passes with the camera axes tilted during one of the passes; this type of photography is referred to as convergent photography. The success of this experiment contributed to the decision that mission III would be a site-certification mission. It provided additional coverage of the most promising candidate Apollo landing sites photographed during missions I and II. Other factors which compelled this decision were the need for makeup high-resolution coverage of areas inadequately covered during mission I and the desire to obtain oblique views of the landing sites to simulate the views which would exist during a manned descent to the surface. The 12 primary sites photographed during mission III included 5 areas previously photographed during mission I, 5 areas previously photographed during mission II, and 2 proposed Surveyor landing sites which had been selected on the basis of Earth-based photography. To photograph all these areas under favorable lighting conditions, the inclination of the orbit to the lunar equator was increased from the 12° value used for missions I and II to a value of 21°. Whereas missions I and II employed standard photographic techniques, diverse techniques were used during mission III to take the required vertical, oblique, and converging coverage. Unfortunately, the spacecraft developed trouble with its film advance motor late in the mission and was unable to read out a substantial number of photographs. This mission provided the Apollo program with sufficient information, however, to allow the remaining two missions to concentrate on more expanded scientific objectives.

In summary, 22 areas were photographed during the first three missions in search of Apollo landing sites. On the basis of this photography and data obtained from Surveyor I, eight candidate Apollo landing sites were selected. Although all three types of required coverage had been obtained at only three of these sites, additional coverage obtained later during mission V enabled complete certification of all sites. Table 3 summarizes all photography during missions I, II, III, and V taken

in search of Apollo landing sites. The most promising areas were those photographed with all three types of photography indicated. Entries for "Area of interest" begin with the east-ernmost site and progress westward.

#### AREAS OF GENERAL INTEREST

Most of the photographs taken during missions I, II, and III were of areas which by nature of their potential use as manned landing sites were smooth and featureless. Although the secondary-site photography of these missions included a substantial number of areas interesting from the standpoint of geology and resulted in some very spectacular views, this photography was scheduled "around" the primary-site photography and was limited to areas located near the equator. The converse was true for missions IV and V, whose primary objective was to increase man's scientific understanding of the Moon.

Mission IV was assigned the task of performing a broad systematic survey of lunar-surface features in order to increase the scientific knowledge of their nature, origin, and processes, and to serve as a basis for selecting sites for more detailed scientific study by subsequent orbital and landing missions. Photography was planned on the basis of the coverage to be obtained by the 610-mm camera. It was desired to obtain vertical high-resolution photographs which would provide monoscopic coverage of the entire near side with a minimum of overlap. This coverage was obtained by taking 5 singleframe sequences on each of 29 consecutive passes. The orbit was highly inclined to the equator (85°) and had a perilune altitude, at the equator, of approximately 2700 km. The spacecraft was oriented with the long dimension of the frames in a north-south direction. Pole-to-pole coverage was obtained by taking, on each pass, four vertical photographs symmetrically spaced about the equator for coverage of the equatorial and temperate regions, and a fifth photograph for coverage of the polar regions. The fifth photograph was used alternately from pass to pass for coverage of the south- and north-polar regions. It was taken slightly off vertical for lighting considerations.

The near-side photography covered the equatorial regions with ground resolutions of approximately 60 meters and the polar regions with ground resolutions of approximately 100 meters. The field of the 80-mm camera encompassed nearly the entire lunar disk. The ground resolution of the medium-resolution frames is comparable to the best obtainable from astronomical photography—on the order of  $\frac{1}{2}$  kilometer.

Many photographs, taken early in the mission, were severely degraded during a period when a thermal door to the cameras failed to operate properly. In some cases the door failed to open and therefore the expected photographs were unexposed. In other cases, the photographs were degraded because of condensation on the camera windows. All the areas covered by these degraded photographs were rephotographed toward the end of the mission by six sequences taken near apolune. As in the case of all photographs taken near apolune, these photographs were taken with evening illumination.

Mission IV coverage of the far side was obtained by five sequences taken near apolune and by a number of the near-perilune sequences. The photographs taken near apolune consisted of seven medium-resolution frames (two of which were severely degraded); the high-resolution frames covered essentially unilluminated areas. Medium-resolution frames taken near perilune (with morning illumination) provided the more significant far-side coverage during mission IV. The photographs taken on the first pass covered extensive areas beyond  $+90^{\circ}$  longitude, and each medium-resolution frame taken on the polar sequences, although centered on the near side, provided coverage which extended beyond the polar caps and on to the far side.

The primary objective of mission V was to photograph 36 areas of particular scientific interest on the near side. Photography was also required to complete the Apollo requirements and to complete the far-side coverage. This combination of requirements necessitated two orbital changes. The orbital parameters are given in table 1. Photographic altitudes for the near side were on the order of 100 km to 250 km. These altitudes, which were two to five times greater than those used for the near-side photography during missions I, II, and III, were required in order to provide adequate areal coverage and acceptable ground resolution of each of the numerous sites with the limited film supply of the spacecraft. In addition, most of the remaining Apollo requirements were for converging coverage. Thus, the increase in altitude was desirable, since it enabled these photographs to be taken with less cross-track tilt than had been utilized previously.

Mission V was executed precisely as planned and accomplished each of its assigned objectives. One dual frame was also taken which shows a view of a nearly full Earth.

In summary, photography for purposes other than locating or confirming Apollo landing sites was taken during each of the five missions. It consists of low-altitude photography of the near side taken during missions I, II, III, and V; and highaltitude photography taken during each of the missions. The low-altitude photography provided detailed coverage of 88 areas from altitudes ranging from approximately 44 to 250 km; this photography is summarized in table 4. The photography at a selected number of these areas is summarized in table 5; features photographed are in alphabetical order. The high-altitude photography provided broad coverage of essentially the entire Moon from altitudes ranging from approximately 1350 to 6000 km. Whereas mission IV alone provided the broad coverage of the near side, each mission contributed to the broad coverage of the far side.

## MAPS

Figures 4 to 11 are small-scale maps showing all the areas photographed during each mission, with the exception of the areas covered by the medium-resolution frames from mission IV. Figure 4 (p. 22) is a composite plot for missions I, II, III, and V and indicates the missions during which any given area was photographed. Figures 5 to 9 break down the coverage shown in figure 4 and present separately the coverage obtained during each mission. The photographic coverage obtained during mission IV is shown in figures 10 and 11.

Figures 4 to 11 show, where the scale permits, the areas covered by individual photographs. Where scale limitations precluded showing these areas, they show only the envelope of the total coverage at each site. For each of these sites, the areas covered by the individual photographs are shown in figures 12 to 15, which are large-scale maps.

Thus, figures 4 to 15 permit one to determine all photographs covering a given area. One should first consider the coverage of missions I, II, III, and V and, secondly, that of mission IV.

Missions I, II, III, and V.—Figure 4 shows the total area photographed during each of these missions. It should be noted that this figure presents only the envelope of the total coverage by a given mission in any region. Where an area of nearvertical or converging coverage is contained within an area of oblique coverage photographed during the same mission, only the boundary of the oblique coverage is indicated. Thus, the boundaries of coverage for sites IIP-8, IIIP-7, IIIP-8, IIIP-10, IIIP-11, IIIP-12, IIIS-15, IIIS-16, V-8, V-11, V-12, V-16, and V-18 cannot be separately identified in figure 4. However, they are individually outlined in figures 5 to 9, which show the area covered at each site with the Lunar Orbiter site designa-

tion. Figures 5, 6, and 7 pertain to missions I, II, and III, respectively; figures 8 and 9 pertain to mission V. The coverage shown is the envelope of coverage of the medium-resolution frames for the near-side sites, and with the exception of mission I sites IS-3 and IS-9, the coverage of individual photographs for the far-side sites.

Table 6 gives the exposures allocated to each site for these missions; table 7 is a permuted form of table 6, and indicates the site to which each exposure was assigned. Table 8 lists the mission I, II, III, and V sites for which photographs were incompletely read out or degraded.

Figures 12, 13, 14, and 15 are photographic indexes of all near-side sites, except site IS-1, for missions I, II, III, and V, respectively. They show individual photographic outlines portrayed on the U.S. Air Force Aeronautical Chart and Information Center (ACIC) series of lunar charts (Lunar Aeronautical Charts (LAC) or Apollo Intermediate Charts (AIC)). The photographic outlines are accompanied by numbers which uniquely identify the photographs and which should be used in ordering photographs from NSSDC. The photographic outlines were determined by ACIC personnel who matched the photographic images to the shaded relief features on the charts. Thus, the inferred coordinates of the corners of the photographs and the features contained therein are only as accurate as the charts.

At many sites, especially the sites photographed for Apollo, the high-resolution frames have not been indexed. They were not indexed because there was insufficient detail on the base maps with which to make an image match, the photographs were either not read out or were degraded, or their inclusion would have cluttered the figure. The approximate coverage of these frames can be determined, for vertical or near-vertical photography, by scaling the fields of the two cameras, shown in figure 3, to the map scale of the photographic index. To determine which photographs were degraded or incompletely read out, reference should be made to table 8.

At multiple-exposure sites, the exposure numbers increase from west to east for missions I, II, and III sites (figs. 12, 13, 14), and from south to north for mission V sites (fig. 15). The maps are oriented in the standard aeronautical convention with north at the top. They incorporate the selenographic coordinate system with east (positive) and west (negative) longitudes measured from the central meridian at Sinus Medii and the longitudes increase in magnitude to 180 at the center of the far side.

Mission IV.—Whereas missions I, II, III, and V were assigned to photograph selected areas, mission IV was assigned to photograph broad areas and to cover the entire near side. Both medium- and high-resolution frames from mission IV cover the entire near side, and the medium-resolution frames provide the only coverage of some regions of the far side. The maps of coverage of these photographs are presented independently of those from the other missions.

With the exception of two small areas near the poles, any area which figure 4 indicates was not photographed during missions I, II, III, and V was photographed during mission IV. Figure 10 shows the area covered by each mission IV high-resolution frame. Any area for which neither figure 4 nor figure 10 indicates as having been photographed was covered only by mission IV medium-resolution frames. Figure 11 shows the area covered by a selected number of these photographs (or portions thereof). In most cases the near-side areas covered by these photographs are not shown. The outlines shown figures 10 and 11 are accompanied by the appropriate exposure number.

Table 9 gives the selenographic distribution of mission IV exposures. Table 10 is a permuted form of table 9 and indicates

the site to which each exposure was assigned. Table 11 summarizes all mission IV photographs incompletely read out or degraded.

Map summary.—For any given area, the photographs covering that area are determined as follows:

Refer to figure 4 to determine whether the area was photographed during missions I, II, III, and V and, if so, during which missions(s). Then, depending on the mission(s), refer to the appropriate figure(s) among figures 5 to 9 to determine the site(s). If the area in question is on the near side, the site number is used to locate the photographic index for that site in figures 12 to 15. If the area in question is on the far side, refer to table 6 to determine the exposure number(s).

Refer to figures 10 and 11 to determine whether the area was photographed during mission IV and, if so, by which photograph(s). These figures show the area covered by individual photographs (and the exposure number) for all the high-resolution frames, but only for a selected number of medium-resolution frames. The approximate locations of the principal ground point and condition of the photograph, for the medium-resolution frames not considered in figures 10 and 11, are given in tables 9 and 11, respectively.

## Copies of Photographs

Each Lunar Orbiter spacecraft was supplied with sufficient film to record as many as 426 photographs—213 pairs of medium-resolution and high-resolution frames. The negative images on the spacecraft film were read out in parts, termed "framelets," and reconstructed on Earth on 35-mm film as positive images of the Moon at a scale (GRE scale) 7.18 × spacecraft scale. The framelets were then used to make reassembled frames in various forms.

#### 20- BY 24-INCH SECTIONS

The framelets reconstructed in the GRE represent the original flight data and are designated as zero-generation positives. The original framelets (or copies) were reassembled and contact printed on to 20- by 24-inch sheet film. One medium-resolution frame required just one 20- by 24-inch section, whereas the high-resolution frame required three component sections. All Lunar Orbiter photographs have been reassembled into a 20- by 24-inch format, with the exception of the smeared high-resolution frames of mission 1.

By using second-generation duplicate positives of the original flight data, the U.S. Army Topographic Command (TOPOCOM) prepared third-generation 20- by 24-inch master negatives for all photographs from missions III, IV, and V and for the high-resolution frames from mission II. These negatives were made to provide quick copies for Government agencies for interpretation and mission planning and to provide the National Space Science Data Center (NSSDC) with material from which early copies could be made generally available to the scientific community.

At the completion of the Lunar Orbiter program, the NASA Langley Research Center (LRC) produced an improved set of 20- by 24-inch negatives from which high-quality copies could be made and disseminated by the NSSDC. The video tapes were used to generate a new set of positive framelets which had generally improved tonal qualities over those secured during the missions. These positive framelets were made by electronic preprocessing of the video signal prior to input to the GRE. (However, because the video signal was intentionally distorted prior to input to the GRE, the 35-mm film exhibits density variations which are not accurate representations of the true lunar reflectance properties and should not, therefore, be used for densitometric or photometric analysis.) The positive

framelets thus obtained were reassembled and contact printed on to 20- by 24-inch sheet film to make first-generation master negatives. This procedure was followed for all photographs except those not graded A or B in tables 8 and 11. Each 20by 24-inch section is labeled with a photo number consisting of mission number, a Roman numeral; exposure number, an Arabic numeral; and frame type, M (medium-resolution frame) or H (high-resolution frame). Sections of high-resolution frames are additionally labeled with subscripts 1, 2, or 3 following the photo number to distinguish the component sections. For example, the sections labeled V-141M and V-141H<sub>2</sub> are mission V medium-resolution frame no. 141 and the center section of mission V high-resolution frame no. 141, respectively. For the photographs listed in table 12, the video tapes were replayed additional times to produce 35-mm film with optimum detail in the highlight areas or the lowlight areas. The photographs made from reassemblies of this film are additionally labeled with "SP," indicating a special play for highlight areas, or "SP-1," indicating a special play for lowlight areas.

Table 13 gives some characteristics of Lunar Orbiter vertical photographs. Values given for the photographic scale apply to the 35-mm framelets reconstructed in the GRE and also to the 20- by 24-inch sections. The ground resolutions given are in direct proportion to the altitudes given. The reassembly code given for the high-resolution frames is useful for orienting the photographs. The long axis of all photographs is oriented either in a primarily north-south or an east-west direction. With the edge data at the top, the left, center, and right sections (of a high-resolution frame for those frames reassembled at the Langley Research Center (LRC)) are numbered 1, 2, and 3, respectively. The reassembly code tells which of these sections provides the northernmost (N) or easternmost (E) coverage. It applies only for frames reassembled at LRC. (TOPOCOM numbered the three-component sections of a highresolution frame in the reverse order: sections 1, 2, and 3 in the LRC convention are sections 3, 2, and 1, respectively, in the TOPOCOM convention.)

## SOURCE OF COPIES

The results of all space science flight experiments are made available through the National Space Science Data Center (NSSDC). Copies of all Lunar Orbiter photographs and background information including photographic system calibrations and photographic supporting data are available from the NSSDC. For further information, scientists located within the United States should address their inquiries to—

National Space Science Data Center Code 601.4 Goddard Space Flight Center Greenbelt. Md. 20771

Scientists from abroad, to-

World Data Center A Rockets and Satellites Code 601 Goddard Space Flight Center Greenbelt, Md. 20771, U.S.A.

In ordering copies, the photographs should be specified by mission number, exposure number, and frame type (M or H). When interested in a particular section of a high-resolution frame, the position of that section relative to the central section—northern, eastern, etc.—should be stated. The quantity of Lunar Orbiter photographs available from the NSSDC, in terms of 20- by 24-inch sections, is given in table 14.

 ${\bf TABLE~1.} \color{red} -Lunar~Orbiter~Flight~Log$ 

	Mission	Mission	Mission	Mission	Mission
	I	II	III	IV	V
Launch: Date Hr:min (GMT)	8/10/66	11/6/66	2/5/67	5/4/67	8/1/67
	19:26	23:21	01:17	22:25	22:23
Injection into lunar orbit: Date Hr:min (GMT)	8/14/66	11/10/66	2/8/67	5/8/67	8/5/67
	15:43	22:58	22:03	15:17	16:49
Photographic dates: First exposure Last exposure	8/18/66	11/18/66	2/15/67	5/11/67	8/6/67
	8/29/66	11/25/66	2/23/67	5/25/67	8/18/67
Mission termination:  Date of impact  Hr:min (GMT)  Impact location:  Longitude, deg  Latitude, deg	10/29/66 13:29 160.71 E 6.35 N	10/11/67 07:17 119.13 E 2.96 N	10/9/67 10:27 92.70 W 14.32 N	*7/17/67 06:30 ≈26 W	1/31/68 07:58 83.04 W 2.79 S
Orbital parameters utilized for photography: First set: Perilune altitude, km Inclination, deg Period (hr:min) Exposures taken Second set: Perilune altitude, km Apolune altitude, km	189 1866 12.16 3:37 5 to 42 56 1853	50 1853 11.89 3:28 5 to 215	55 1847 20.91 3:28 5 to 215	2706 6114 85:48 12:01 5 to 196	195 6028 85.01 8:27 5 to 22
Inclination, deg Period (hr:min) Exposures taken Third set:	12.05 3.29 44 to 133	}	Not applicable		85.61 8:21 24 to 30
Perilune altitude, km Apolune altitude, km Inclination, deg Period (hr:min) Exposures taken	40 1817 12.00 3:26 134 to 215	}	Not applicable		99 1500 84.76 3:12 31 to 217

<sup>\*</sup>Last communication with spacecraft. Date of impact estimated at 10/31/67.

 ${\bf TABLE~2.} \color{red} - Number~of~Photographs~Obtained$ 

					Number of photogr	aphs obtained —		
			Med	lium-resolution fr	ames	н	igh-resolution fra	mes
Mission	Number of sites	Number of exposures	Useful for in	terpretation	Not useful for	Useful for in	terpretation	Not useful for
		!	Complete frames	Partial frames	interpretation or not read out	Complete frames	Partial frames	interpretation or not read out
			Ap	ollo				
Mission I	9	136	136	0	0	0	0	136
Mission II	13	184	180	1	3	175	7	2
Mission III	18	162	116	4	42	102	31	29
Mission V	9	44	44	ō	0	44	0	0
Subtotal	49	526	476	5	45	321	38	167
			General inter	est; near side	_ <del>  </del>			
35: 1 Y	18	57	57	0	0	6	0	51
Mission I	13	23	23	0	0	23	ŏ	0
Mission II	23	47	35	1	11	35	3	9
Mission III	23 36	130	129	0	1 1	129	1 1	0
Mission V				ļ <u>.</u>			9	24
Mission IV (subtotal)_	148	165	117	3	45	132	9	
Missions I, II, III V (subtotal)	90	257	244	1	12	193	4	60
			General inte	rest; far side	<u> </u>	<u> </u>		
		11	11	0	0	6	0	5
Mission I	2 4	4	4	0	0	4	ŏ	ő
Mission II	2	2	1	0	1	1	ì	Ŏ
Mission III	6	7	5	0	2	0	ō	7
Mission IV	23	37	37	0	0	35	Ö	2
Mission V				<u> </u>				14
Subtotal	37	61	58	0	3	46	1	14
			Ea	rth	<del>-, ,</del>		T.	_
Earth	<b></b> _	3	3	0	0	3	0	0
Grand total	324	1012	898	9	105	695	52	265

Table 3.—Photography in Search of Apollo Landing Sites

## (a) Photographic information

hotography looking)	Oblique pl (west l	phy	erging photogra	Conv		ind near-vertical otography				Area of interest	
Expos	Site	Convergent sequence, exposures	Near- vertical sequence,	Site	Note reference in table 3(b)	Second sequence, exposures	First sequence, exposures	Site	ion	Approx locati	Vicinity of search site*
↓		CAPOSITES	exposures		table 3(b)	exposures	exposures		Latitude	Longitude	site*
38 (ze pha phot	V-3.1	33 to 36	35 to 32	IIIP-2	2		52 to 67	IP-1	1° S	42° E	IP-1
42	V–6	48 to 51	44 to 47	V-8	<b>4</b> 6		25 to 32 44 to 47	IIIP-2			
42	v –6				2		5 to 20	V-8 IIP-1	4° N	070 E	TTD 4
1					2		68 to 83	IP-2	4° N	37° E	IIP-1
52	V-9.1	59 to 62	55 to 58	V-11	2		35 to 42	IIP-2	2° N	36° E 34° E	IP-2 IIP-2
32	V-3.1	33 10 02	00 10 30	V-11	4		5 to 20	IIIP-1	2 IN	34 E	11P-2
64	V-13	52 to 59	60 to 67	IIIP-5	2		85 to 100	IP-3	1° N	26° E	IP-3
"		75 to 78	71 to 74	V-16	2	84 to 91	76 to 83	IIP-6	1 11	20 15	11-3
1		10 10 10	11 10 14	. 10	4	04 00 01	44 to 51	IIIP-4		}	
					4		60 to 67	IIIP-5		1	
]					2		67 to 74	IIP-5	3° N	25° E	IIP-5
]					2	51 to 58	43 to 50	IIP-3	4° N	21° E	IIP-3
]					5		40 to 43	HIP-3	2° N	21° E	IIIP-3
]					5		68 to 71	IIIP-6	0°	21° E	IIIP-6
_					2		59 to 66	IIP-4	4° N	16° E	IIP-4
_					3		105 to 112	IP-4	0°	14° E	IP-4
84	IIIS-11	86 to 93	94 to 101	IIIP-7	2		118 to 133	IP-5	0°	1° W	IP-5
		112 to 115	108 to 111	V-27	2	121 to 128 (129 to 136,	113 to 120	IIP-8			
						3d sequence)		l			
4			·		4		94 to 101	IIIP-7			
-			·- <b>-</b>		3		141 to 148	IP-6	4° S	2° W	IP-6
1					2	104 to 111	96 to 103	IIP-7	2° N	2° W	IIP-7
1	TITO 01				2	454 4 450	138 to 145	IIP-9	1° N 0°	13° W	IIP-9
120	IIIS-21				2	171 to 173	163 to 170	IIP-11	0,	20° W	IIP-11
136	IIIS-24	137 to 144	145 to 152	IIIP-9	4 2		124 to 131 157 to 172	IIIP-8 IP-7			
100	1115-24	157 10 144	145 to 152	1117-9	4	153 to 160	145 to 152	IIIP-9	3° S	22° W	
1					2	154 to 161	146 to 153	IIP-9 IIP-10	3- 5	22° W	IP-7
1				<b></b>	2	187 to 194	179 to 186	IIP-10 IIP-12	3° N	27° W	IIP-10
171	IIIS-27	173 to 176	169 to 172	V-42	2	187 10 194	176 to 183	IP-8.1	3° N	34° W	IIP-10 IIP-12
1 111	1110-21	113 10 110	105 10 172	v —— 2	4		173 to 180	IIIP-11	3° S	36° W	IP-12 IP-8.1
161	IIIS-25	163 to 170	205 to 212		2	205 to 212	197 to 204	IIP-13	2° N	41° W	IIP-13
1	4110-20	(IIIP-10)	(IIP-13)		-	400 10 212	131 10 204	111-10	2 14	41 11	111-10
172	IIIS-28	205 to 212	185 to 192	IIIP-12	2	200 to 215	184 to 199	IP-9.2	2° S	43° W	IP-9.2
1	1110-20		130 00 132	1111 - 12	-	185 to 200	181 to 184	IIIP-12	2 0	20 11	11-5.2
			İ			(201 to 204,					
1						3d sequence)					

<sup>\*</sup>Search site: Area of vertical coverage photographed in search of Apollo landing sites. Candidate Apollo landing sites selected on the basis of this photography were certified by the additional vertical, converging, and oblique photography listed. Search sites which

did not reveal areas suitable for Apollo were not rephotographed except in an incidental manner. These areas of common coverage are not indicated in this table but may, however, he determined by reference to the index maps.

## (b) Vertical photography

	Exposure		ic coverage, lution frames	Monoscopic coveras fram	
Note reference	internal rate	Forward overlap, percent	Lateral overlap,* percent	Forward overlap, percent	Lateral overlap,* percent
2	Fast	88	66	Continuous	11
3	Slow	52	66	Discontinuous	11
4	Fast	88	42	Continuous	Discontinuou
5	Slow	52	42	Discontinuous	Discontinuou
6	Fast	88	Not applicable	Continuous	Not applicable

<sup>\*</sup>Lateral overlap given for photographs taken on adjacent sequences.

Site		ate center erage Latitude	Type of photography *	Remarks	Site	Approxim of cov	nate center verage Latitude	Type of photography	Remarks
	Longituae	Latitude							
IS-1 <sup>b</sup>	90° E	1° S	NV, 16f, 4f	Mare Smythii	V-29	3° W	12° N	NV, 4f	Rima Bode II
IS-2b	72° E	2° N	NV		IS-12b	5° W	3° N	NV	N
V-1	61° E	26° S	NV, 4f	Petavius	IIIS-15	6° W	0°	NV, 4s	Near Schröter, north of Mösting
IS-4b		1° N	NV		IIIS-16	6° W	0° 3° S	NV NV. 4s	Mösting Mösting C
V-2.1	58° E	20° S	NV	Petavius B	IIIS-18	8° W		Oblique	Candidate Surveyor site
V-4	53° E	32° S	NV	Stevinus A	IIIS-14	8° W	6° N 42° S	NV. 4f	Tycho
IS-5b	50° E	2° N	NV,*	Taruntius	V-30 IIS-10.2_	11° W	13° N	NV NV	Gambart C, thermal anomaly
IIIS-1b	47° E	1° S	NV, 4f	Messier and Messier A	V-32	11° W	13° N	NV.4s	Erathosthenes
V-5.1	42° E	2° S	Oblique	Messier	IS-13b	15° W	2° N	NV	Gambart
IIS-1	42° E	3° N	NV, 4f		V-33	15° W	6° N	NV	Area of Copernicus CD
IS-6b	40° E	3° N	NV*		V-34	16° W	8° S	NV, 4f	Fra Mauro
IIS-2	37° E	3° N		Experiment on convergent	V-35		14° N	NV. 4s	Copernicus secondaries
			ļ	photography	IIIS-23_		4° S	NV, 4s	Fra Mauro
V-12	34° E	1° S	NV	Censorinus	IS-15*		0.	NV	
IS-7b	32° E	5° N	NV*	I-48M shows domes near	V-36	18° W	7° N	NV, 4f	Copernicus H
				Maskelyne A	IS-14	20° W	1° N	NV	Soporations ==
V-14	30° E	22° N	NV, 4f	Littrow	IIS-12		10° N	Oblique	Copernicus, northerly oblique
IIIS-8		14° S	Oblique	Theophilus	V-37		10° N	NV.8f	Copernicus
V-10	26° E	30° S	Oblique	Altai Scarp	IS-19b	22° W	5° S	NV	•
V-15.1	1	17° N	NV	Dawes	IIIS-22	22° W	1° N	NV	Candidate Surveyor site
IIIS-5b	25° E	1° S	Oblique	Moltke	V-38	22° W	33° N	NV, 4f	Imbrian flows
V-18	19° E	2° N	NV, 4f	Dionysius	IS-16b	24° W	0°	NV	Near Reinhold, grooves and chain
IIIS-9	. 18° E	2° S	NV	Delambre					craters radial to Copernicus
IS-8b	17° E	3° N	NV	Dionysius	IIS-11	27° W	4° N	NV	Southwest of Copernicus near
V-19	14° E	15° S	NV	Abulfeda crater chain					Hortensius
IIIS-10	. 14° E	2° S	NV, 4s	Candidate Surveyor landing site	IIIS-20	27° W	12° N	Oblique	Hortensius domes
V-21		39° N	NV, 4f	South of Alexander	IS-17b		1° N	NV	
IIS-8		0°	NV		V-40	31° W	12° N	NV, 4f	Tobias Mayer dome
IS-10b	. 9° E	2° N	NV		IS-21b	35° W	4° S	NV	
IIIS-6b	. 9° E	13° N	Oblique	Hyginus Rilles	IS-18b	36° W	0°	NV	
V-22	9° E	20° N	NV, 4f	Sulpicius Gallus Rilles	IIIS-26.		11° N	Oblique	Kepler
IIIS-75	7° E	4° N	NV, 4s	Vicinity of Dembowski	V-41		31° S	NV	Vitello
V-23.2	6° E	8° N	NV, 4f	Hyginus Rilles	V-43.2	1	18° S	NV, 4f	Gassendi
IIS-6	. 5° E	4° N	NV	Near Triesnecker	V-45.1	41° W	36° N	NV, 4f	Jura Domes
IIIS-17	. 4° E	5°S	NV,4s	Candidate Surveyor site,	IIS-13		3° N	NV	Braided ridge southwest of Kepler
_				floor of Hipparchus	V-46		27° N	NV, 8f	Harbinger mountains Aristarchus
V-24	4° E	5° S	NV, 4f	Hipparchus	V-48	47° W	23° N	NV, 8f	Cobra Head
V-26.1	4° E	36° N	NV, 4s	Hadley Rille Murchison and Pallas	V-49 IS-20b		25° N 4° S	NV, 4f NV	Cobra Head
IIIS-13	. 2° E	11° N 2° N	Oblique	Sinus Medii, southwest of	IIS-20°		12° N	Oblique	Marius, northerly oblique
IIS-9	1°E	Z N	NV	Triesnecker	V-50	52° W	28° N	NV, 4f	Aristarchus plateau
IIS-7	1° W	3° S	Oblique	Sinus Medii, southerly oblique	IIS-16		3° N	NV	South of Reiner
V-25		46° N	Oblique	Alpine valley	IIS-17		13° N	Oblique	Reiner Gamma
V-31	2° W	50° N	NV, 4f	Sinuous rille east of Plato	V-51	56° W	13° N	NY, 8f	Marium Hills
V-28	3° W	14° S	NV,4f	Alphonsus	HIS-29	62° W	10° S	Oblique	Damoiseau
IIIS-19	. 3° W	4° S	NV, 4s	Candidate Surveyor site,	IIIS-30		12° N	Oblique	Cavalerius, Luna 9 area
				Flammarion	IIIS-31_	67° W	1° N	NV	Floor of Hevelius

<sup>a</sup> Type of photography: NV, vertical or near vertical

Oblique

overlap between consecutive medium-resolution frames and discontinuous highresolution coverage.

xxf or xxs (applies to multiple exposure sequences). The number of exposures taken per sequence followed by the exposure interval rate; f, fast rate to give 88 percent forward overlap between consecutive medium-resolution frames and continuous high-resolution coverage; and s, slow rate to give 52 percent forward

b Sites at which photographs were incompletely read out or secured in degraded form. See table 8.

Photographs of these sites were taken on separate orbits having different orbital parameters. Although they were taken independently of each other, they provided continuous coverage of specific areas.

 ${\bf Table}\ 5. {\bf ---} Sites\ of\ Selected\ Areas\ of\ Special\ Interest$ 

Feature	Site	Feature	Site
Near-vertical photograph	у	Near-vertical photography—Conti	nued
Abulfeda Alphonsus Aristarchus Aristarchus Plateau Rima Bode II Censorinus Cobra Head Copernicus Copernicus CD Copernicus H Copernicus Secondaries Dawes Delambre Dionysius Dionysius Eratosthenes Fra Mauro Gambart Gambart C Gassendi Hadley Rille Harbinger Mountains Hevelius (floor) Hipparchus Hyginus Rille Imbrian Flows Jura Domes	V-19 V-28 V-48 V-50 V-29 V-12 V-49 V-37 V-33 V-36 V-35 V-15.1 IHIS-9 V-18 IS-8 V-32 V-34 IHIS-23 IS-13 IHS-10.2 V-43.2 V-46.1 V-46 IHIS-31 V-24 V-23.2 V-38 V-28 V-28 V-28	Sinus Medii Messier Messier A Moltke Mösting C Petavius Petavius B Mare Smythii Stevinus A Sulpicius Gallus Rilles Taruntius Tobias Mayer Dome Tycho  Oblique photography  Alpine Valley Altai Scarp Cavalerius Copernicus Damoiseau Hortensius Domes Hyginus Rille Kepler Marius Hills Sinus Medii	IIS-9 IIIS-1 IIIS-1 IIIS-1 IIIS-5 IIIS-16 IIIS-18 V-1 V-2.1 IS-1 V-4 V-22 IS-5 V-40 V-30  V-25 V-10 IIIS-30 IIS-12 IIIS-29 IIIS-26 IIIS-26 IIS-15 IIS-7
LittrowMarius Hills	V-14 V-51	Messier Murchison Pallas Reiner Gamma Theophilus	V-5.1 IIIS-13 IIIS-13 IIS-17 IIIS-8

Table 6.—Exposures Allocated to Each Site for Missions I, II, III, and V

Mi	ssion I	Mis	sion II	Miss	ion III	Mission V		
Site	Exposures	Site	Exposures	Site	Exposures	Site	Exposures	
			Near	side				
S-8* S-10* S-12* S-13*	68 to 83 85 to 100 105 to 112 118 to 133 141 to 148 157 to 172 176 to 183 184 to 215 5 to 24 25, 26, 27 29, 33, 34 31, 32, 44 41, 50, 51 42, 46 to 49 84 103 113, 114 134, 135 137, 139, 140 138 149, 151 150 153 to 156 173	IIP-1* IIP-2 IIP-3 IIP-4 IIP-6 IIP-6 IIP-10 IIP-11* IIP-13 IIP-13 IIS-1 IIS-1 IIS-2 IIS-6 IIS-8 IIS-9 IIS-10.2. IIS-11 IIS-12 IIS-13 IIS-13 IIS-14 IIS-15 IIS-16 IIS-16 IIS-17	5 to 20 35 to 42 43 to 58 59 to 66 67 to 74 76 to 91 96 to 111 113 to 136 138 to 145 146 to 161 163 to 178 179 to 194 197 to 212 21 to 24 25 to 32 92 93 94 95 112 1137 162 195 213 214	HIP-1*   HIP-2*   HIP-3*   HIP-4*   HIP-5*   HIP-6*   HIP-7   HIP-6*   HIP-10   HIP-11   HIP-12   HIS-1*   HIS-5*   HIS-6*   HIS-7*   HIS-6*   HIS-7*   HIS-10   HIS-11   HIS-13*   HIS-15*   HIS-15*   HIS-16*   HIS-16*   HIS-17*   HIS-16*   HIS-17*   HIS-16*   HIS-16*   HIS-17*   HIS-16*   HIS-16*   HIS-17*   HIS-16*   HIS-17*   HIS-16*   HIS-17*   HIS-18*   HIS-19*   HIS-20*   HIS-21*   HIS-21*   HIS-22*   HIS-23*   HIS-24*   HIS-	25 to 36 40 to 43 44 to 51 52 to 67 68 to 71 68 to 71 68 to 101 124 to 131 137 to 160 163 to 170 173 to 180 181 to 212 22 to 24 38 39 72 73 74 to 77 78 79 80 to 83 84 85 102 103 to 106 107 108 to 111 112 to 115 116 to 119 123 120 122 132 to 135	V-1	37 38 40 41 42 44 to 51 52 54 55 to 62 63 64 66 to 69 70 71 to 78 80 to 83 84 86 to 89 90 to 93 94 to 97 98 to 101 102 104 to 107 108 to 115 116 to 119 120 to 123 125 to 128 129 to 132 133 to 136 137 138 to 141 142 to 145 146 to 149	
				HIS-24 HIS-25 HIS-26 HIS-27 HIS-28 HIS-29 HIS-30 HIS-31	161 162 171 172 213 214	V-37. V-38	159 to 162 164 to 167 168 169 to 176 177 to 180 182 to 185 186 to 193 194 to 201 202 to 205 206 to 209	
			Far	side <sup>b</sup>		<u></u>	<u> </u>	
[S-3* [S-9*		IIS-3 IIS-4 IIS-5 IIS-14		IIIS-2* IIIS-21.5_		VA-1 VA-2 VA-3 VA-4 VA-6 VA-7.1 VA-8 VA-10 VA-11.2 VA-12 VA-13 VA-14 VA-15 VA-16.1 VA-17.1 VA-18.1 VA-19 VA-20 VA-21 VA-21	13 to 20 21 22 22 24 25 26 28 29 30 31 31 32 39 43 65 79 85	

<sup>&#</sup>x27;Sites at which photographs were incompletely read out or secured in degraded form. See table 8.

<sup>&</sup>lt;sup>b</sup> Photographs centered on the far side were all taken obliquely except for the following near-vertical exposures: for mission 1, 28, 30, 35, 36, 37, 38, 39, 40, 115, 116, 136; for mission II, 33, 196.

<sup>\*</sup>Earth photographs. Lunar Orbiters took 3 exposures of Earth. Mission I exposures 102 and 117 yielded photographs showing a crescent Earth and an oblique view of the far side of the Moon just beyond its eastern limb (as seen from Earth). Mission V, exposure 27 (designated as site VA-9), yielded photographs showing a nearly full Earth.

Table 7.—Assignment of Exposures for Missions I, II, III, and V

Mi	ssion 1	1	sion II	1	#118810NS I, II	T	sion V
Site	Exposure number(s)	Site	Exposure number(s)	Site	Exposure number(s)	Site	Exposure number(s)
IS-1	5 to 24 25 to 27 28 29 30 31, 32 33, 34 35 to 40 41 42 43 44 45 46 to 49 50, 51 52 to 67 68 88 84 85 to 100 101 102 113, 114 115, 116 118 to 133 134, 135 136 137 138 139, 140 141 to 148 149 150 151 152 153 to 156 157 to 172	IIP-1 IIS-1 IIS-2 IIS-3 IIP-2 IIP-3 IIP-5 IIP-6 IIS-6 IIS-7 IIS-8 IIS-10 IIP-8 IIS-11 IIP-9 IIP-10 IIP-10 IIP-11 IIP-12 IIP-13 IIS-14 IIS-15 IIS-16 IIS-17	5 to 20 21 to 24 25 to 32 33 34 35 to 42 43 to 58 59 to 66 67 to 74 75 76 to 91 92 93 94 95 96 to 111 112 113-136 137 138 to 145 146 to 161 162 163 to 178 179 to 194	IIIP-1	5 to 20 21 to 24 25 to 36 37 38 39 40 to 43 44 to 51 52 to 67 68 to 71 72 73 74 to 77 78 79 80 to 83 84 85 86 to 101 102 103 to 106 107 108 to 111 112 to 115 116 to 119 120 121 123 124 to 131 132 to 135 136 137 to 160 161 162 163 to 170 171 172 173 to 180 181 to 212 213 214 215	VA-1	5 to 12 13 to 20 21 22 23 24 25 26 27 28 29 30 31 31 32 33 to 36 37 38 39 40 41 42 43 44 to 51 52 53 54 55 to 62 63 64 65 66 to 69 70 71 to 78 79 80 to 83 84 85 86 to 89 90 to 93 94 to 97 98 to 101 102 103 104 to 107 108 to 115 116 to 119 120 to 123 124 125 to 128 129 to 132 133 to 136 137 138 to 141 142 to 145 146 to 149 150 to 157 168 169 to 157 168 169 to 167 168 169 to 176 177 to 180 181 182 to 185 186 to 193 194 to 201 1902 to 205 206 to 209
						V-49	202 to 205

<sup>\*</sup>Film-handling considerations required that this frame be advanced through the cameras without being exposed.

<sup>&</sup>lt;sup>b</sup> An exposure taken for diagnostic test purposes. The medium-resolution frame was unexposed; the high-resolution frame was smeared during exposure.

Table 8.—Missions I, II, III, and V Sites for Which Photographs Were Incompletely Read Out or Degradeda

1	[	Photo	rank b	<u> </u>			o rank b
Site	Exposure number	Medium- resolution frame	High- resolution frame	Site	Exposure number	Medium- resolution frame	High- resolution frame
	Mi	ssion I			Mission 1	II—Continued	l
S-2	25	A100	C100	IIIP-2	25	A100	NRO
	26, 27	A100	A100		26	NRO	A22
S-3	28	A100	C100	1	27	NRO	A1
	30	A100	A100	i l	28	NRO	NRO
	35	A100	C100		29	NRO	NRO
	36	A100	A100		30	NRO	NRO
	37	A100	C100	li l	31	A100	NRO
	38	A100	A100		32	NRO	A19
ļ	39	A100	C100		33	A100	A9
	40	A100	B100	1	34	NRO	A5
S-4	29	A100	A100		35	NRO	A22
~	33, 34	A100	C100	li l	36	NRO	NRO
S-5	31	A100	A100	IIIP-3	40	NRO	A65
	32, 34	A100	C100		41	NRO	NRO
S-6	41	A100	A100		42	NRO	A75
	50, 51	A100	C100	1	43	NRO	NRO
IS-7	42	A100	A100	IIIP-4	44	NRO	A75
	46 to 49	A100	C100	1	45	NRO	NRO
S-9	102, 115	A100	A100	I	46	NRO	A89
	116	A100	C100		47	NRO	NRO
	117, 136	A100	A100		48	NRO	NRO
	·				49	NRO	NRO
				1	50	A03	NRO
	Mi	ssion II		II.	51	NRO	NRO
			1	IIIP-5	52	NRO	A93
	_		NDO		53	AO3	NRO
IIP-1	5	A100	NRO	il	54	NRO	A96
	6	NRO	A27		55	NRO	NRO
	7	A100	A15		56	NRO	NRO
	8	NRO	A36		57	NRO	NRO
	9	A100	A23		58	A83	NRO
	10	NRO	A27		59	NRO	NRO
	11	A100	A27		60	A100	A71
	12	A100	NRO		61	NRO	NRO
	13 to 20	A100	A100		62	NRO	A74
IIP-11	163 to 167	A100	A100		63	A.100	NRO
	168	A50	A100		64	NRO	A02
	169	A100	A89		65	NRO	A81
	170 to 178	A100	A100		66	A100	NRO
	l			-	67	NRO	A06
	Mi	ssion III		IIIP-6	68	A100	A84
	212.1	331011 111			69	NRO	A18
	Γ' ''-			7	70	A52	A75
IIIP-1	5	A100	NRO		71	NRO	NRO
	6	NRO	A31	IIIS-1	21	NRO	A11
	7	NRO	NRO		22	NRO	NRO
	8	NRO	NRO		23	NRO	NRO
	9	A100	NRO		24	NRO	NRO
	10	NRO	A74	IIIS-2	37	NRO	61°
	11	A100	NRO	IIIS-3	. 38	NRO	NRO
	12	NRO	A47	IIIS-4	. 39	NRO	NRO
	13	A100	A23	IIIS-5	72	NRO	A77
	14	NRO	A48	IIIS-6		A86	NRO
	15	A100	A12	IIIS-7		NRO	NRO
	16	NRO	A51	1	75	NRO	A83
	17	A100	A12	1	76		NRO
	18	NRO	A33		77	NRO	NRO
	19	A100	A13	L			
	20	NRO	A32		M	lission V	
				V F1	216	B100 <sup>d</sup>	A100
				V-51	1		A96
				V-51			No exp
			1	VA-16.1 VA-18			No exp

<sup>&</sup>lt;sup>a</sup> The photo rank is given for all photographs at each site, but only for those sites where one or more photographs was incompletely read out or degraded. All photographs not listed are ranked Al00 except for mission I high-resolution frames which are ranked C100.

degraded during exposure in the spacecraft and which is unusable for interpretation. Consideration is given only to those degradations associated with the operation of the photographs expression of the photographs contain blemishes associated with the craft's development process and others are over the craft's development process and others are over the craft's development process and others are over the craft's development process and others are overlap of the process of the photograph for interprets was not read out at all and No exp indicates the spacecraft film was unexposed. An experimental zero-phase photograph which was appreciably overexposed. It has questionable utility for interpretation of This photograph was incompletely developed in the spacecraft but is useful for interpretation.

Explanation of photo rank. An image quality grade of A, B, or C, based on subjective evaluation, is assigned to each photo graph and represents the state of the original film as secured from the spacecraft. This letter is followed by a number expressing the percent of the frame that was read out. Letter grades are:
A, a photograph free of image degradation: B, a photograph slightly degraded during exposure in the spacecraft, but which is usable for interpretation; and C, a photograph which was severely

Table 9.—Selenographic Distribution of Mission IV Exposures

	Latitude							Exposure	Exposure number for pass	1888 1						
Designation	n spacecraft nadir	34	33	32	31	30	53	28	27	26	25	24	23	22	21	20
Band G	33° N		191, 192		177		165							1001		
Band N	72° N		190		176		164		+	146, 147	140		128	179	116	!
Band D	45° N		189	183	175		163	158			139	134	127	122	115	-110 -110
Band C	14° N	961,	188	182	174	169	162	157	151	145	138	133	126	121	114	109
Band B	14° S	195	187	181	173		161	156	150	144	137	132	125	120	113	108
Band A	42° S	194	186	180	172		160	155	149	143	136	131	124	119	112	107
Band S	72° S	193	1	179	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	166	1	154	148	142		130	1	118	1	106
Band H	33° S		184, 185	178			1				1	1				
Approximate perilune longitude of spacecraft	: perilune if	92° W	85° W	M .6L	72° W	M .99	59° W	53° W	46° W	40° W	33° W	27° W	20° W	14° W	W	1° W
Approximate apolune longitude of spacecraft	apolune f		95° E	101° E	108° E		121° E			140° E				166° E		

						Ø	xposure num	Exposure number for passContinued	Continue	_				
Designation spacecraft nadir	19	18	17	16	15	14	13	12	11	10	6	8	7	9
Band G		l i												
		1 66 <sub>q</sub>				- 475				221				9.5
Band N.	104		92		.80	1	89 <sub>q</sub>	1	,26		542		,30 ,30	ì
Band D	103	86	16	984	624	₽74	.9 <sub>e</sub>	,e2	22	₽ <b>4</b> 8	•41 •41	984	229	21 to 24°
Band C	102	164	064	282	824	b73	994	▶61	224	47	₽40	\$35	<sub>2</sub> 28	17 to 20°
Band B.	101	96	684	84	LLa	272	<b>,</b> 62	9.	<b>22</b>	₽46	684	<b>34</b>	<sup>6</sup> 27	13 to 16 <sup>b</sup>
Band A	100	95	884	83	94	₽71	₽94	*59	,25	\$45	8£4	233	<sup>2</sup> 26	9 to 12b
Band S.	1 1 1	P94		, 82 , 82		024		, 58 128	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	P44	1	- 32		5 to 8
		111111	2.1			* I .		3.1		9.1				
Approximate perilune longitude of spacecraft	2 2	ਤ ਤ ਤ	3 .6I	32° 52°	35° E	च 88	45° E	ਤਾ <sub>•</sub> ਤਿ	58° E	64° E	71° E	77° E	84° E	ਬ .06
Approximate apolune longitude of		168° W				142° W				116° W				₩ °06
spacecraft														

Exposures in lattude bands S, A, B, C, D, and N were taken near perliune with morning illumination. Photographs in bands F, G, and H were taken near apolune with evening illumination.

 $^{\rm b}\,\rm Exposures$  for which photographs were incompletely read out or secured in degraded form. See table 11.

 ${\tt TABLE~10.} {\color{red} -} Assignment~of~{\it Mission~IV~Exposures}$ 

	Site b	Exposure number(s)	Site b	Exposure number(s)	Site <sup>b</sup>	Exposure number(s)	Site <sup>b</sup>	Exposure number(s)	Site <sup>b</sup>	Exposure number(s)	Site b	Exposure number(s)
		*1.0	T31 11 4	E9	TV 160	99	TV 91 A	119	TV_26A	142	TV-31A	172
												173
												174
												175
			- ,									176
			1V-11N		14-100		14-2114			1		177
	IV-6F	25		-91		0,		11,		1		
IV-7B	IV-7A	26	IV-12S	58	IV-17A	88	IV-22S	118				
IV-7C   28		27	IV-12A	59	IV-17B	89	IV-22A					
IV-7D   29				60	IV-17C	90	IV-22B					
IV-1N   30					IV-17D	91	IV-22C	121	IV-27D			181
IV-8S   31				62	IV-17N	92	IV-22D	122	IV-27N	152		182
	.,			*63	1	*93	IV-22F	123		*153	IV-32D	183
					*** *00	0.4	117 00 4	104	137 900	154	IV-33H	184
												_
IV-8D												
IV-9B												
IV-9A	IV-8D		IV-13N				1V-23N		17-200			1
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$		37*		-69	IV-18F	99		129		100		
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	IV-9A	38	IV-14S	70	IV-19A	100				i		191
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$		39	IV-14A	71	IV-19B	101				ţ		192
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$			IV-14B	72	IV-19C	102	ì					_
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$			IV-14C	73	IV-19D	103	IV-24S			1		193
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$		42	IV-14D	74	IV-19N	104	IV-24A	131	IV-29B	161		194
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$		43*	IV-14F	75		*105	IV-24B	132	IV-29C	162		
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$							IV-24C	133	IV-29D	163	IV-34C	196
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$							IV-24D	134	IV-29N	164		
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$								135	IV-29G	165		
IV-10D							l		*** 000	100		
111   IV-25C 138   IV-30B 168   IV-10F 150   IV-25D 140   IV-30D 170   IV-10F 170   IV-25N 140   IV-30D 170   I								1				
IV-25D 139   IV-30C 169   IV-10F 170   IV-25N 140   IV-30D 170   IV-30D	IV-10D		IV-15N		IV-20D							
IV-10F 51   IV-25N 140   IV-30D 170				*81		*111						
17-107						ł						
	IV-10F	51			ŀ	1	1V-25N		114-30D			
				1				141		171		

<sup>\*</sup> Film-handling considerations required that this frame be advanced through the cameras without being exposed.

<sup>&</sup>lt;sup>b</sup> Sites designated by: IV denotes mission IV; arabic numeral, pass number; and letters indicate the latitude band of photography. See table 9.

Table 11.—Mission IV Sites for Which Photographs Were Incompletely Read Out or Degraded\*

Site	Exposure number	Photo rank b				Photo rank b	
		Medium- resolution frame	High- resolution frame	Site	Exposure number	Medium- resolution frame	High- resolution frame
IV-6S	5	NRO	A100	IV-12S	58	B100	B100
	6	NRO	A93	IV-12A	59	B100	B100
	7	NRO	NRO	IV-12B	60	B100	B100
	8	A100	A59	IV-12C	61	C68	B100
				IV-12D	62	C100	B100
IV-6A	9	A100	A100				
	10	A100	A93	IV-13A	64	C100	B100
	11	A100	A100	IV-13B	65	C100	B100
	12	A100	A99	IV-13C	66	C100	B100
IV-6B	13	No exp	No exp	IV-13D	67	C39	B100
	14	No exp	No exp	IV-13N	68	C100	B100
	15	No exp	No exp	IV-14S	70	C100	A100
	16	No exp	No exp	IV-14A	71	B100	B100
				IV-14B	72	B100	B100
IV-6C	17	A32	A100	IV-14C	73	B100	B100
	18	NRO	A100	IV-14D	74	B100	A100
	19	C36	NRO	IV-14F	75	C100	No exp
	20	NRO	C54	1			1
IV-6D	21	B100	NRO	IV-15A	76	B100	A100
	22	NRO	C34	IV-15B	77	B100	A100
	23	B100	C01	IV-15C	78	B100	A100
	24	NRO	A09	IV-15D	79	B100	A100
			1	IV-15N	80	C100	B100
IV-6F	25	C100	No exp	IV-16S	82	B100	A100
V-7A	26	No exp	No exp	IV-16C	85	B50	B100
IV-7B	27	C100	B100	IV-16D	86	A100	B100
V-7C	28	C89	C100	1		1	
V-7D	29	C100	C100	IV-17A	88	B100	A100
V-7N	30	C100	C04	IV-17B	89	B100	A100
				IV-17C	90	B100	A100
IV-8S	32	C100	C100	IV-18S	94	B100	A100
V-8A	33	C100	C100	IV-18C	97	A89	A100
V-8B	34	C100	C100	IV-18F	99	B100	No exp
V-8C	35	C71	C100				110 CXP
V-8D	36	C21	C100	IV-20D	110	B100	A100
[V-9A	38	B100	B100	IV-22F	123	B100	No exp
IV-9B	39	C100	B100		120	D100	110 CKP
V-9C	40	C100	C100	IV-26F	146	A100	No exp <sup>c</sup>
V-9D	41	C64	C100	IV-26F	147	A100	No exp°
V-9N	42	C100	C100	IV-27N	152	A100	A98
V-10S							
	44	B100	B100	IV-31N	176	No exp	A100
V-10A V-10B	45	C39	B05	IV-31G	177	B100	B100
V-10B V-10C	46	C100	B100	IV-32H	178	B100	A100
	47	C11	C100		119	D100	VIOO
V-10D	48	C100	C100	IV-33H	184	B100	A100
V-10F	51	B100	NRO	IV-33H	185	B100	A100
V-11A	52	A100	B74	IV-34B		MDO	1
V-11B	53	B100	B100		195	NRO	A100
V-11C	54	C68	B100	IV-34C	196	NRO	A38
V-11D	55	C100	B100				
V-11N	56	C100	B100				1
- 7	50		2100				í

<sup>&</sup>lt;sup>a</sup> The photo rank is given for all photographs at each site, but only for those sites where one or more photographs were incompletely read out and/or secured in degraded form. All photographs not listed are ranked A100.

for interpretation. Consideration is given only to those degradations associated with the operation of the photographic system. Many photographs contain blemishes associated with the spacecraft's development process and others are overexposed to varying degrees. Generally, neither of these blemishes seriously affect the usefulness of the photograph for interpretation and are not considered here. NRO indicates the photograph was not read out at all and No exp indicates the spacecraft film was unexposed.

\*Mission IV exposures taken at apolune for coverage of the far side are 25, 51, 75, 99, 123, 146, and 147. For each exposure, the high-resolution coverage is situated on the unilluminated side of the evening terminator except for small portions of photographs IV-99H, IV-123H, IV-146H, and IV-147H.

b Explanation of photo rank. An image quality grade of A, B, or C, based on subjective evaluation, is assigned to each photograph and represents the state of the original film as secured from the spacecraft. This letter is followed by a number expressing the percent of the frame that was read out. Letter grades are: A, a photograph free of image degradation; B, a photograph slightly degraded during exposure in the spacecraft, but which is usable for interpretation; and C, a photograph which was severely degraded during exposure in the spacecraft and which is unusable

 ${\tt TABLE~12.--Photographs~Processed~for~Emphasis~of~Detail~in~Highlights~and~Shadows}$ 

Area of interest	Photographs for which SP highlights are available	Photographs for which SP-1 lowlights are available		
Structures on bright peak and wall of Petavius	V-33M, V-34M, V-36M			
Details of areas in and around Censorinus	V-63M, V-63H	V-63M, V-63H		
Area of Littrow Rilles	V-66M, V-66H			
Bright crater in Sulpicius Gallus Region	V-90H			
Bright walls of Hyginus Rille craters	V-96H, V-97H			
Slopes in vicinity of Alpine Valley	V-102H	V-102H		
Slopes in vicinity of Hadley Rille	V-104M, V-105M, V-106M, V-107M	V-104M, V-105M, V-106M, V-107M		
Bright walls of crater Alphonsus	V-116H, V-117H, V-118H, V-119H			
Walls of crater near Rima Bode II	V-122H			
Bright walls of crater Tycho	V-125H			
Steep slopes in Rima Plato II Region	V-130M, V-131M, V-132M, V-129H, V-130H, V-131H, V-132H	V-130M, V-131M, V-132M, V-129H, V-130H, V-131H, V-132H		
Walls of crater Fra Mauro and other slopes	V-138H	V-138H		
Bright walls of crater Copernicus	V-152M			
Scarps of the Imbrian flows	V-160H, V-161H	V-160H, V-161H		
Steep slopes in vicinity of Tobias Mayer	V-164M, V-165M, V-166M, V-167M	V-164M, V-165M, V-166M, V-167M		
Slopes of Gassendi and adjacent territory	V-177M, V-178M, V-179M, V-180M, V-178H, V-179H	V-177M, V-178M, V-179M, V-180M, V-178H, V-179H		
Bright slopes of Jura domes and the terra ridges	V-182M, V-184H, V-185H			
Slopes of rilles and craters in the Harbinger Mountains $_{--}$	V-187H, V-188H, V-189H, V-190H, V-191H, V-192H	V-187H, V-188H, V-189H, V-190H, V-191H, V-192H		
Walls of crater Aristarchus	V-198H, V199H, V-200H	V-198H, V-199H, V-200H		
Sinuous rille and bright walls of Schröter's Valley	V-204H	V-204H		

Table 13.—Some Characteristics of Lunar Orbiter Photographs

		Photo characteristics (average values)							
Mission	Typical spacecraft altitude, km	Medium-resolution frames			High-resolution frames				
		Photo scale (GRE scale)	Ground resolution, m	Framelet width,* km	Photo scale (GRE scale)	Ground resolution, m	Framelet width,* km	Reassembly code <sup>c</sup>	
			Photographs o	f near side					
Mission I:									
Exposures 5 to 27, 29,									
31 to 34, 41, and 42	240	1:420 000	40	7.60	1:55 000	5 to 10	1.0	N=3	
Other exposures	55	1:96 000	10	1.75	1:12 500	≈40	0.23	N=3	
Mission II	50	1:87 000	10	1.60	1:11 400	1	0.21	N=3	
Mission III	55	1:96 000	10	1.75	1:12 500	1	0.21	N=3	
Mission IV:	80	1.30 000	10	1.10	1.12 000	1	0.28	N = 3	
Perilune photos:									
Equatorial regions	2710	1:4 700 000	500	86	1:620 000	60	11	N = 3	
Temperate regions	2940	1:5 100 000	500	93	1:670 000	64	11 12	N=3 $N=3$	
Polar regions	3520	1:6 100 000	600	111	1:800 000	76	15	$N \equiv 3$ $N \equiv 3$	
Apolune photographs	5650	1:9 800 000	1000	180	1:1 300 000	120	24	N=3 N=1	
Mission V:	0000	2.0 000 000	1000	-00	1.1 000 000	120		1	
Extreme value	97	1:169 000	20	3.1	1:22 000	2	0.4	E=1	
Extreme value	243	1:423 000	40	7.60	1:55 000	5	1.0	E=1	
			Photographs of	of far side					
Mission I	1375	1:2 400 000	240	43	1:310 000	30	5.7	N = 1	
Mission II	1500	1:2 600 000	260	48	1:340 000	30	6.2	N = 1	
Mission IIIb	1463	1:2 500 000	260	46	1:330 000	30	6.1	N=1	
Mission IV: Apolune photo-	1400	1.2 000 000	200	40	1.000 000	30 1	0.1	24 — 1	
graphs <sup>b</sup>	6150	1:10 700 000	100	195		Not appli	ionhla		
Mission V:	0100	1.10 100 000	100	130		1400 appli	icable		
Exposures 5 to 30:	1			1					
Extreme value	2548	1:4 400 000	450	81	1:580 000	55 1	11	N = 3	
Extreme value	5758	1:10 000 000	1000	183	1:1 300 000	125	24	N=3 $N=3$	
Other exposures:	5.00	1.10 000 000	2000	-00	1.1 000 000	120	24	** - 0	
Extreme value	1181	1:2 000 000	200	37	1:270 000	30	5	N = 3	
Extreme value	1396	1:2 400 000	240	44	1:320 000	30	6	N=3	

<sup>&</sup>lt;sup>a</sup> All Lunar Orbiter photographs are distinguished by faint parallel lines running widthwise. These lines are spaced at approximate 20-mm (0.75-inch) intervals on the 20- by 24-inch sections and provide a convenient rule for measuring the ground distances given.

unilluminated area; however, the values given provide a gross characterization of these photographs and are given for a comparison with the other listings and for completeness. Photographs without superscripts are vertical photographs.

b All photographs on this row were taken obliquely. The values given for photographic characteristics apply only to the nadir point which in most cases were located in an

c See text, page 6, for explanation of reassembly code.

Table 14.—Lunar Orbiter Photographs Available From the National Space Science Data Center

Photo rank	Mission I *	Mission II	Mission III	Mission IV	Mission V	Total
		Medium-reso	lution frames			
A100 or B100 A (<100) or B (<100)	206 0	207 1	152 5	122 3	211	898 9
Total frames	206	208	157	125	211	907
20-inch by 24-inch sections (subtotal)	206	208	157	125	211	907
		High-resolu	ition frames			
A100 or B100A (<100) or B (<100)	14 0	202 7	138 35	132 9	209	695 52
Total frames	14	209	173	141	210	747
20-inch by 24-inch sections (subtotal)	42	615	482	417	630	2186
		Medium-reso	olution frames			
C100	0 0 0	0 0 0 3	0 0 0 54	22 10 6	0 0 0	22 10 6
Total frames	0	3	54	47	1	105
20-inch by 24-inch sections (subtotal)	0	0	0	32	0	32
	-	High-resol	ution frames			
C100 C (<100) Not exposed Not read out	192 0 0 0	0 0 0 2	0 0 0 38	12 4 11 4	0 0 2 0	204 4 13 44
Total frames	192	2	38	31	2	268
20-inch by 24-inch sections (subtotal)		. 0	0	41	0	41
20-inch by 24-inch sections (subtotal)	248	823	639	615	841	3166

<sup>\*</sup>Copies of all photographs are available from the NSSDC as 20- by 24-inch sections with the exception of the smeared high-resolution frames of mission I. Copies of these photographs are, however, available as 9½-inch roll film or paper.

## References

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- BEELER, M.; AND MICHLOVITZ, K.: Lunar Orbiter Photographic Data. Data Users' Note, NSSDC 69-05, NASA Goddard Space Flight Center, June 1969.
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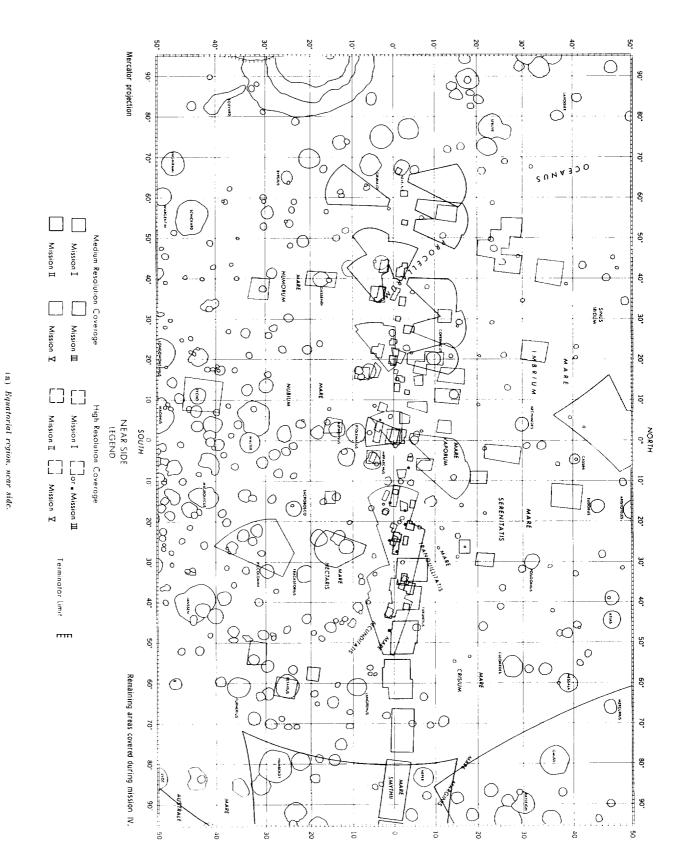


Figure 4.—Mission Index for missions I, II, III, and V.

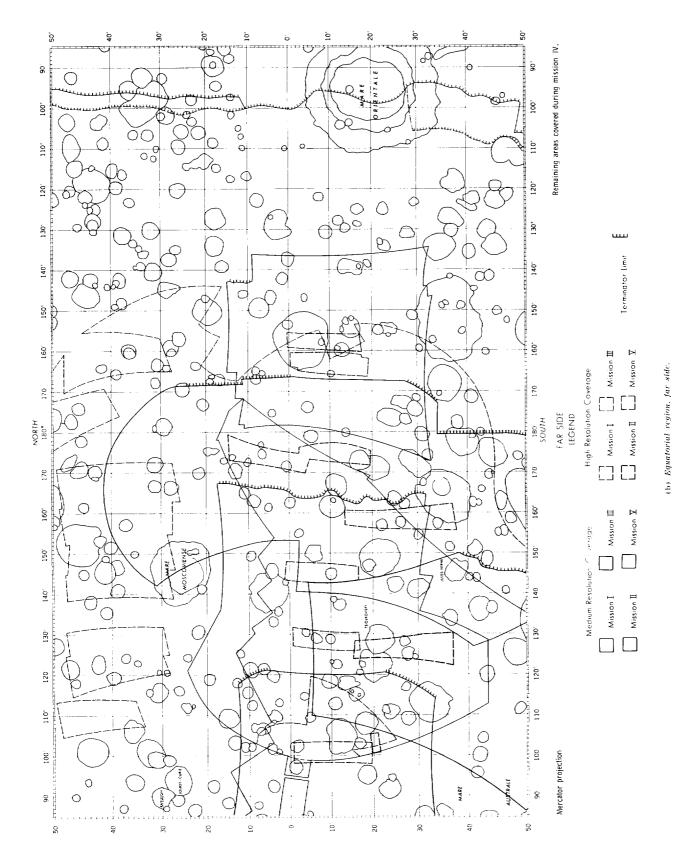
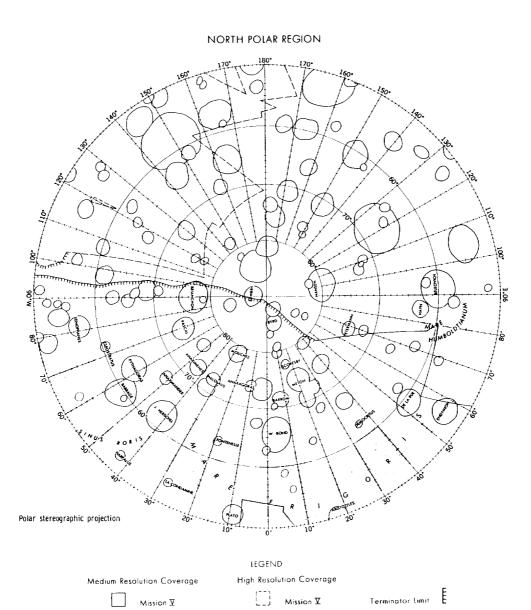


Figure 4.—Mission Index for missions I. II. III. and V.—Continued.

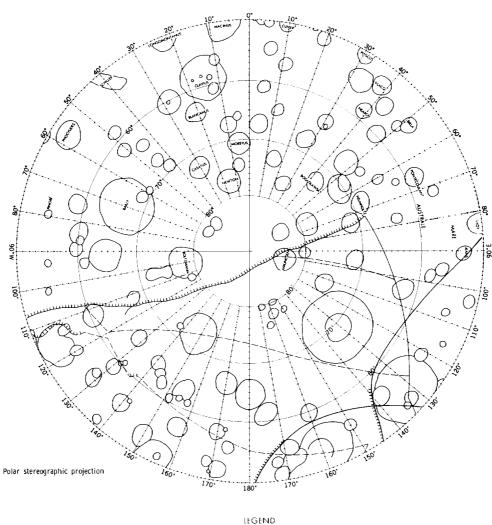


Remaining area covered by mission  ${\sf IV}$ .

(c) North polar region.

Figure 4.—Mission Index for missions 1, 11, 111, and V.—Continued.

# SOUTH POLAR REGION

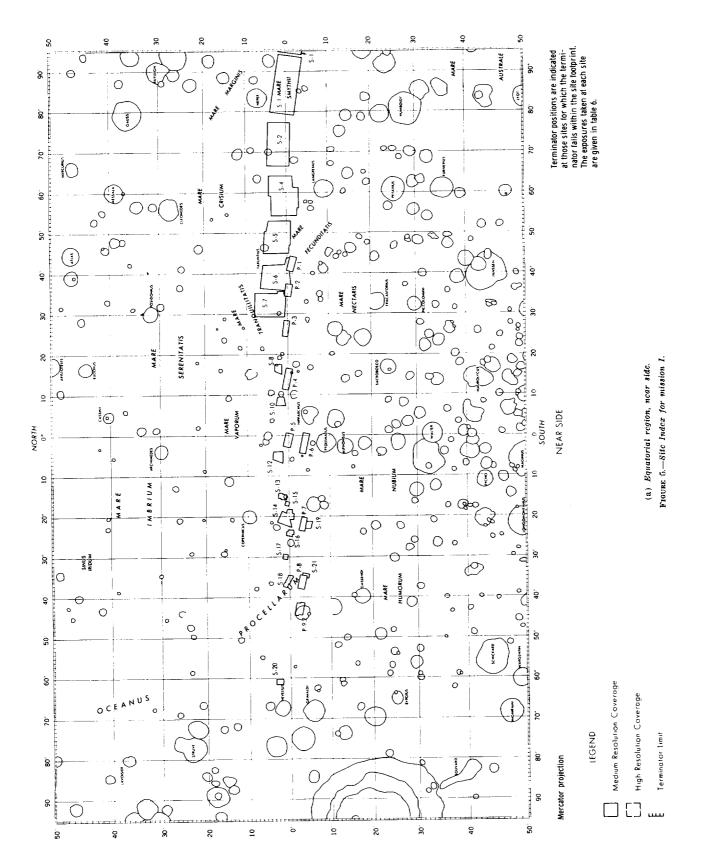


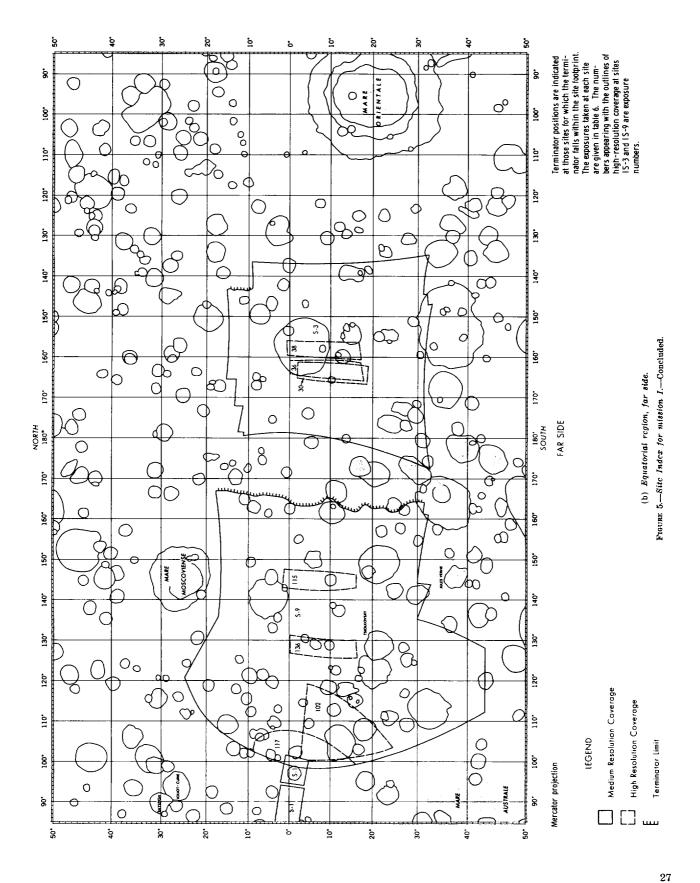
## 

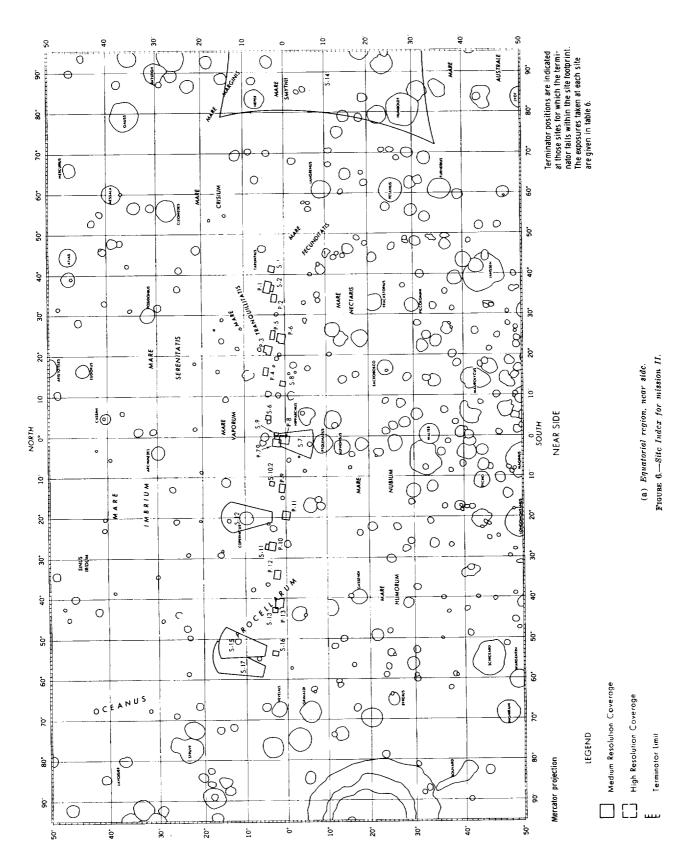
Remaining area covered by mission  $\ensuremath{\text{IV}}.$ 

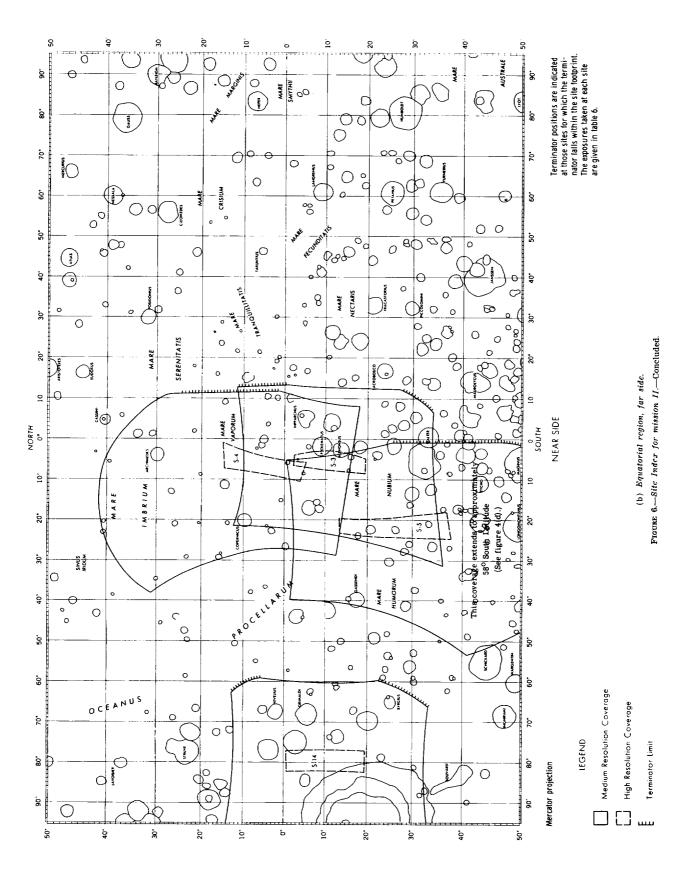
(d) South polar region.

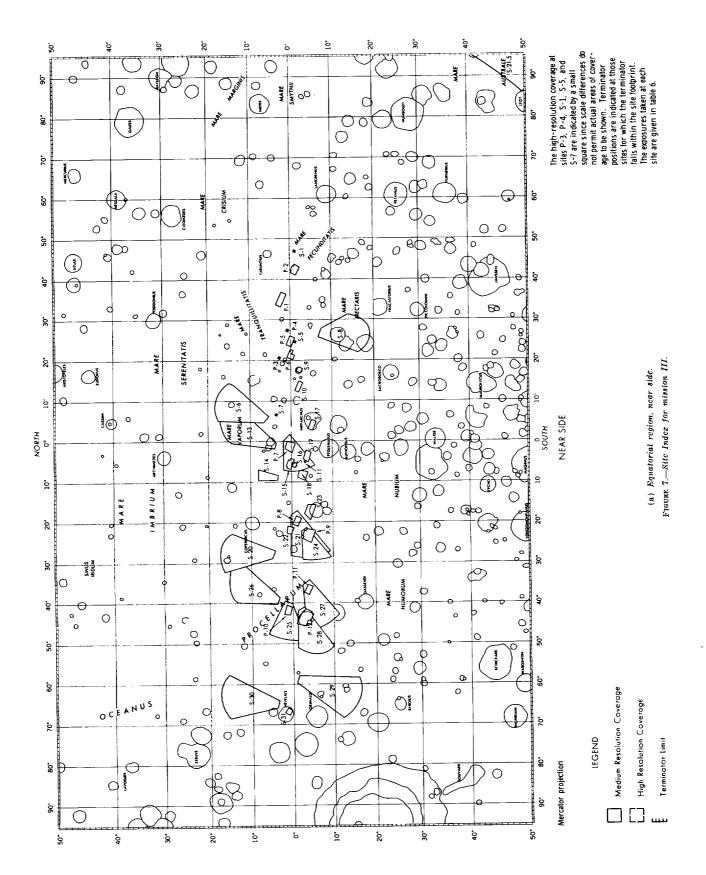
Figure 4.—Mission Index for missions 1, 11, 111, and V.—Concluded

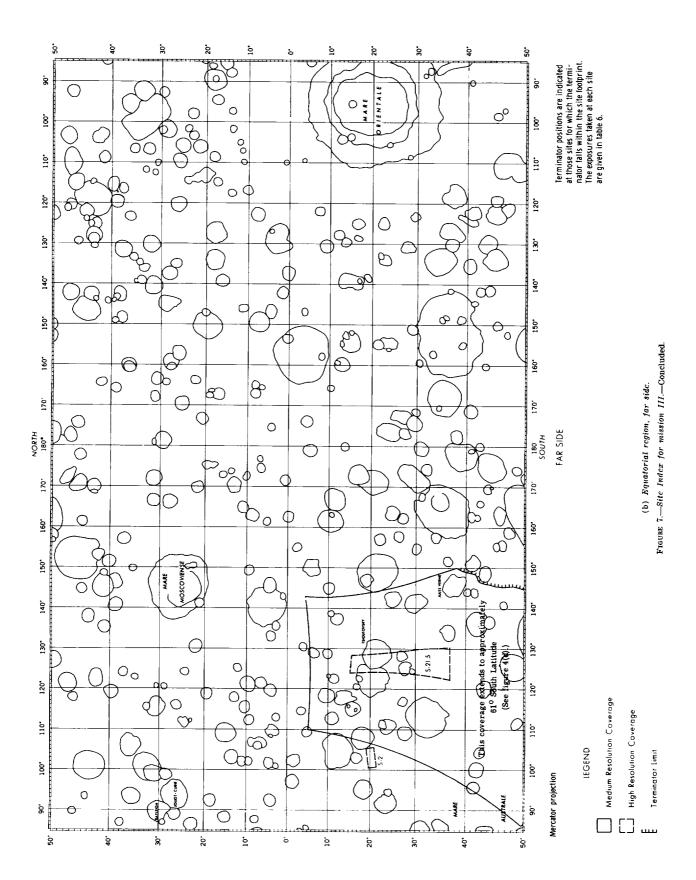


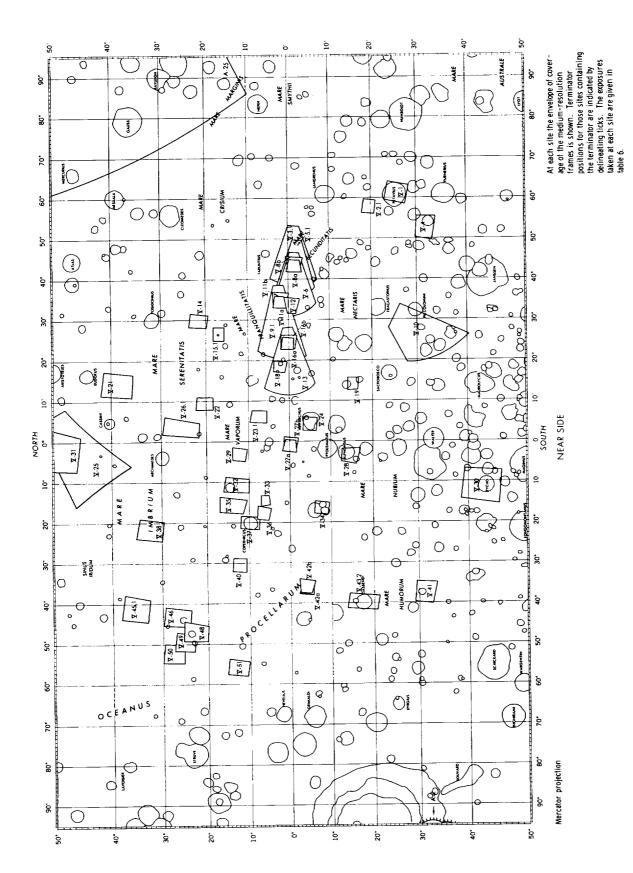




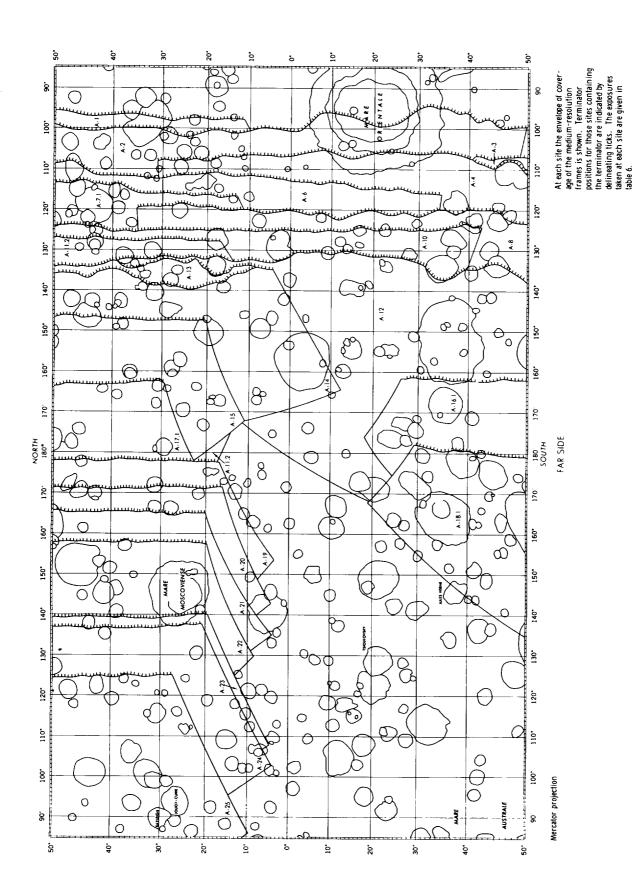




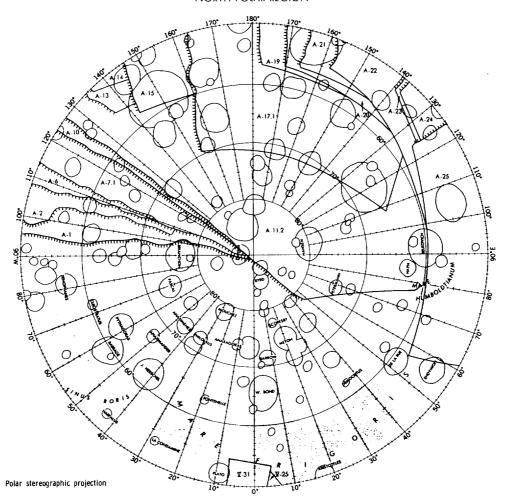




(a) Equatorial region, near side. Floure 8.—Site Index for mission V.



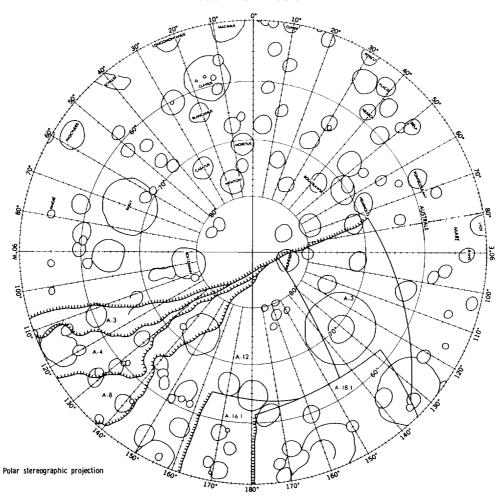
(b) Equatorial region, far side.
Figure 8.—Site Index for mission V.—Continued.



At each site the envelope of coverage of the medium-resolution frames is shown. Terminator positions for those sites containing the terminator are indicated by delineating ticks. The exposures taken at each site are given in table 6.

(c) North polar region.

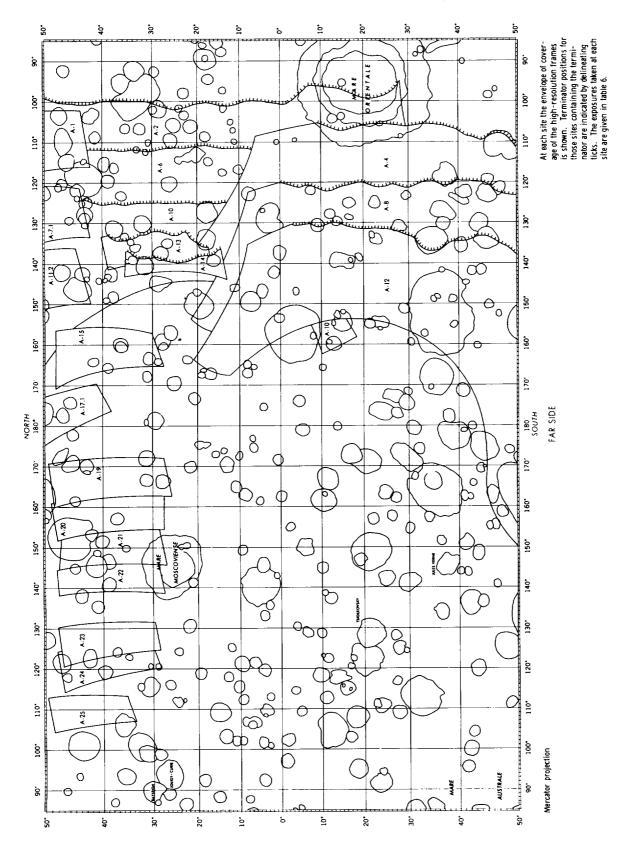
FIGURE 8.—Site Index for mission V.—Continued.



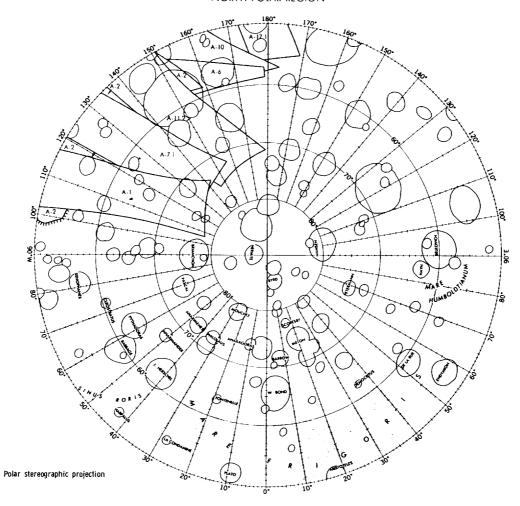
At each site the envelope of coverage of the medium-resolution frames is shown. Terminator positions for those sites containing the terminator are indicated by the terminator are indicated by the terminator sites. The exposures taken at each site are given in table 6.

(d) South polar region.

FIGURE 8.—Site Index for mission V.—Concluded.



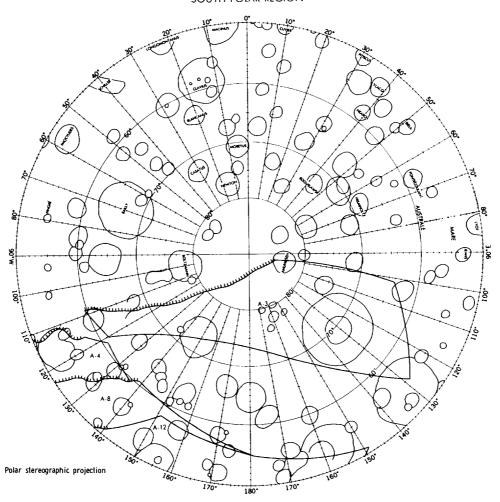
(a) Equatorial region, far side. Flaune 9.—Photographic Indexes for mission V high-resolution frames of the far side.



At each site the envelope of coverage of the high-resolution frames is shown. Terminator positions for those sites containing the terminator are indicated by delineating ticks. The exposures taken at each site are given in table 6.

(b) North polar region.

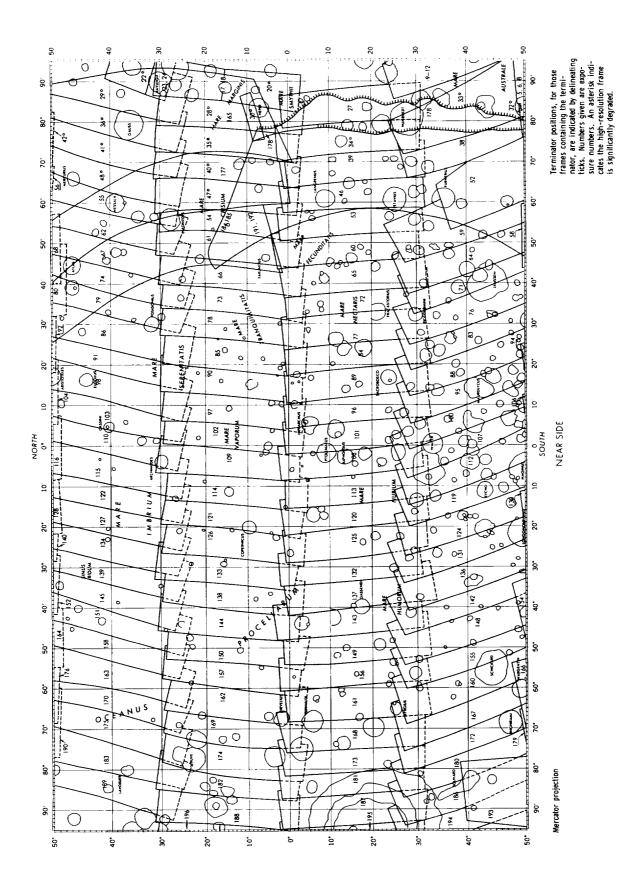
 $\textbf{Figure 9.} \\ \textbf{--Photographic Index} cs~for~mission~V~high-resolution~frames~of~the~far~side.\\ \textbf{--}Continued.$ 



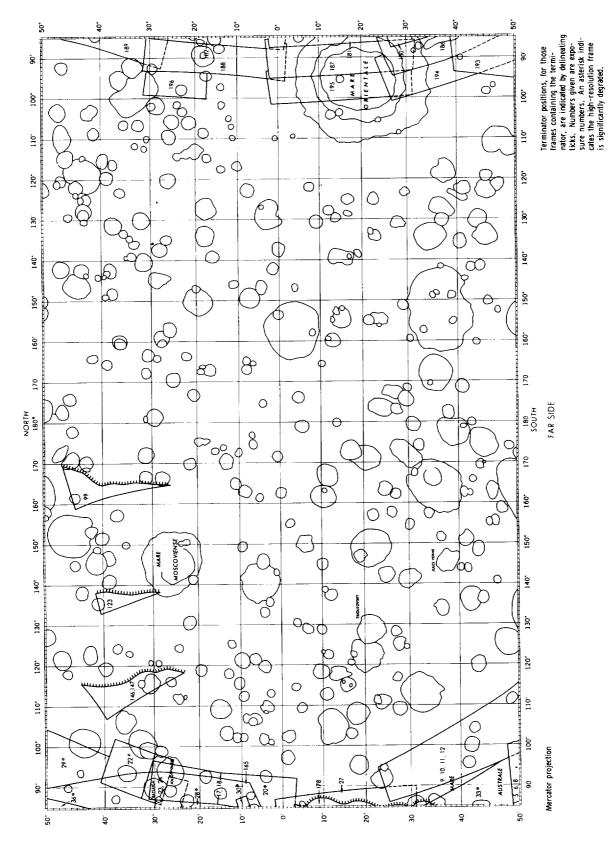
At each site the envelope of coverage of the high-resolution frames is shown. Terminator positions for those sites containing the terminator are indicated by delineating ticks. The exposures taken at each site are given in table 6.

(c) South polar region.

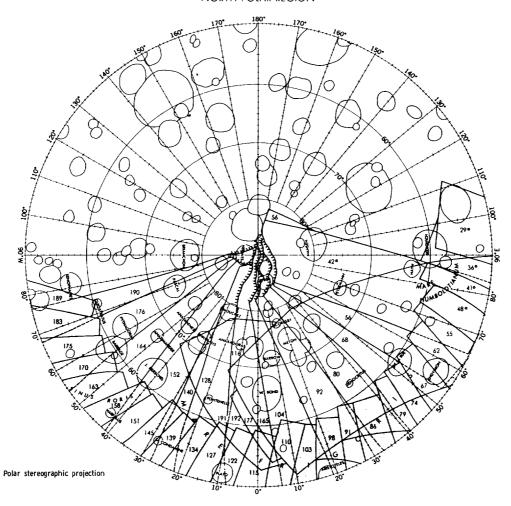
 $\textbf{Figure 9.} \\ \textbf{-Photographic Indexes for mission $V$ high-resolution frames of the far side.} \\ \textbf{-} \textbf{Concluded.}$ 



(a) Equatorial region, near side.
Figure 10.—Photographic Indexes for mission IV high-resolution frames.



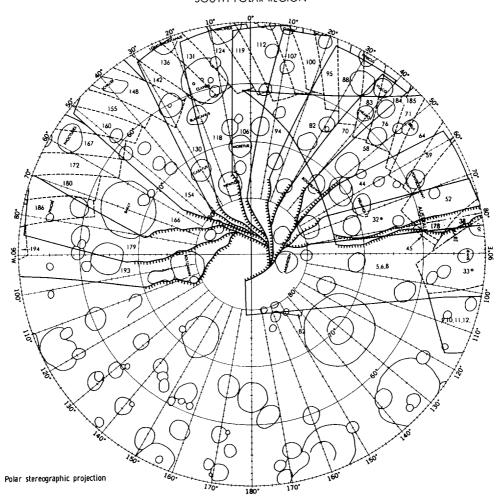
(b) Equatorial region, far side. Floure 10.—Photographic Indexes for mission II high-resolution frames.—Continued.



Terminator positions, for those frames containing the terminator, are indicated by delineating ticks. Numbers given are exposure numbers. An asterisk indicates the high-resolution frame is significantly degraded.

(c) North polar region.

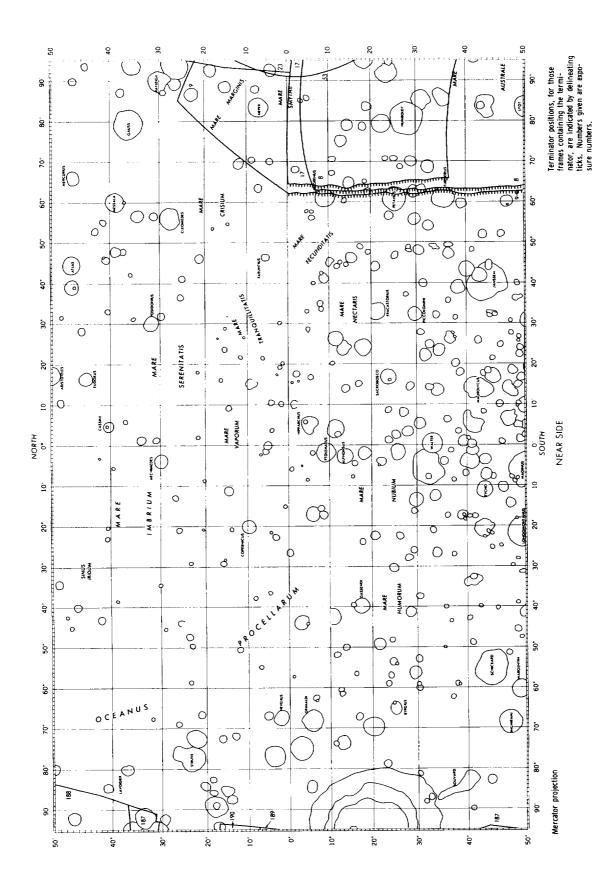
 $\textbf{Figure 10.--Photographic Indexes for mission } IV \ \textit{high-resolution frames.--} \textbf{Continued.}$ 



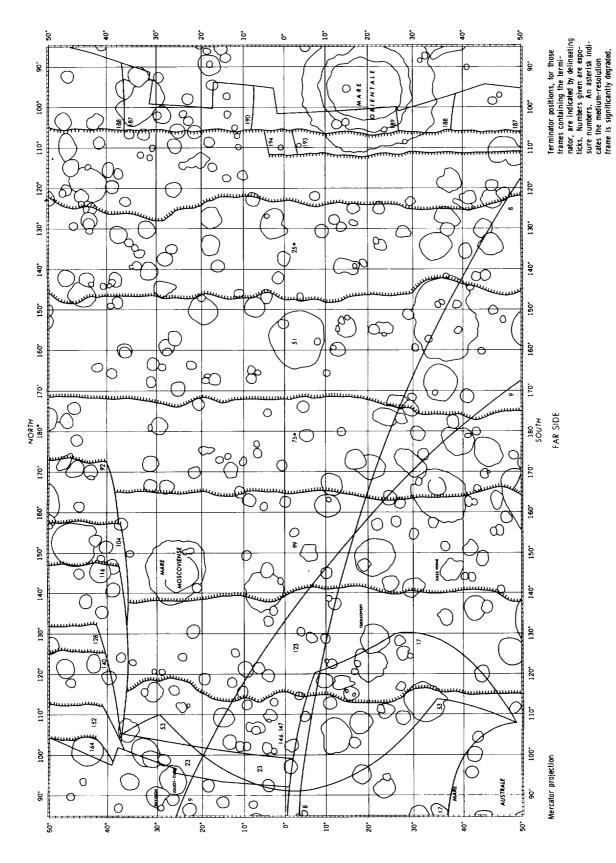
Terminator positions, for those frames containing the terminator, are indicated by delineating ticks. Numbers given are exposure numbers. An asterisk indicates the high-resolution frame is significantly degraded.

(d) South polar region.

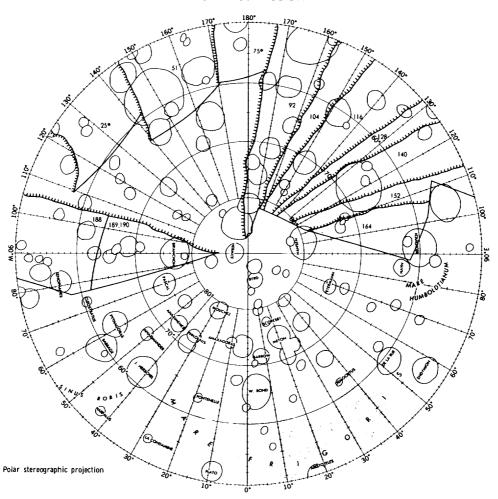
Figure 10.--Photographic Indexes for mission IV high-resolution frames.-- Concluded.



(a) Equatorial region, near side. Figure 11.—Photographic Indexes for selected mission IV medium-resolution frames of the far side.



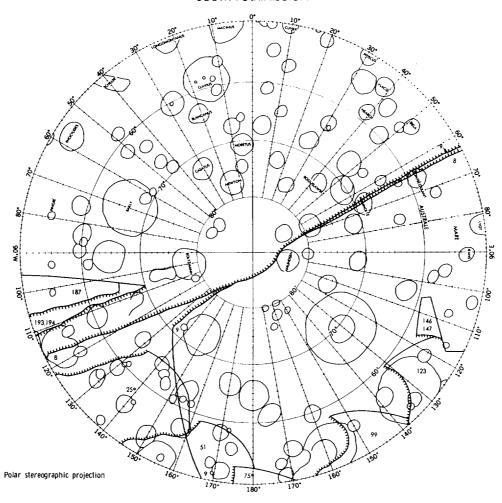
(b) Equatorial region, far side.
Froune 11.—Photographic Indexes for selected mission IV medium-resolution frames of the far side.—Continued.



Terminator positions, for those frames containing the terminator, are indicated by delineating ticks. Numbers given are exposure numbers. An asterisk indicates the medium-resolution frame is significantly degraded.

(e) North polar region.

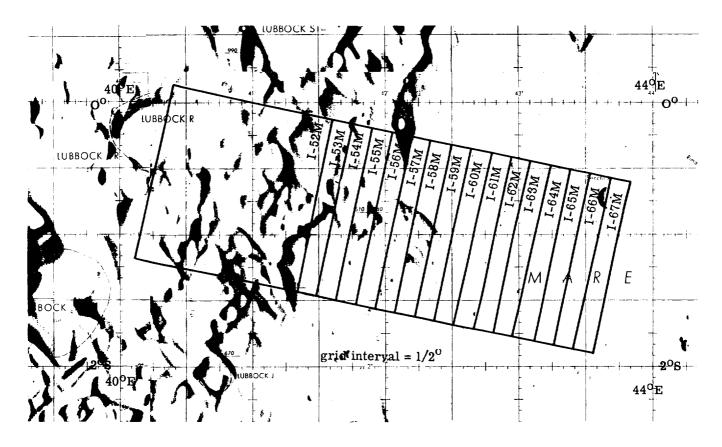
FIGURE 11.—Photographic Indexes for selected mission IV medium-resolution frames of the far side.—Continued.



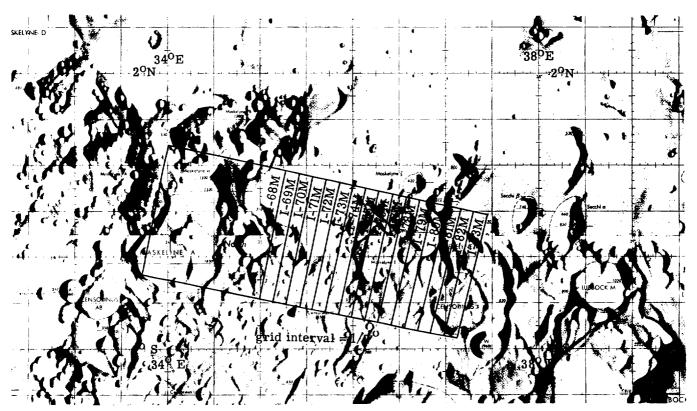
Terminator positions, for those frames containing the terminator, are indicated by delineating ticks. Numbers given are exposure numbers. An asterisk indicates the medium-resolution frame is significantly degraded.

(d) South polar region.

FIGURE 11.—Photographic Indexes for selected mission IV medium-resolution frames of the far side.—Concluded.



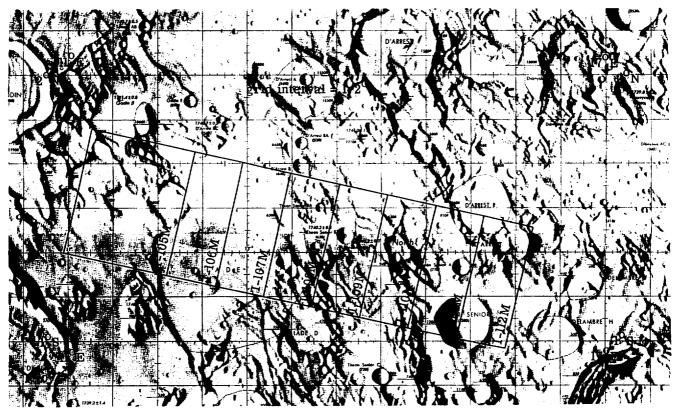
(a) Site IP-1.



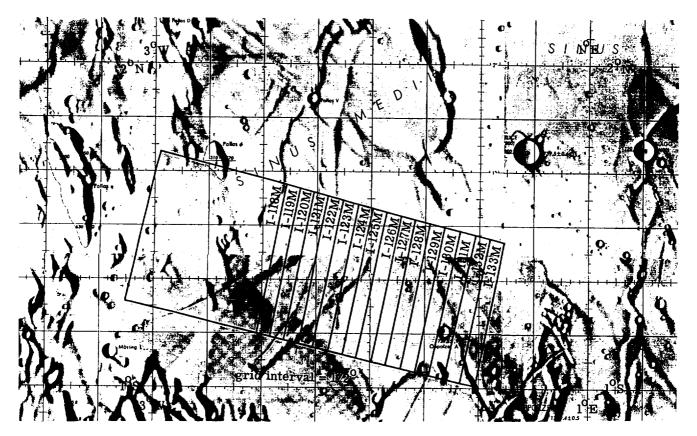
(b) Site IP-2.



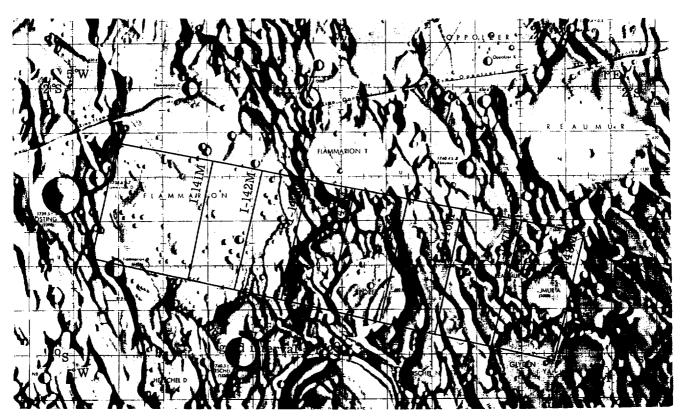
(c) Site IP-3.



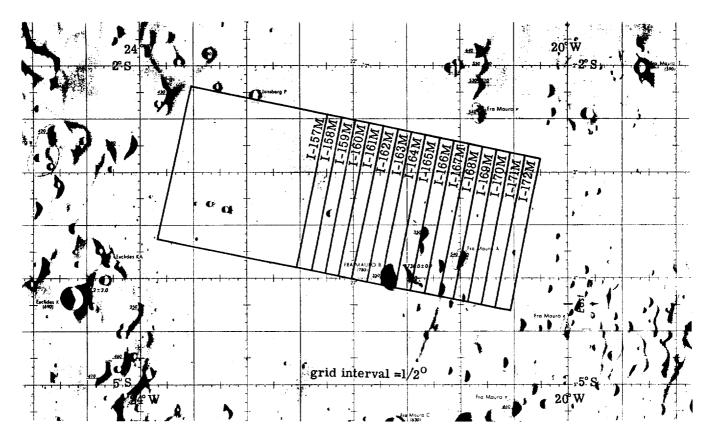
(d) Site IP-4.



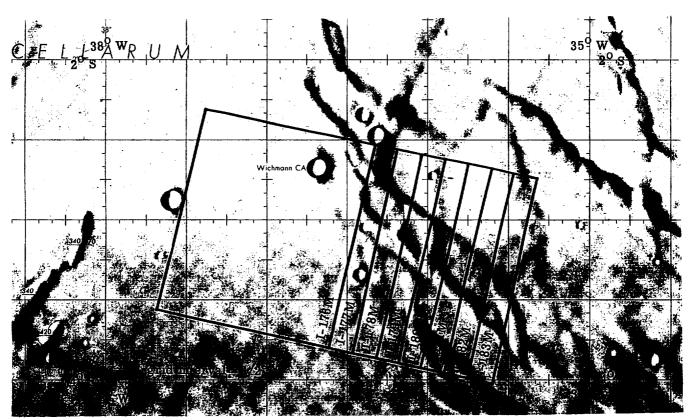
(e) Site IP-5.



(f) Site IP-6.







(h) Site IP-8.1.
Figure 12.—Photographic Indexes to mission I near-side sites.—Continued.

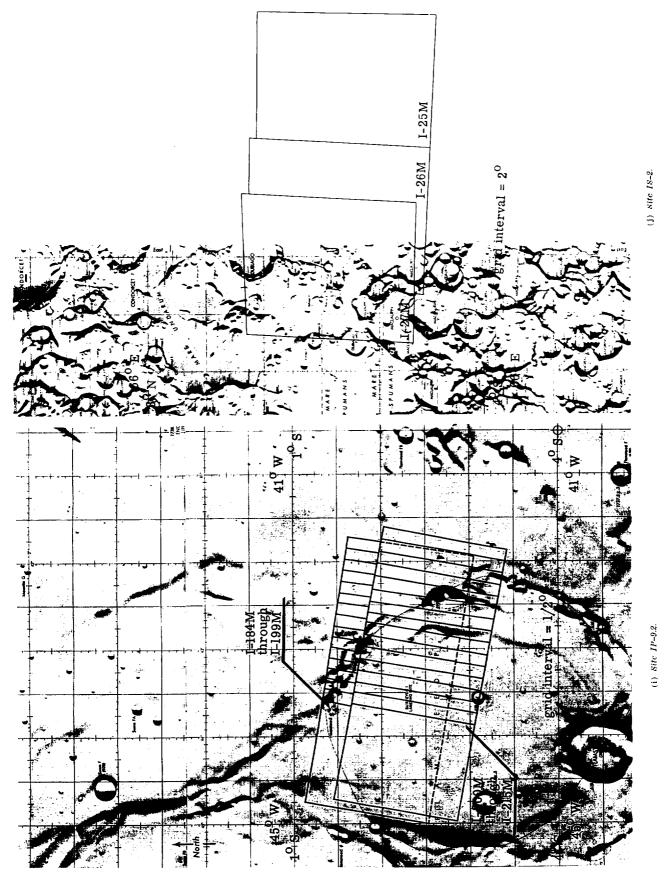
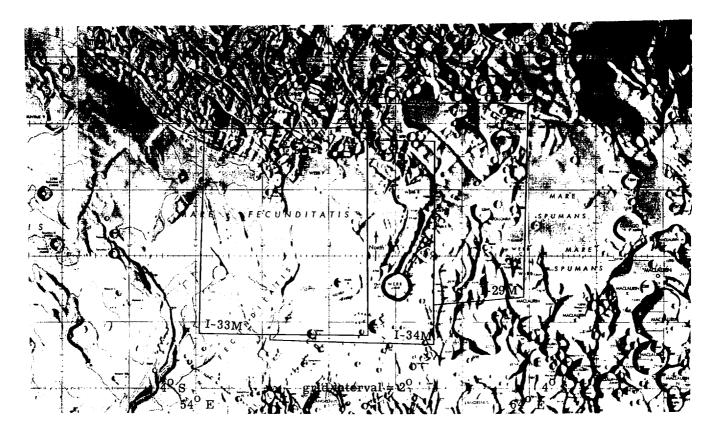
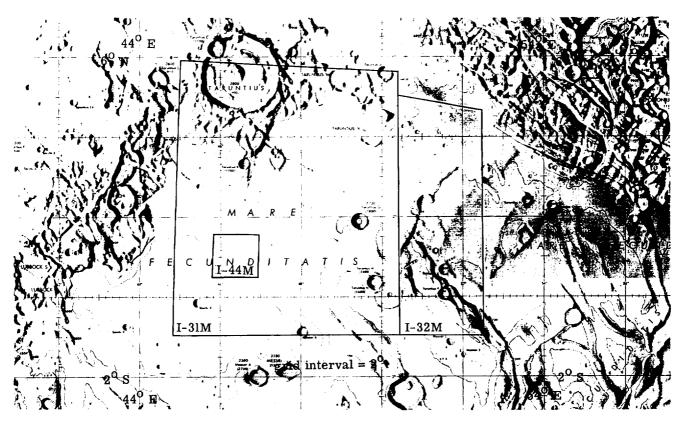


Figure 12.—Photographic Indexes to mission I near-side sites.—Continued.



(k) Site IS-4.



(1) Site IS-5.
Figure 12.—Photographic Indexes to mission I near-side sites.—Continued.

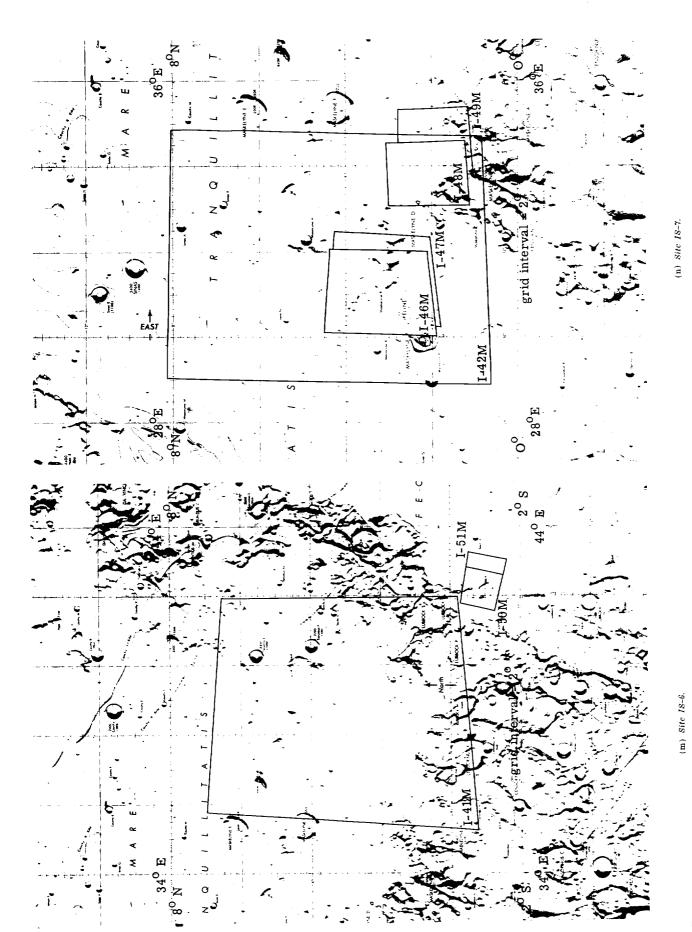
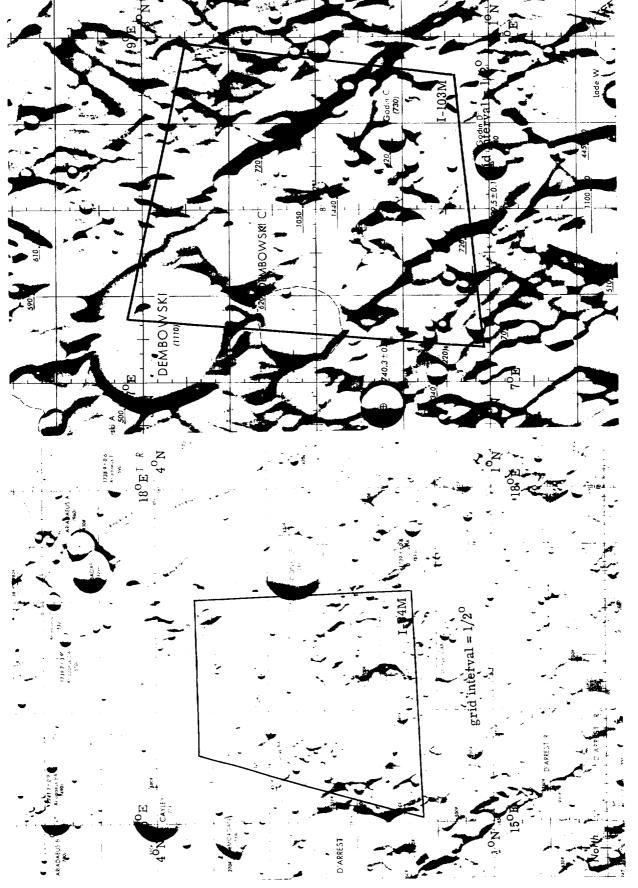
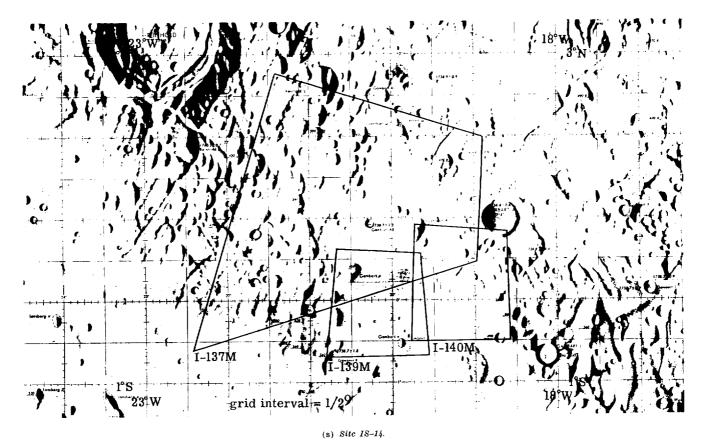


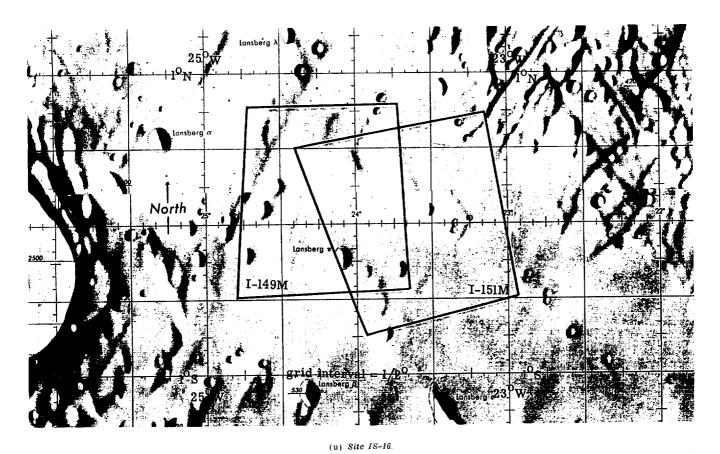
Figure 12.—Photographic Indexes to mission I near-side sites.—Continued.

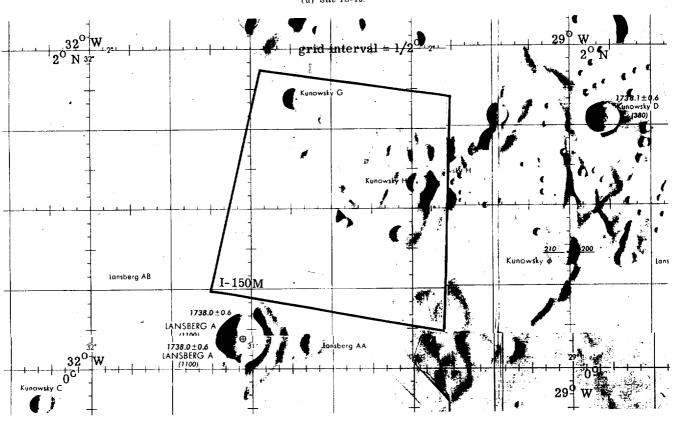
(o) Sitc IS-8.

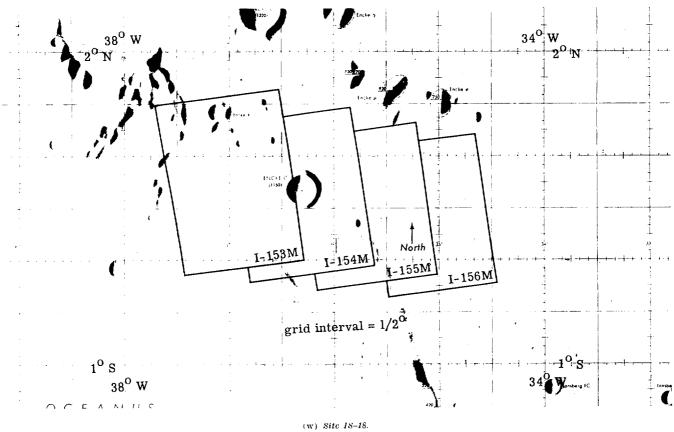




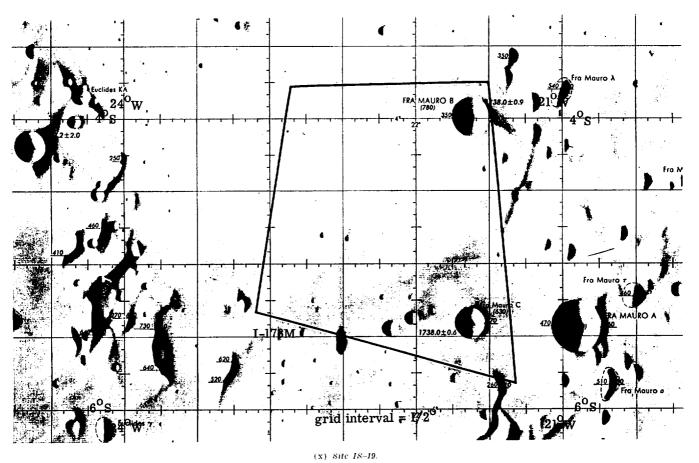
 $\mbox{(t) Sitc IS-15.} \\ \mbox{Figure 12.--Photographic Indexes to mission I near-side sites.--Continued.}$ 



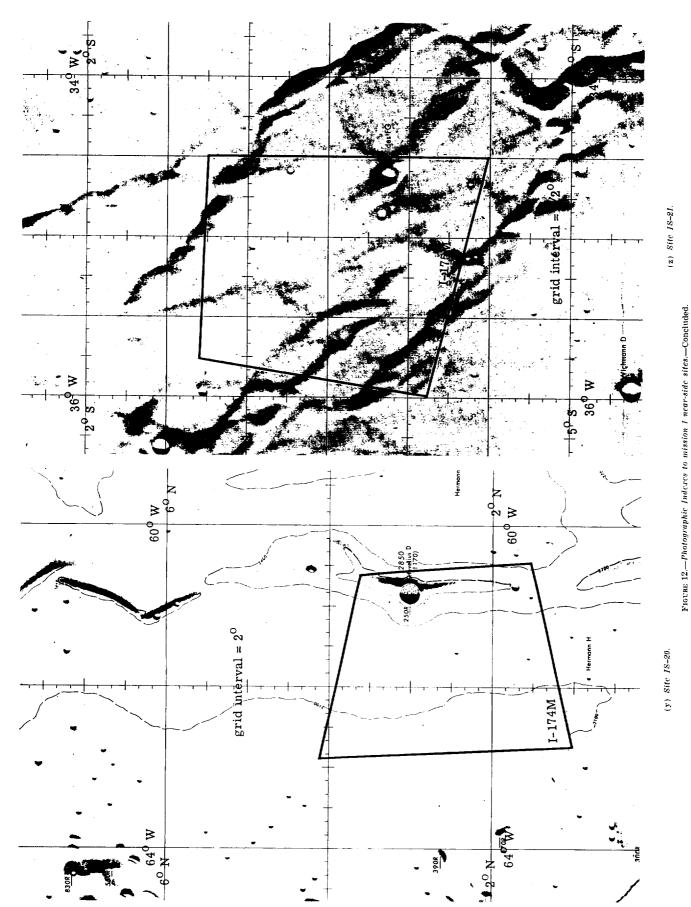


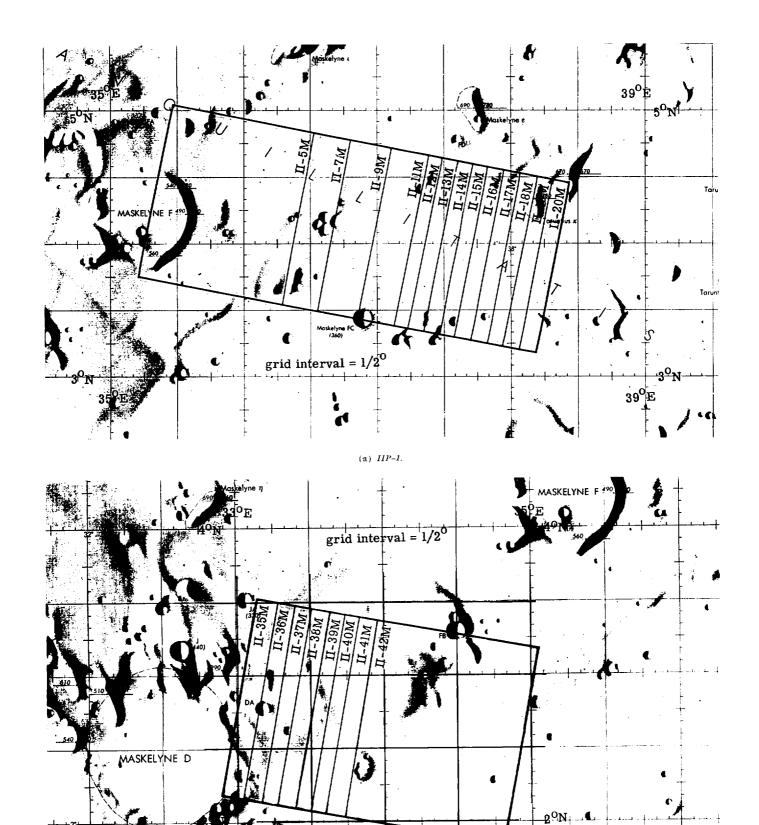




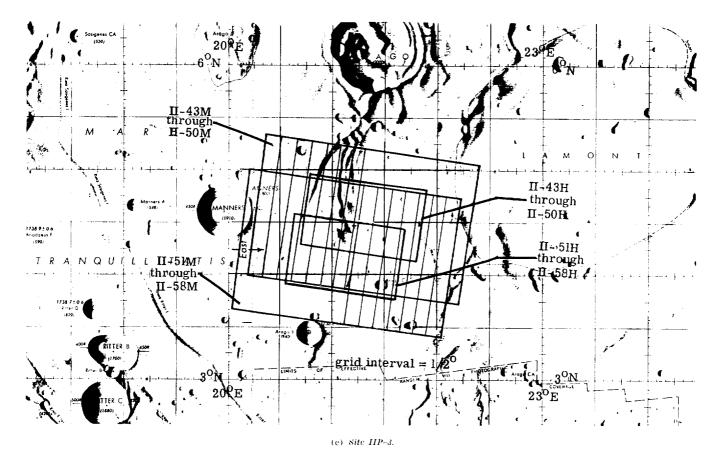


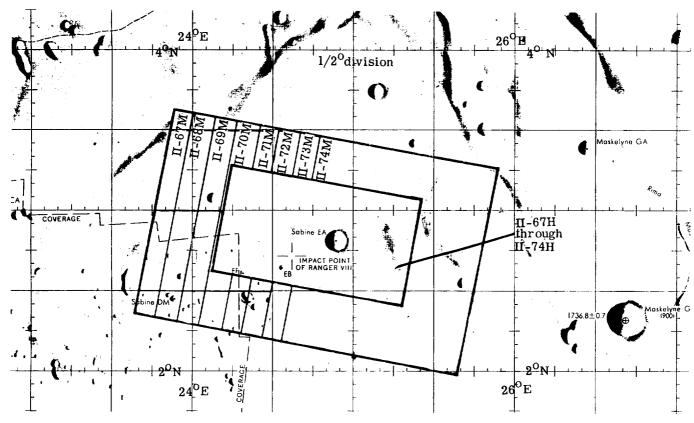
 ${\tt Figure}\ 12. - {\tt Photographic}\ Indexes\ to\ mission\ I\ near-side\ sites. - {\tt Continued}.$ 



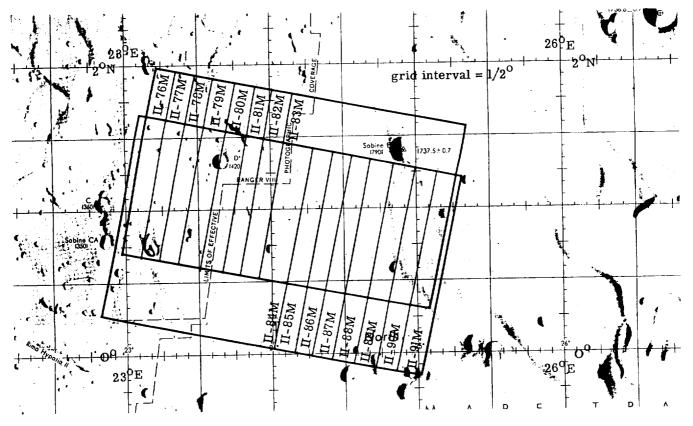


(b) Site IIP-2.
Figure 13.—Photographic Indexes to mission II near-side sites.

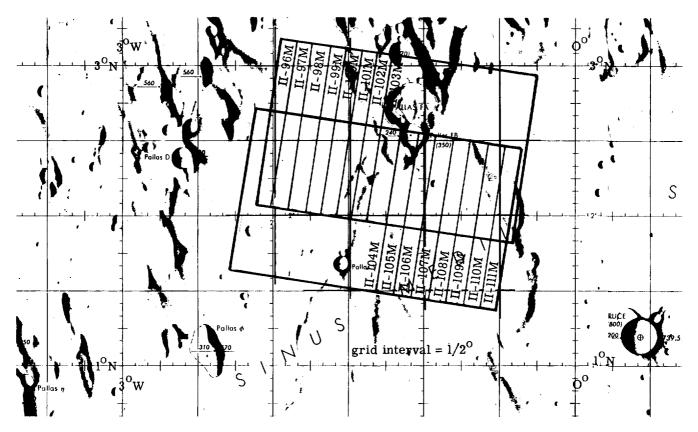




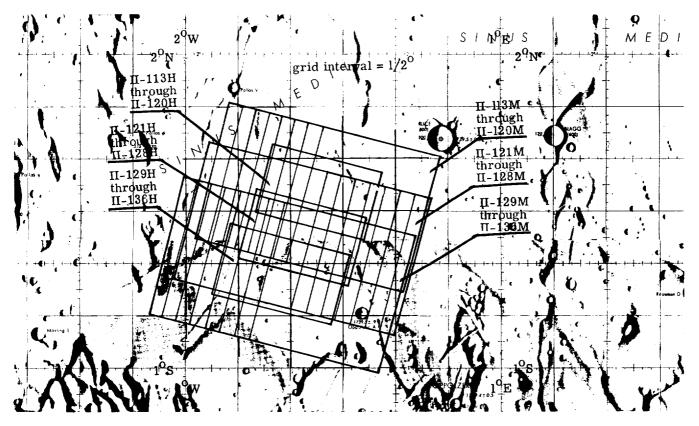
(e) Site IIP-5.



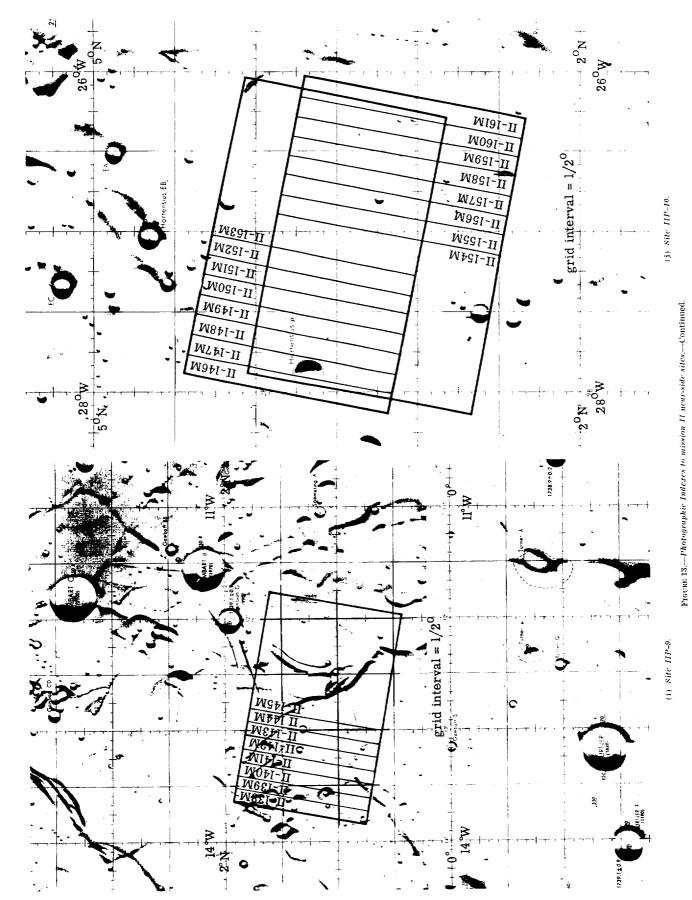
(f) Site IIP=6.

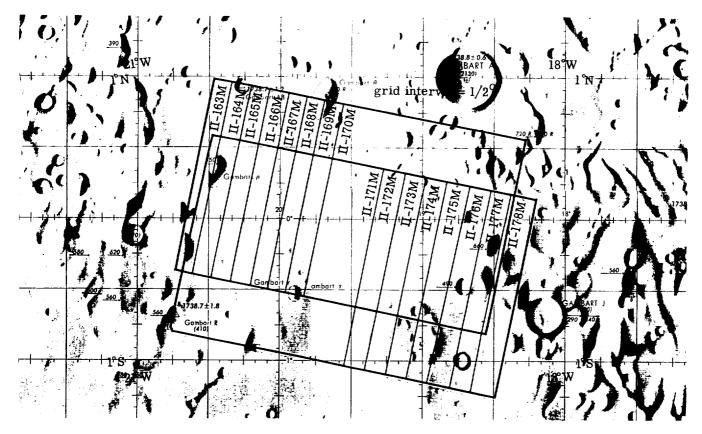


(g) Site IIP-7.

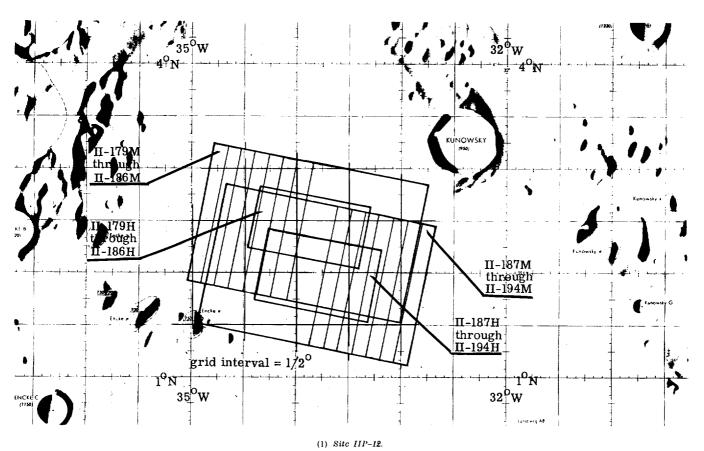


(h) Site HP-8.





(k) Site IIP-11.



 $\textbf{Figure 13.--} Photographic\ Indexes\ to\ mission\ II\ near-side\ sites.-- Continued.$ 



66



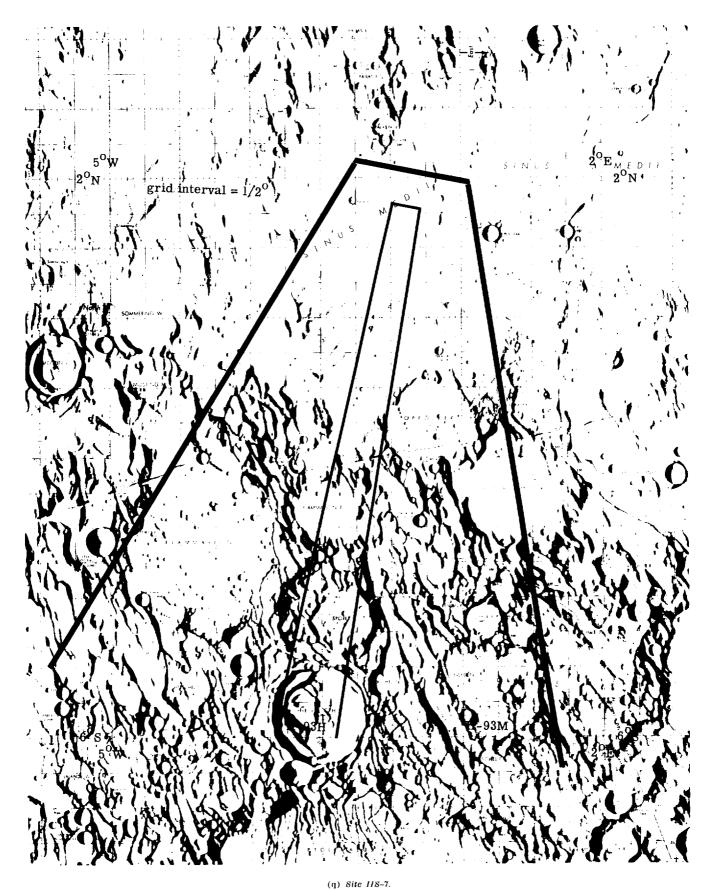
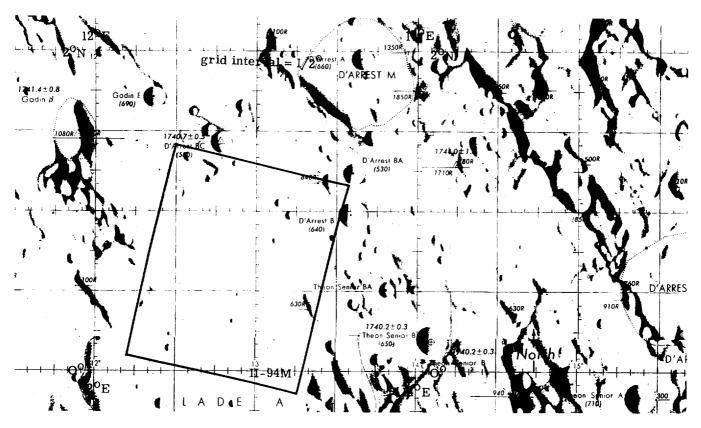
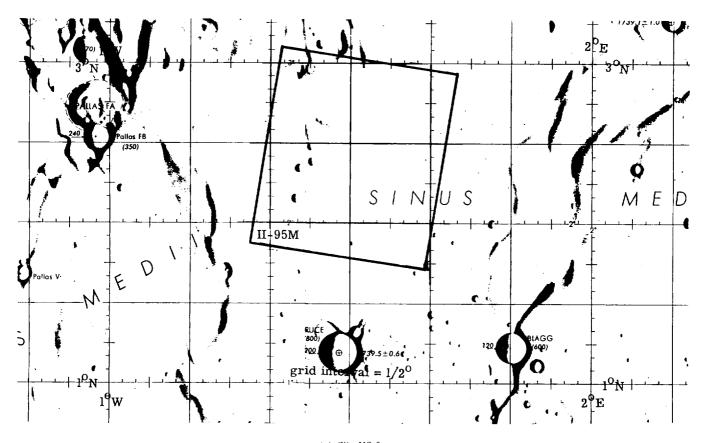


Figure 13.—Photographic Indexes to mission II near-side sites.—Continued.







(s) Site IIS-9.
Figure 13.—Photographic Indexes to mission II near-side sites.—Continued.

Froure 13.—Photographic Indexes to mission II near-side sites.—Continued.

(t) Site 118-10.2.

(u) Sitc IIS-II.

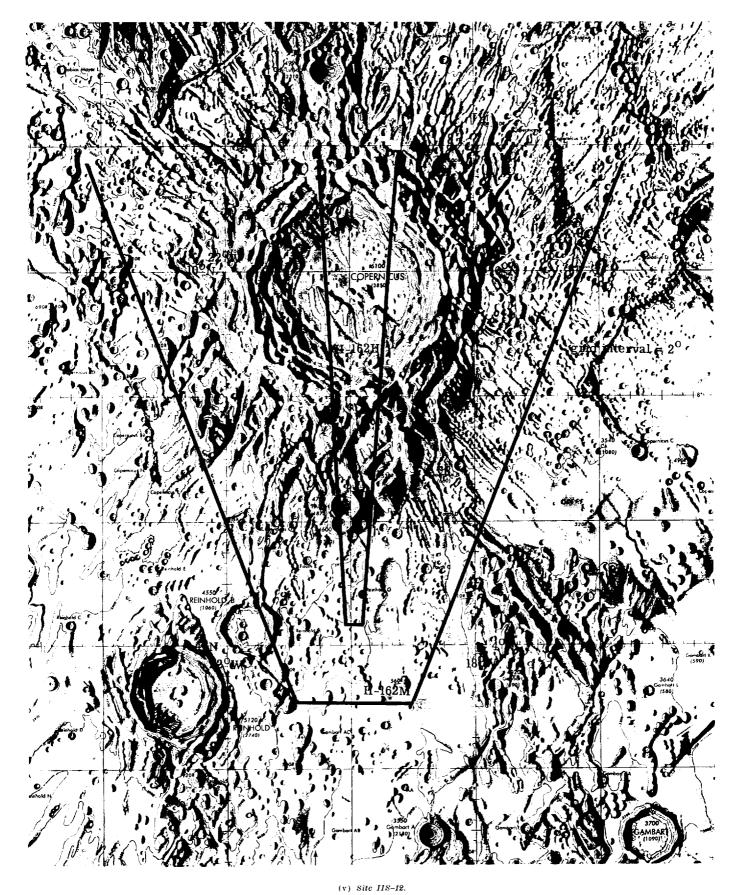
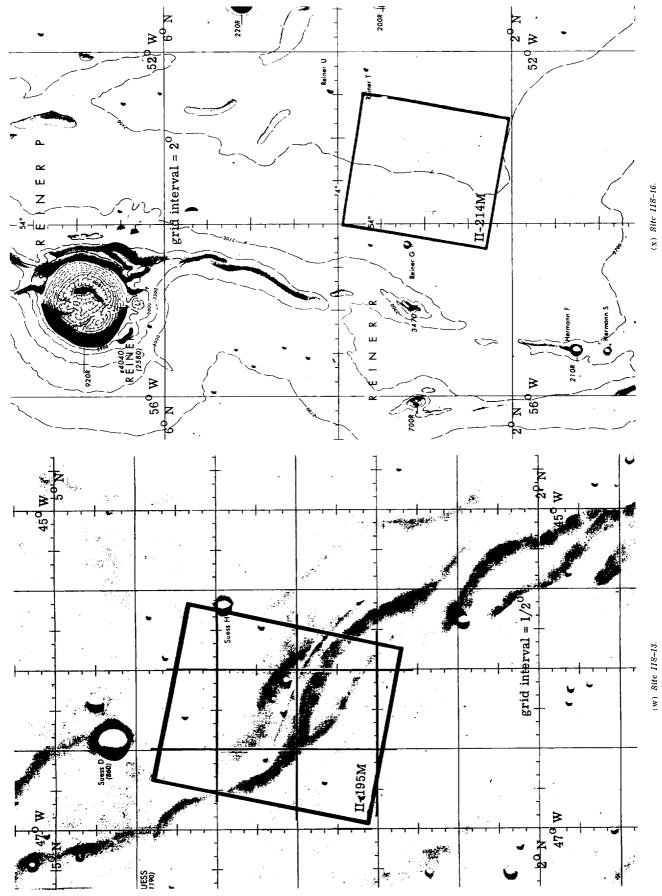


Figure 13.—Photographic Indexes to mission II near-side sites.—Continued.



Fraure 13.—Photographic Indexes to mission II near-side sites.—Continued.



(y) Site IIS-15.

 $\textbf{Figure 13.--} Photographic\ Indexes\ to\ mission\ II\ near-side\ sites.-- Continued.$ 

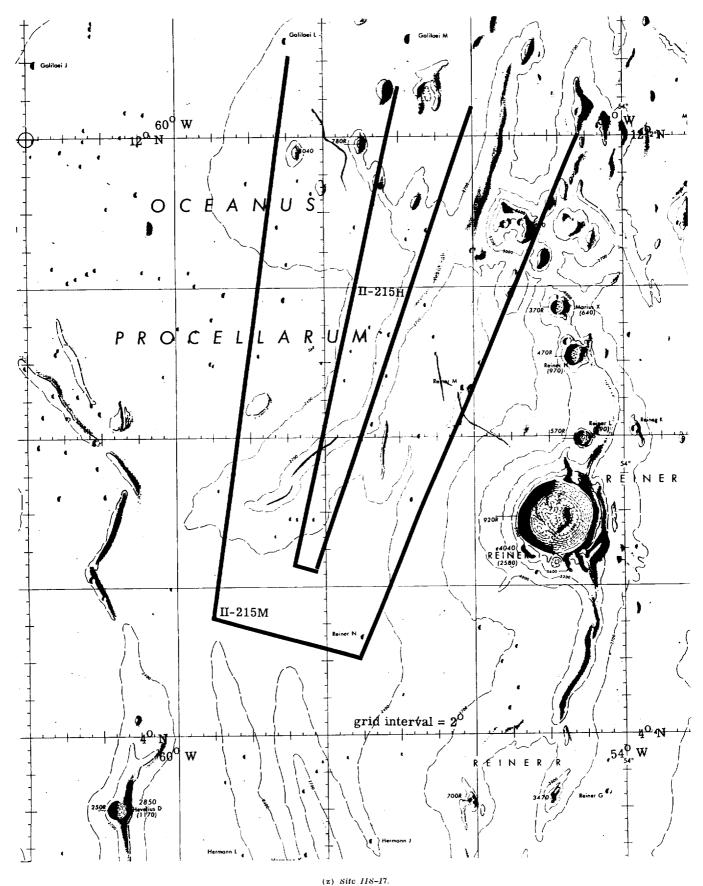
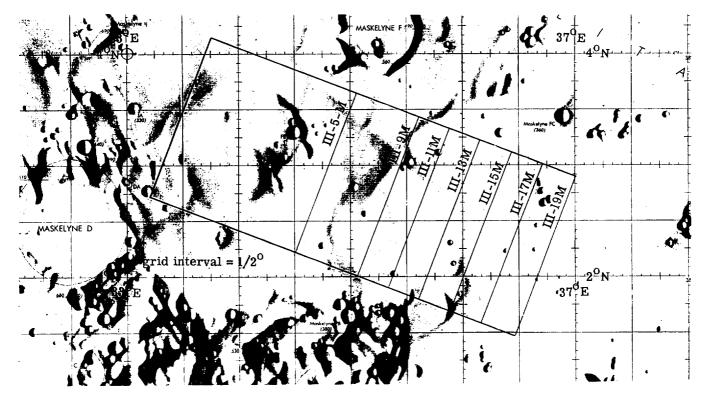
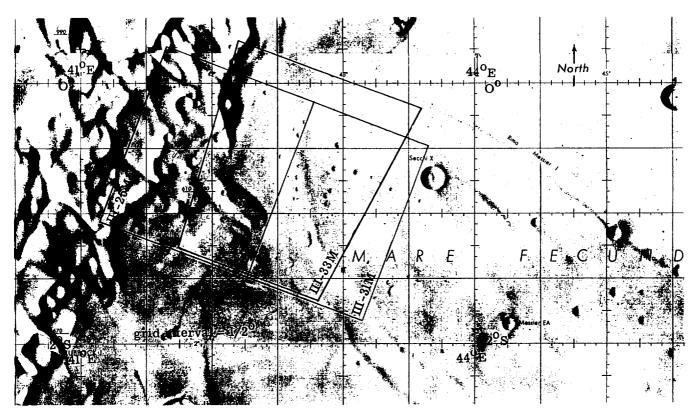


FIGURE 13.—Photographic Indexes to mission II near-side sites.—Concluded.







(b) Site IIIP-2.

FIGURE 14.—Photographic Indexes to mission III near-side sites.

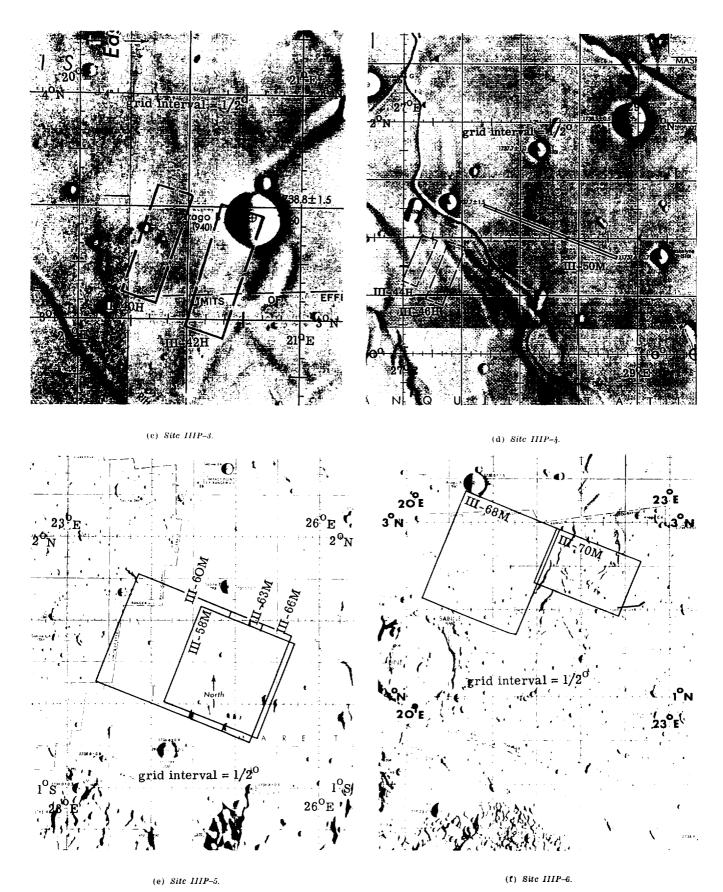
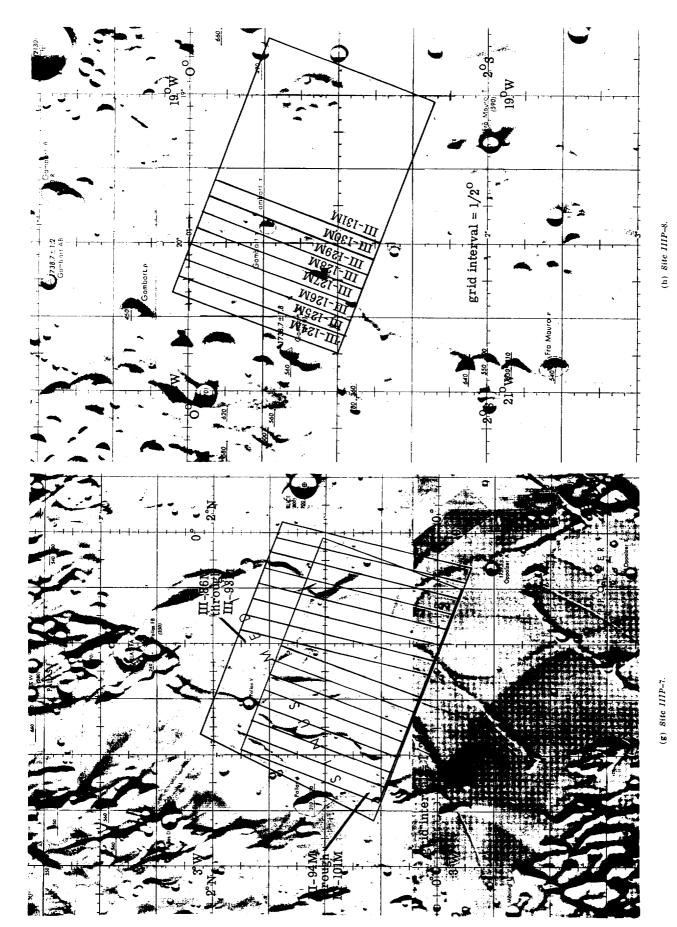
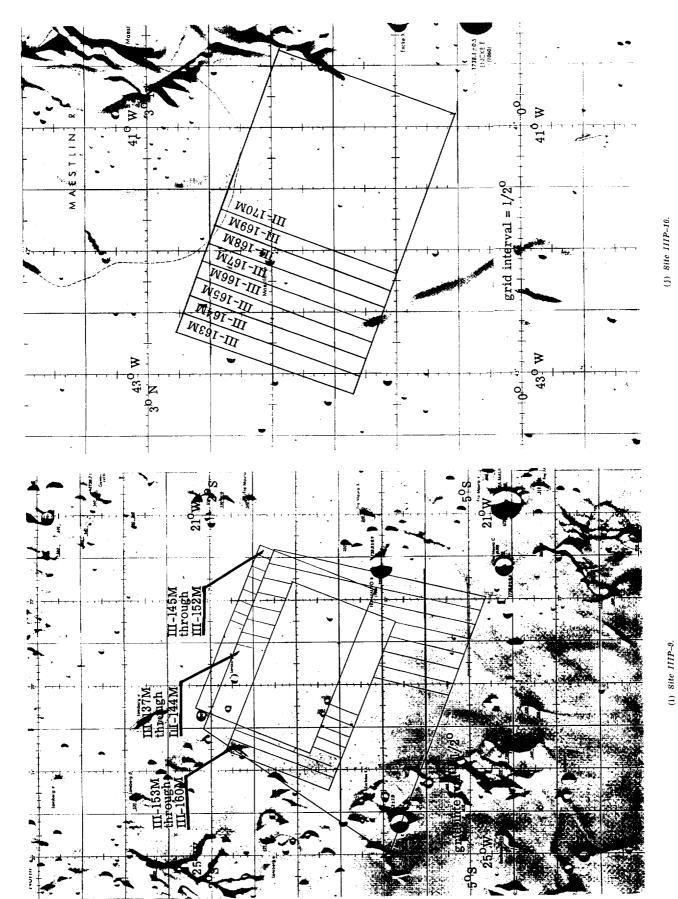


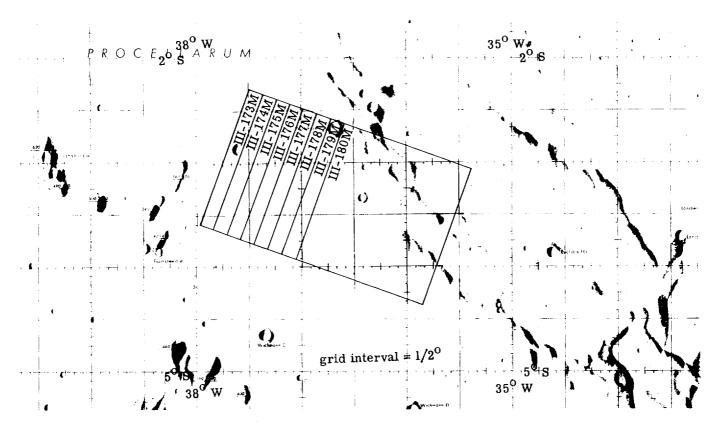
FIGURE 14.—Photographic Indexes to mission III near-side sites.—Continued.



Floure 14.-Photographic Indexes to mission III near-side sites .--Continued.



Froure 14.—Photographic Indexes to mission III near-side sites.—Continued.





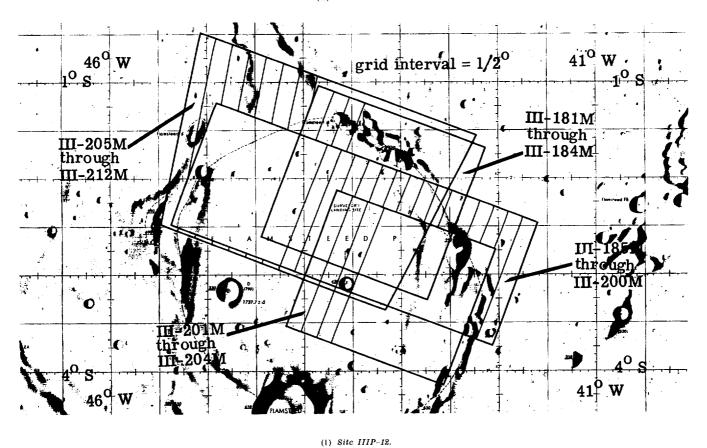


FIGURE 14.—Photographic Indexes to mission III near-side sites.—Continued.

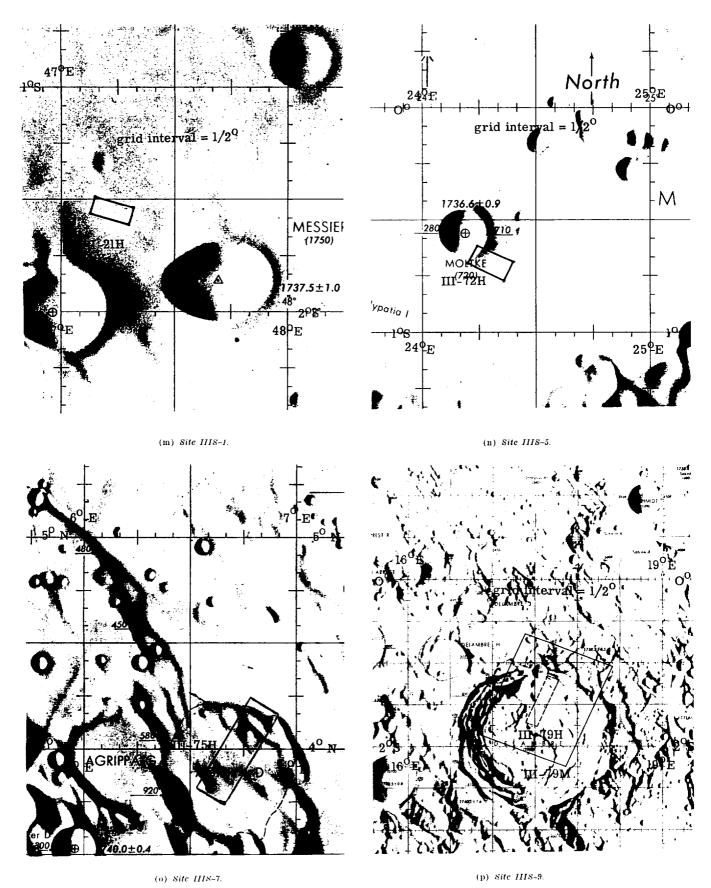


FIGURE 14.—Photographic Indexes to mission III near-side sites.—Continued.

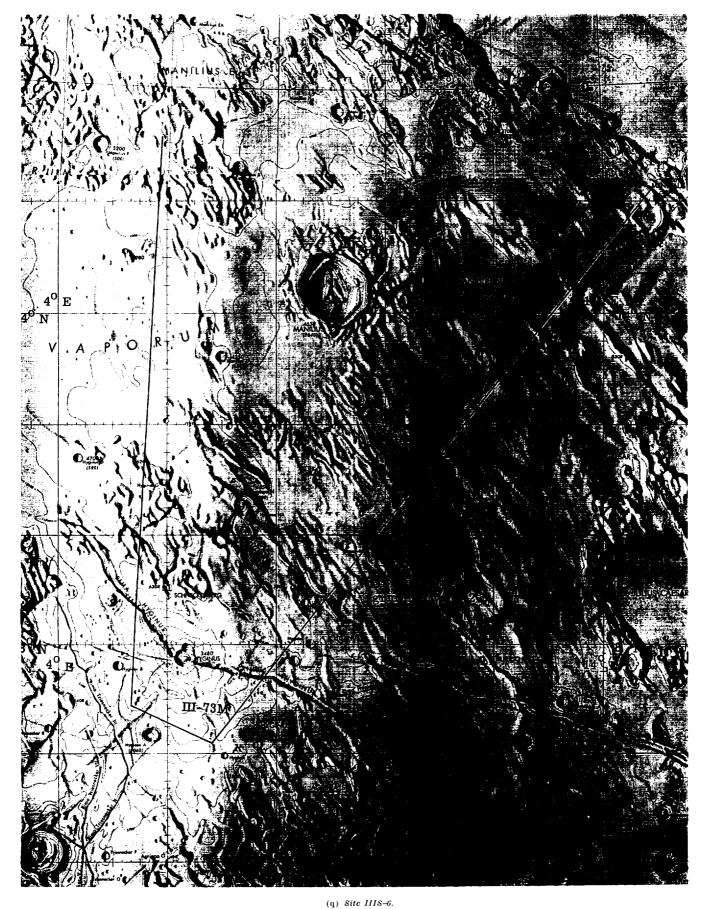


FIGURE 14.—Photographic Indexes to mission III near-side sites.—Continued.

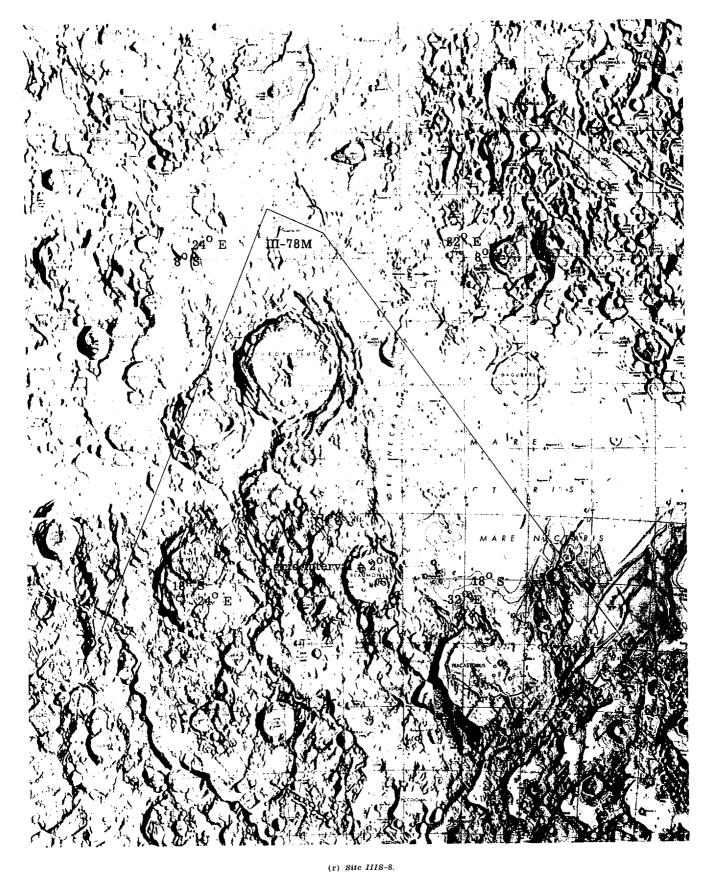
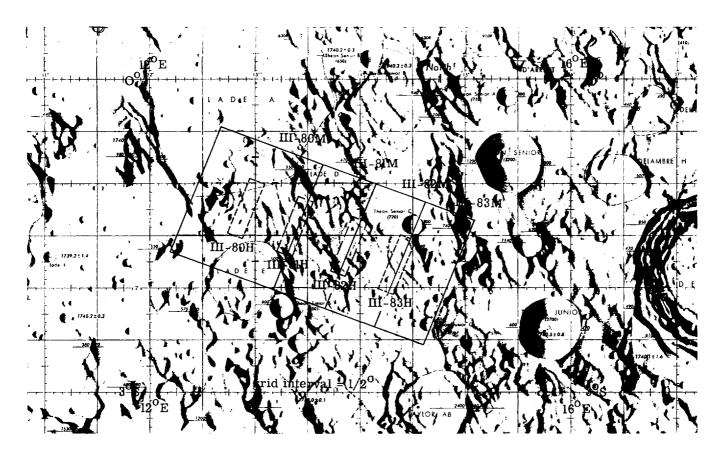
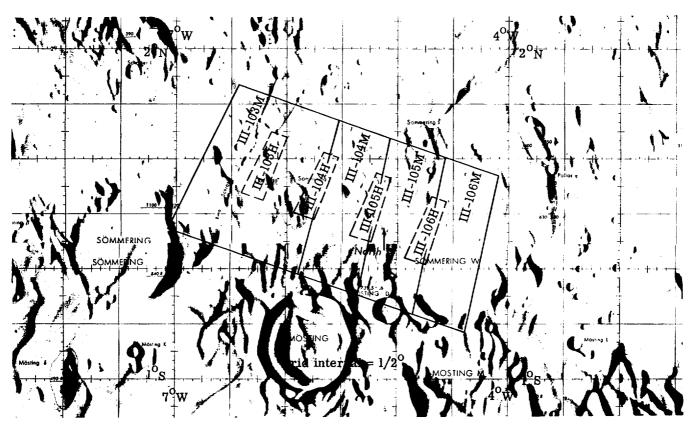


FIGURE 14.—Photographic Indexes to mission III near-side sites.—Continued.



(s) Site IIIS-10.



(t) Site IIIS-15.

FIGURE 14.—Photographic Indexes to mission III near-side sites.—Continued.

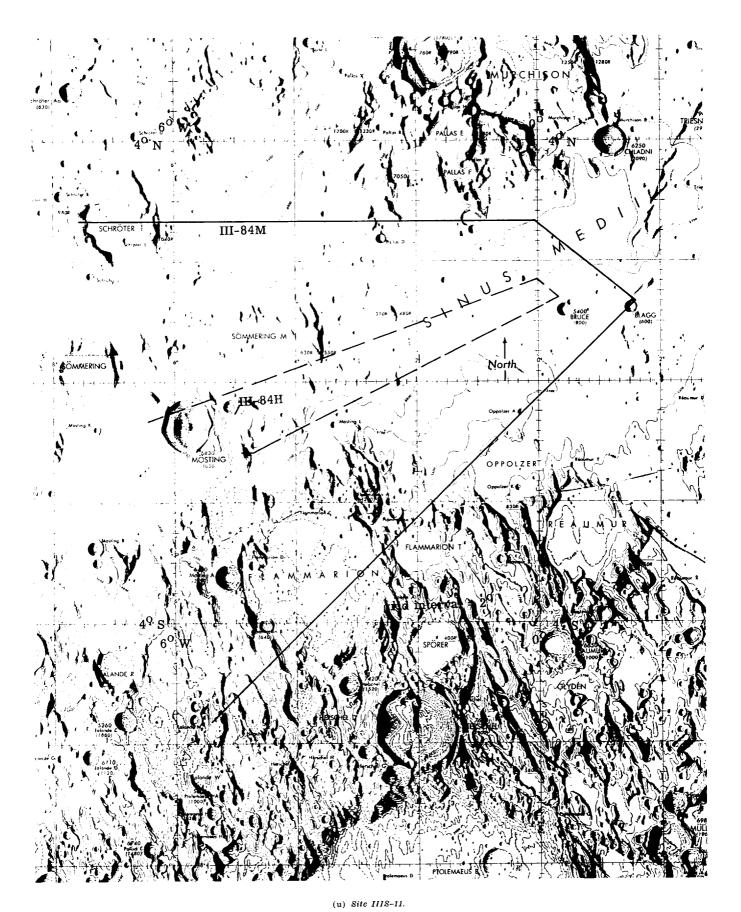
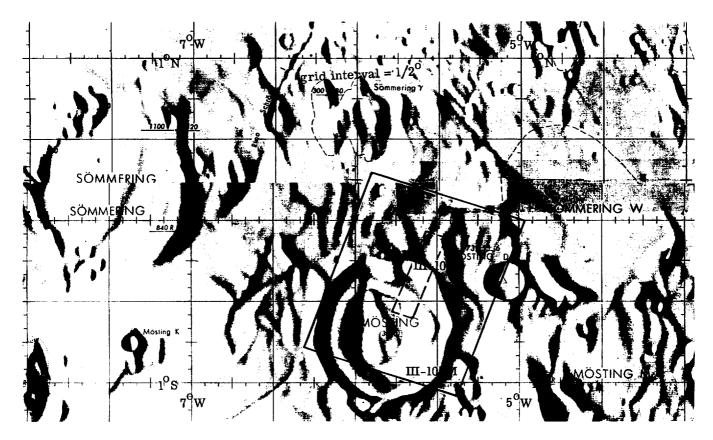


FIGURE 14.—Photographic Indexes to mission III near-side sites.—Continued.

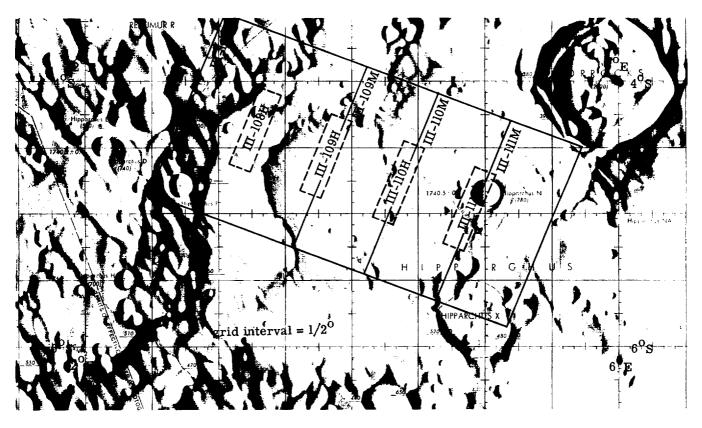


FIGURE 14.—Photographic Indexes to mission III near-side sites.—Continued.

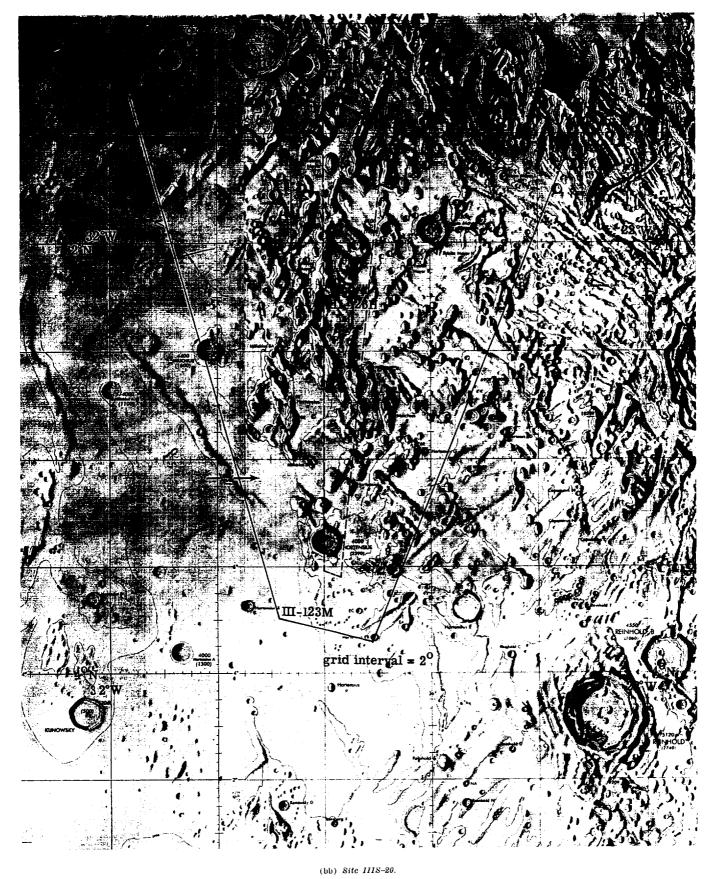
(w) Site IIIS-14. Froune 14.—Photographic Indexes to mission III near-side sites.—Continued.



(x) Site III8-16.



(3) Site IIIS-17.
FIGURE 14.—Photographic Indexes to mission III near-side sites.—Continued.



 ${\bf Figure~14.--} Photographic~Indexes~to~mission~III~near-side~sites.-- Continued.$ 

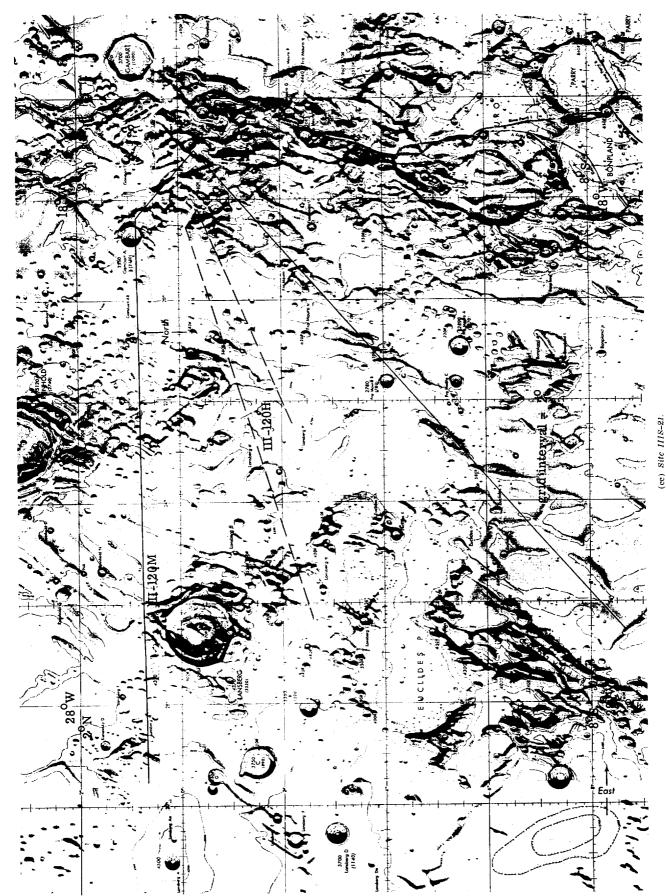
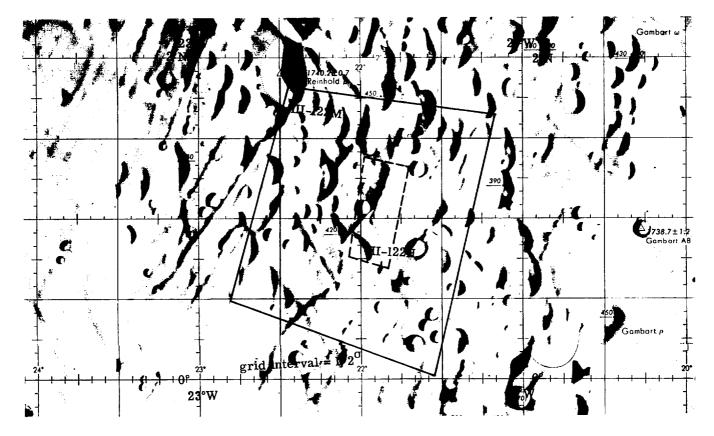


Figure 14.—Photographic Indexes to mission III near-side sites.—Continued.



(dd) Site IIIS-22.

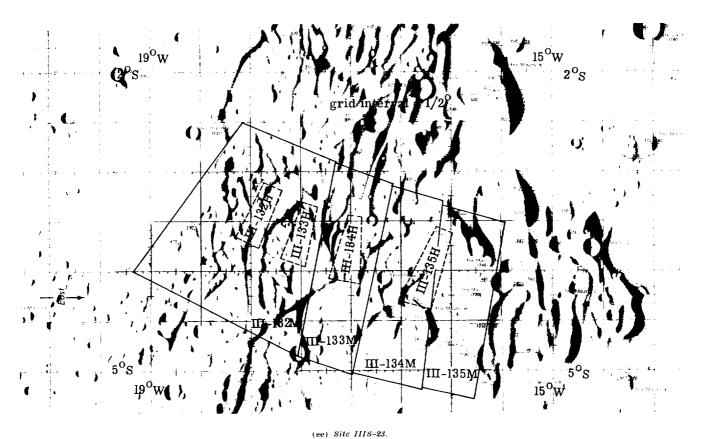
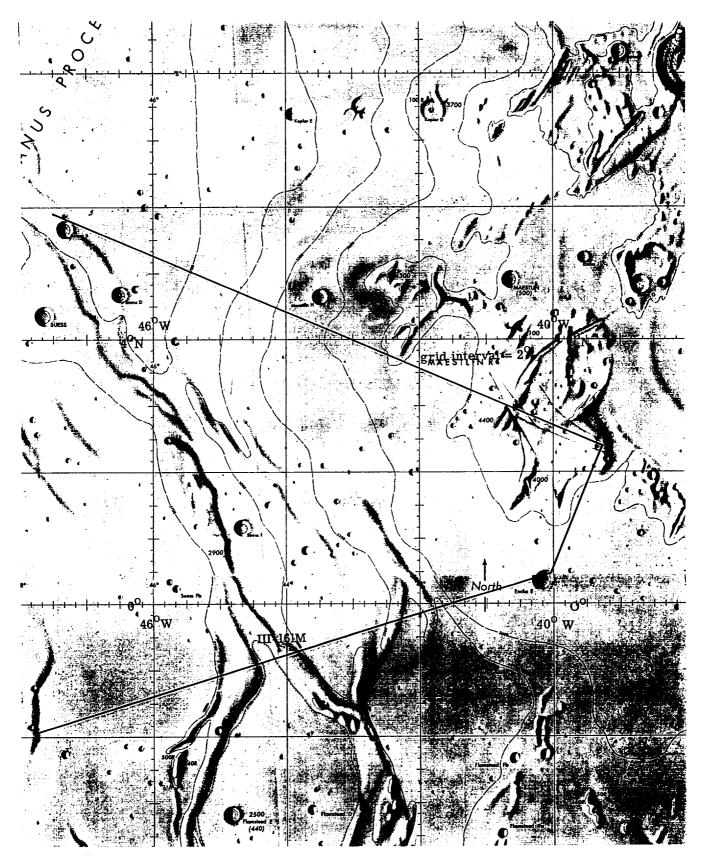


FIGURE 14.—Photographic Indexes to mission III near-side sites.—Continued.



 ${\bf Figure\ 14.--} Photographic\ Indexes\ to\ mission\ III\ near-side\ sites.-- Continued.$ 



(gg) Site IIIS-25.

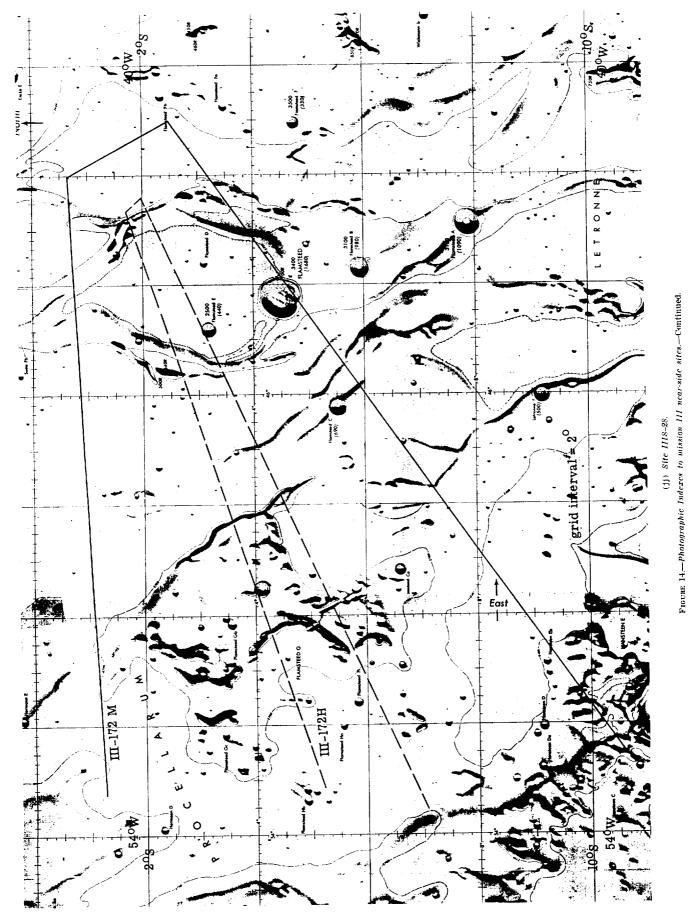
Figure 14.—Photographic Indexes to mission III near-side sites.—Continued.

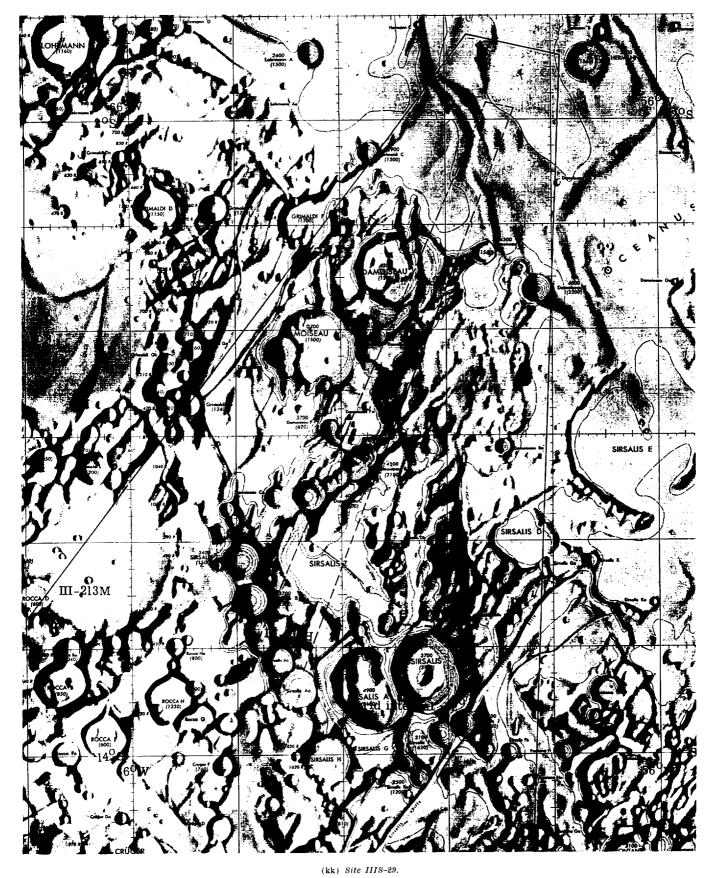
(hh) Site IIIS-26. Froure 14.—Photographic Indexes to mission III near-side sites.—Continued.



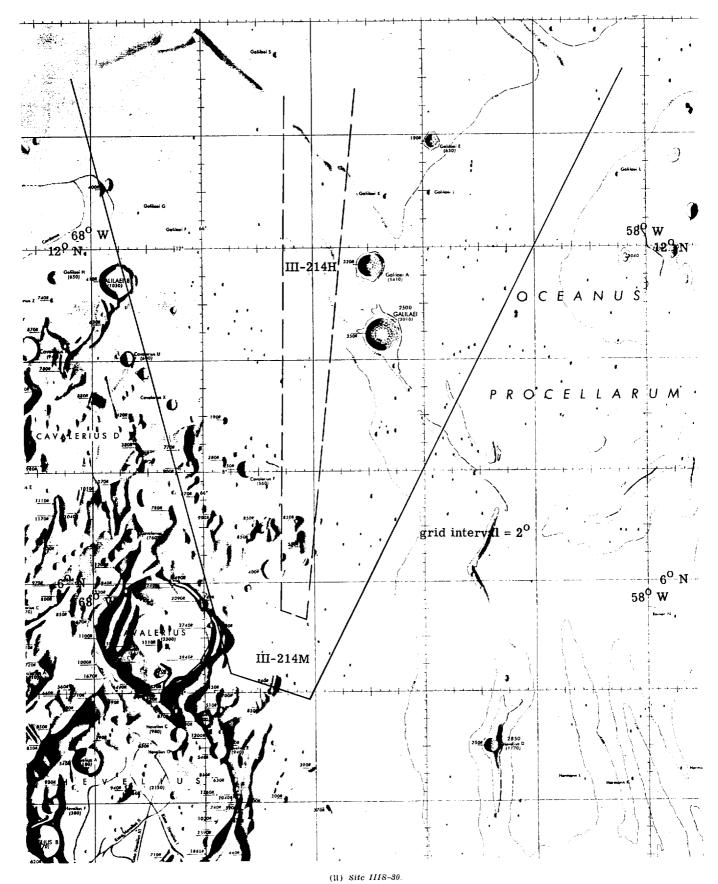
(ii) Site IIIS-27.

FIGURE 14.—Photographic Indexes to mission III near-side sites.—Continued.





 ${\bf Figure\ 14.--} Photographic\ Indexes\ to\ mission\ III\ near-side\ sites.-- Continued.$ 



 ${\bf Figure}\ \ {\bf 14.--Photographic}\ \ Indexes\ \ to\ \ mission\ \ III\ \ near-side\ \ sites.-- Continued.$ 

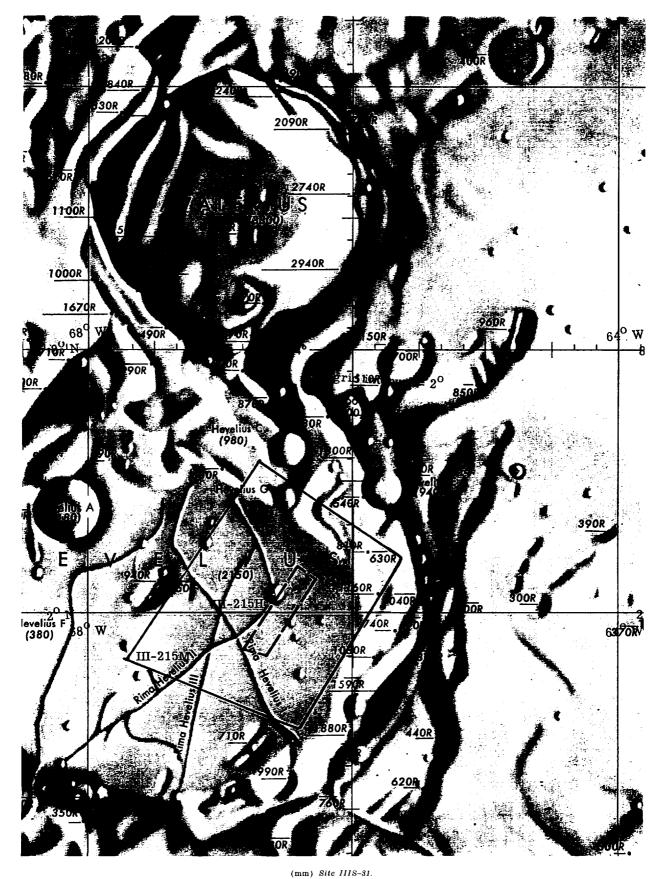
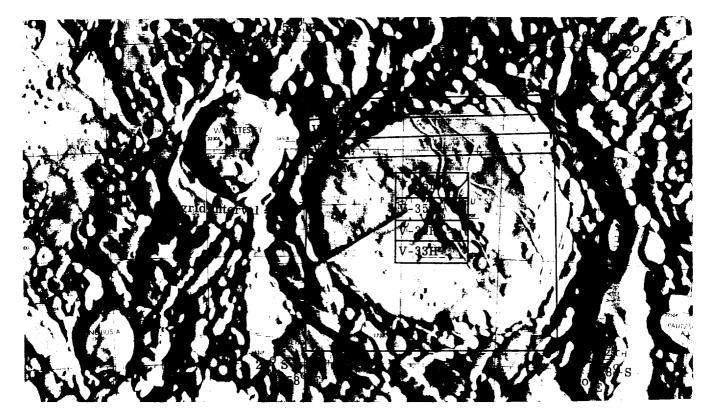
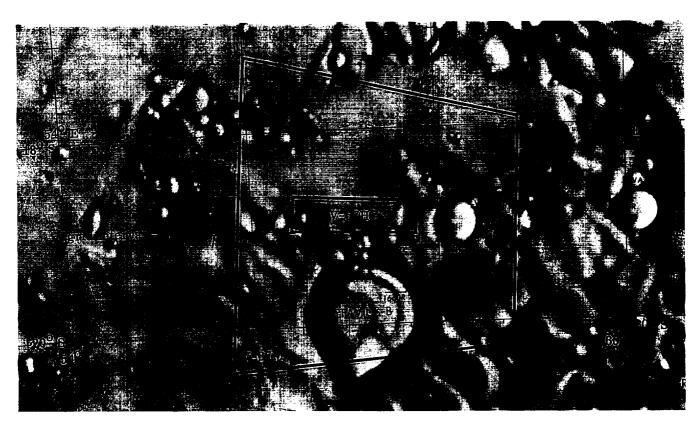


FIGURE 14.—Photographic Indexes to mission III near-side sites.—Concluded.



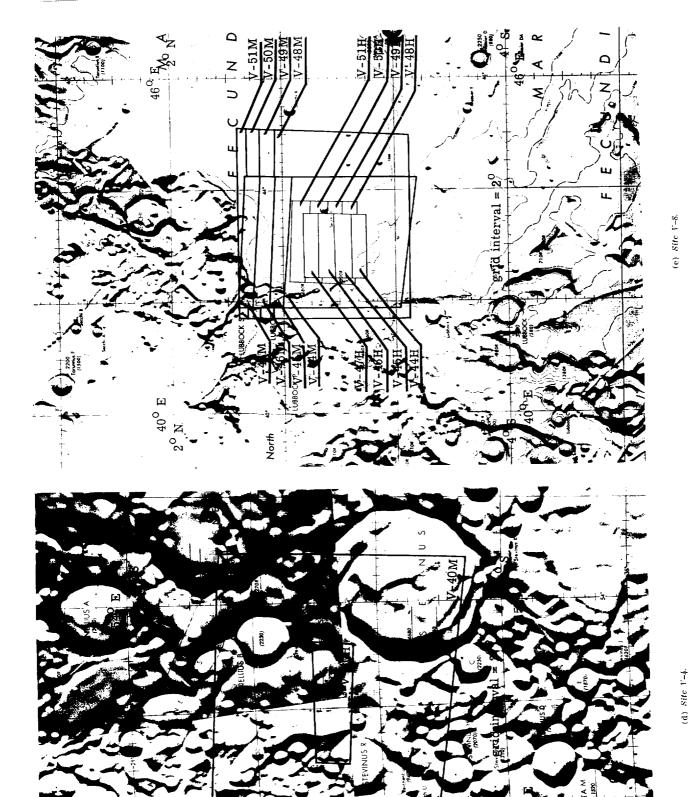
(a) Site V-1.



(b) Site V-2.1.

Figure 15.—Photographic Indexes to mission V near-side sites.

Figure 15.—Photographic Indexes to mission V near-side sites.—Continued.



102



Figure 15.—Photographic Indexes to mission I near-side sifes.—Continued.

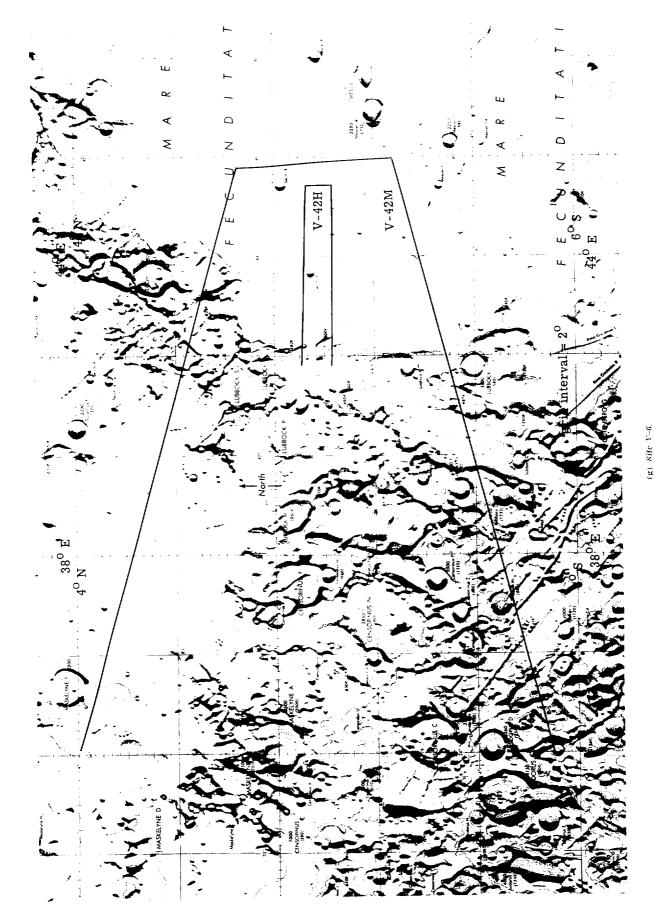


Figure 15.—Photographic Indexes to mission V near-side sites.—Continued.

Figure 15.—Photographic Indexes to mission T near-side sites.—Continued.



 $\mbox{(i) Site $V$-10.} \label{eq:v-10}$  Figure 15.—Photographic Indexes to mission \$V\$ near-side sites.—Continued.

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107

Fraune 15.—Photographic Indexes to mission V near-side sites.—Continued.

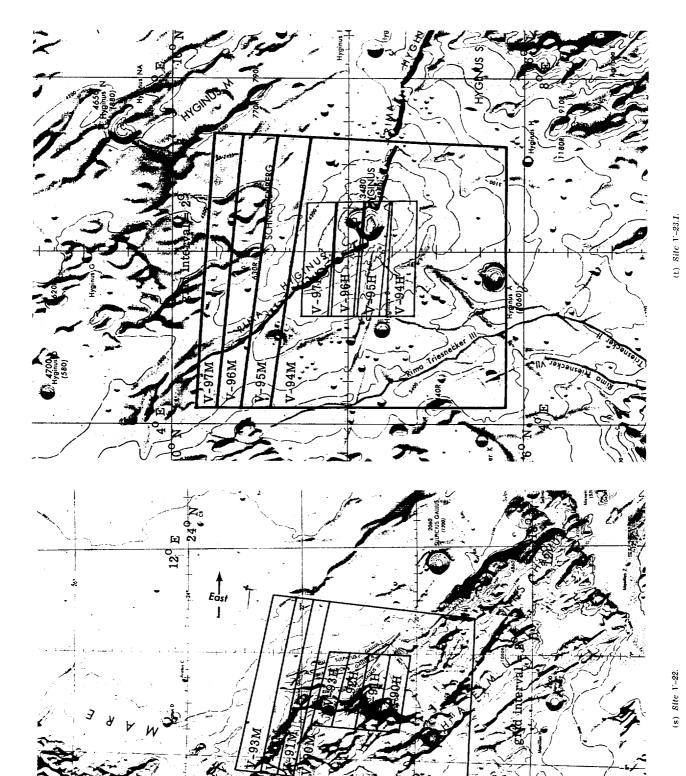
109

Figure 15.-Photographic Indexes to mission V near-side sites .- Continued.

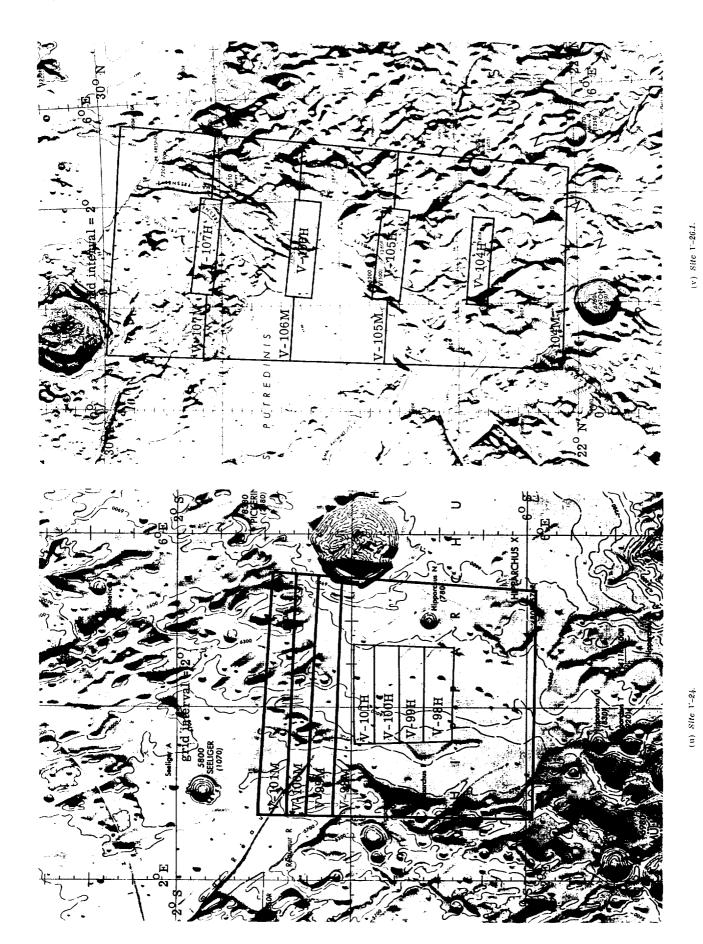
Figure 15.—Photographic Indexes to mission V near-side sites.—Continued.

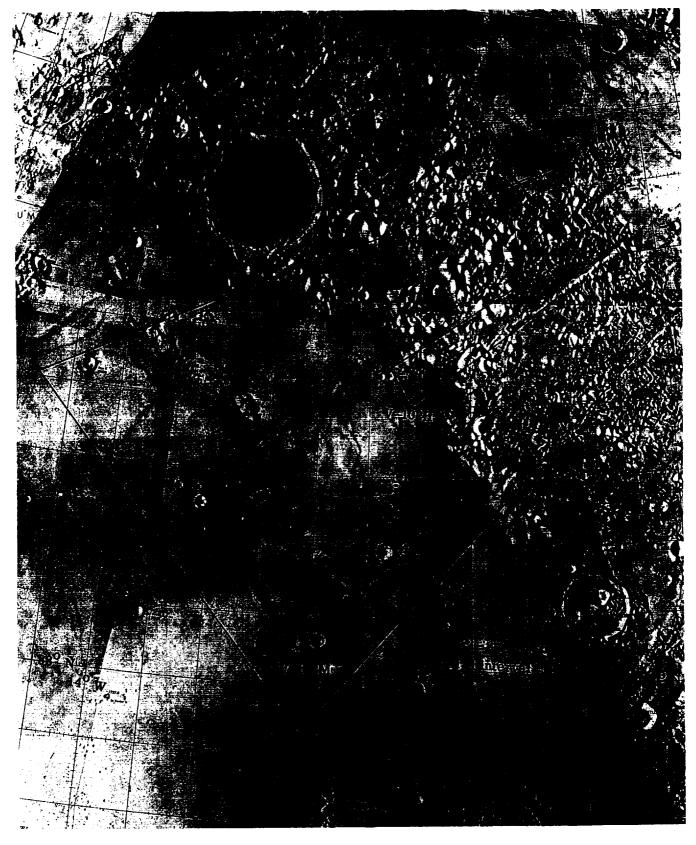
(a) Site V-16.

gure 15.—Photographic Indexes to mission V near-side sites.—Continue

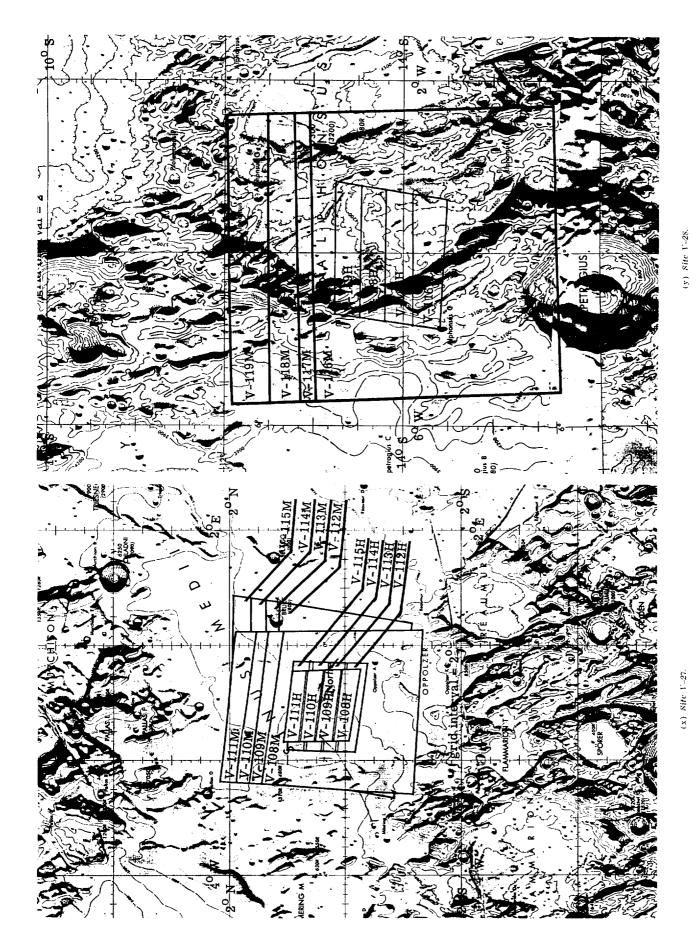


112

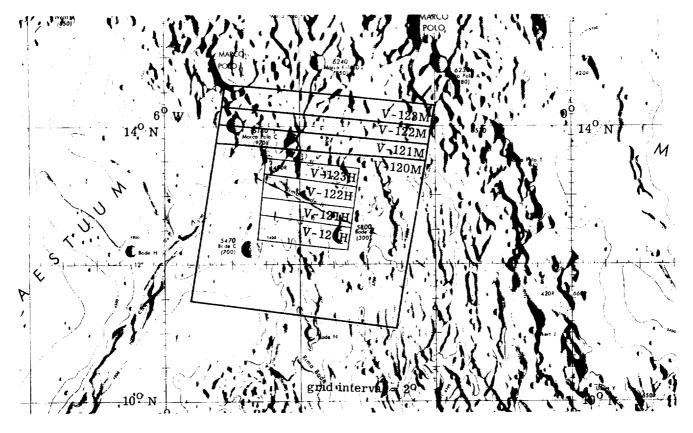




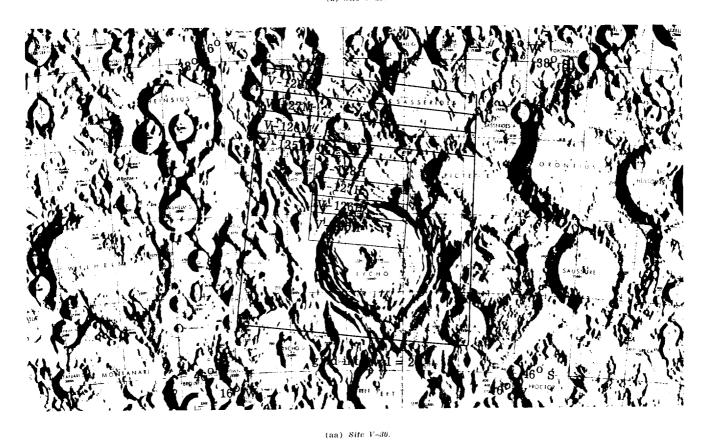
 $\mbox{(w) Site $V$-25.}$  Figure 15.—Photographic Indexes to mission \$V\$ near-side sites.—Continued.



115



(z) Sitc V-29.



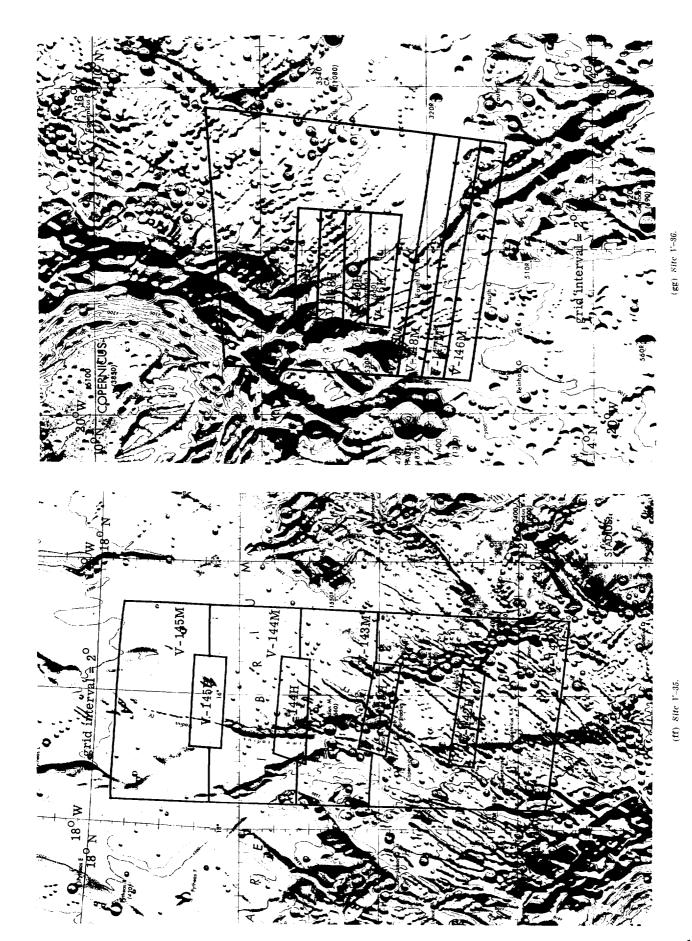
 $\textbf{Figure 15...-} Photographic \ Indexes \ to \ mission \ V \ near-side \ sites. — Continued.$ 

(bb) Site 1-31.

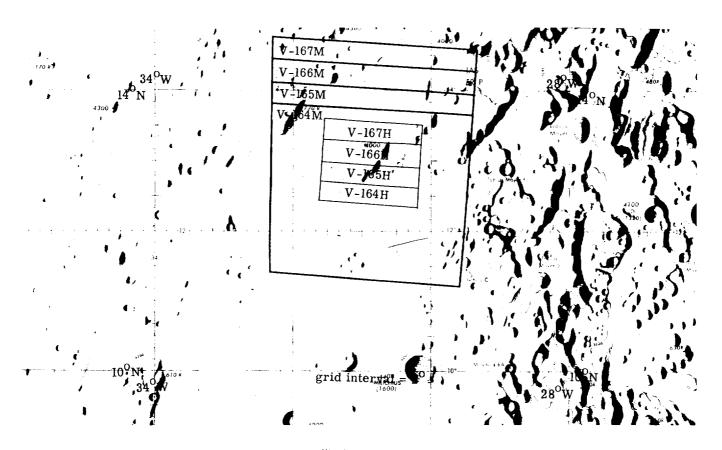
(dd) Site V-33. Flouve 15.—Photographic Indexes to mission V near-side sites.—Continued.

(ee) Site V-34.

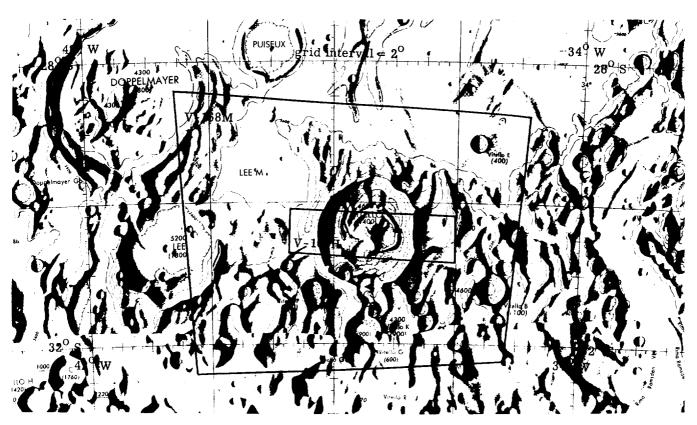
118



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(**jj**) Site Γ-40.



 $\label{eq:continued} (kk) \ \ \textit{Site} \ \ \textit{V-41}.$  Figure 15.—Protographic Indexes to mission V near-side sites.—Continued.

FIGURE 15.—Photographic Indexes to mission V near-side sites.—Continued.

122

Figure 15.—Photographic Indexes to mission V near-side sites.—Continued.

Froure 15.—Photographic Indexes to mission I' near-side sites.—Continued.

(pp) Site V-48.

